# Today in Physics 122: Coulomb's-law examples and useful consequences

- □ Charged disk (from last time)
- Charged line segment; E everywhere
- Approximations and limiting cases
- □ Ion drive for space travel
- □ Ion drives *vs* rockets



NASA's *Dawn* spacecraft, which recently paid visits to asteroids 1 Ceres and 4 Vesta. (Artist's conception: William K. Hartmann and <u>NASA</u>)

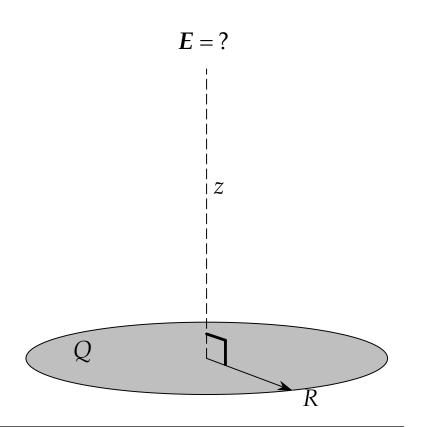
# Example 2: charged disk

*A circular* **disk** *with radius R carries a charge Q which is uniformly distributed on the disk. Calculate the electric field a distance z above the center of the disk.* 

This time it's a two-dimensional charge density, though it's still the same density throughout the charge distribution:

$$\sigma = \frac{Q}{\pi R^2}$$

 The coordinate system from the previous example should be fine.

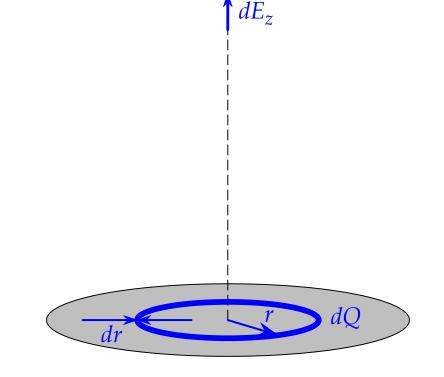


❑ And now for a useful trick: nothing obliges us to use an infinitesimal area element that's infinitesimal in all directions. Let's take as our area element an infinitesimally-thin annulus of radius *r* and width *dr*:

 $dA = 2\pi r dr$  $dQ = \sigma dA = 2\pi \sigma r dr$ 

□ We can regard this as a charged ring, for which we just calculated the field:

$$d\boldsymbol{E} = \frac{1}{4\pi\varepsilon_0} \frac{z\hat{z}}{\left(z^2 + r^2\right)^{3/2}} dQ$$



□ So we add the contributions of all the rings with radius varying between r = 0 and r = R:

$$E(z) = \frac{1}{4\pi\varepsilon_0} \int \frac{z\hat{z}}{(z^2 + r^2)^{3/2}} dQ = \frac{\sigma z\hat{z}}{2\varepsilon_0} \int_0^R \frac{rdr}{(z^2 + r^2)^{3/2}}$$

□ You can look up this integral, as textbooks timidly recommend:

$$\int_{0}^{R} \frac{r dr}{\left(z^{2} + r^{2}\right)^{3/2}} = \left[-\frac{1}{\sqrt{z^{2} + r^{2}}}\right]_{0}^{R} = \frac{1}{z} - \frac{1}{\sqrt{z^{2} + R^{2}}}$$

(This is how you'd finish this problem on an exam, for full credit.)

□ It isn't hard to do the integral oneself, as you learned in MTH 162, with a substitution of variables. One substitution that works is:

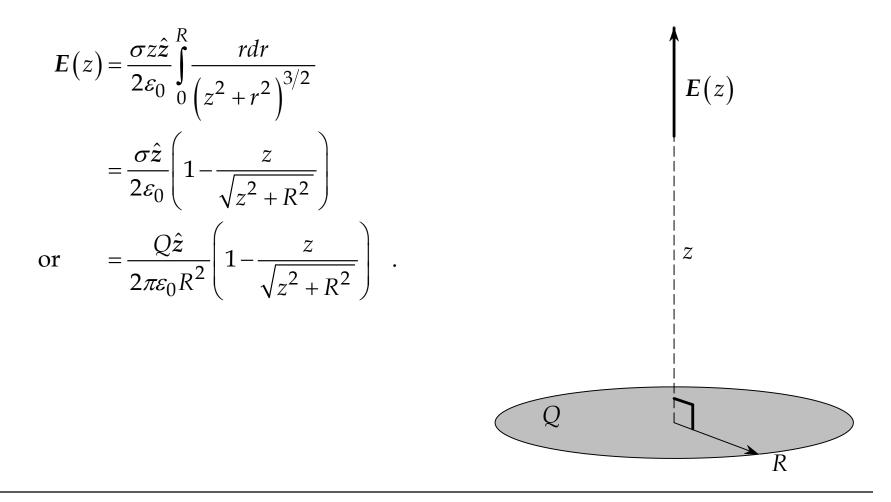
$$u = r^{2} + z^{2} \Rightarrow du = 2rdr$$

$$r = 0 \rightarrow R \Rightarrow u = z^{2} \rightarrow z^{2} + R^{2}$$

$$\int_{0}^{R} \frac{rdr}{\left(z^{2} + r^{2}\right)^{3/2}} = \frac{1}{2} \int_{z^{2}}^{z^{2} + R^{2}} u^{-3/2} du = \frac{1}{2} \left[ -2u^{-1/2} \right]_{z^{2}}^{z^{2} + R^{2}}$$

$$= \frac{1}{z} - \frac{1}{\sqrt{z^{2} + R^{2}}}$$

Either way, we're ready to write down the final answer:



# **Example: field of a line charge**

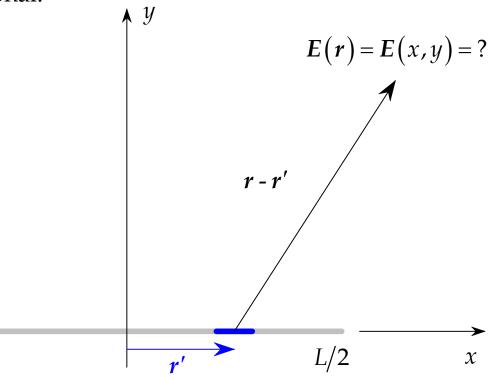
*A straight wire of length* L *carries a total charge* Q. *Calculate the electric field everywhere.* Not quite like the workshop problem.

□ Straight wires are one-dimensional:

 $\lambda = \frac{Q}{L} \quad .$ 

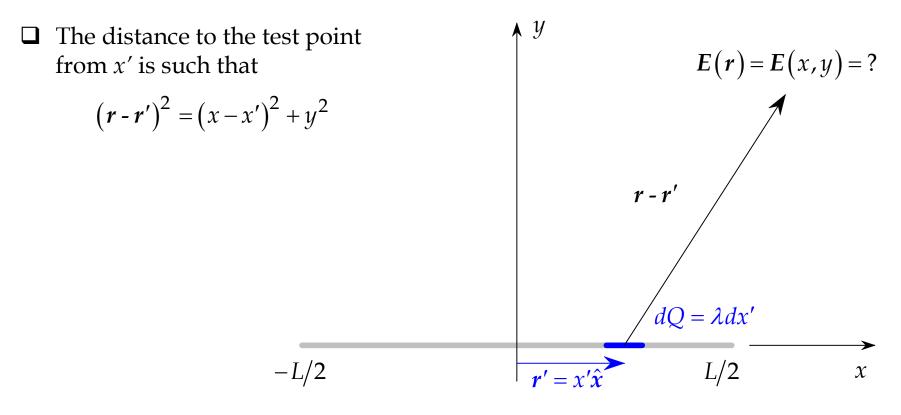
Better run one axis along the wire. It doesn't really matter where along the wire the origin goes, but the midpoint will turn out to be good.

-L/2

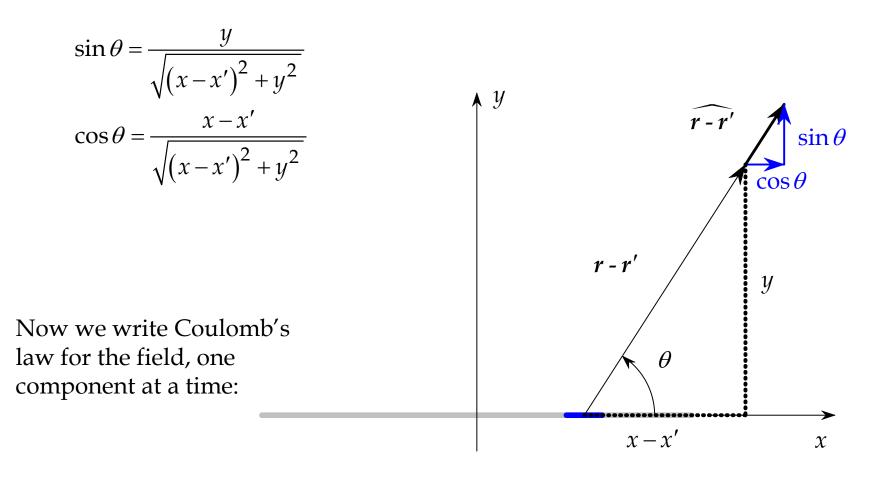


□ The obvious choice for an infinitesimal element of length is dx', making  $dQ = \lambda dx'$ . Then the integral can run from

x' = -L/2 to L/2.



**\Box** From the triangle formed by the axes and r - r',



$$E_{x} = \int \frac{1}{4\pi\varepsilon_{0}} \frac{dQ}{(r-r')^{2}} \cos\theta = \frac{1}{4\pi\varepsilon_{0}} \int_{-L/2}^{L/2} \frac{\lambda dx'}{(x-x')^{2} + y^{2}} \frac{x-x'}{\sqrt{(x-x')^{2} + y^{2}}}$$

□ With a minor variable substitution,

$$X = x - x' \implies dX = -dx'$$
  
$$x' = -L/2 \implies L/2 \implies X = x + L/2 \implies x - L/2$$

we see this is actually the integral we did with the charged disk:

$$E_{x} = -\frac{1}{4\pi\varepsilon_{0}} \lambda \int_{x+L/2}^{x-L/2} \frac{XdX}{\left(X^{2} + y^{2}\right)^{3/2}} = \frac{\lambda}{4\pi\varepsilon_{0}} \left[\frac{1}{\sqrt{X^{2} + y^{2}}}\right]_{x+L/2}^{x-L/2}$$

$$E_{x} = \frac{\lambda}{4\pi\varepsilon_{0}} \left[ \frac{1}{\sqrt{(x - L/2)^{2} + y^{2}}} - \frac{1}{\sqrt{(x + L/2)^{2} + y^{2}}} \right]$$

or

□ Now the *y* component:

$$E_{y} = \int \frac{1}{4\pi\varepsilon_{0}} \frac{dQ}{(r-r')^{2}} \sin\theta = \frac{\lambda y}{4\pi\varepsilon_{0}} \int_{-L/2}^{L/2} \frac{dx'}{((x-x')^{2}+y^{2})^{3/2}}$$

In the spirit of the textbook, we can look up the integral in a table:

$$E_y = -\frac{\lambda}{4\pi\varepsilon_0} \frac{1}{y} \left[ \frac{x - x'}{\sqrt{\left(x - x'\right)^2 + y^2}} \right]_{-L/2}^{L/2}$$

It's not that hard to integrate one's self: make a trigonometric substitution:

$$x - x' = y \tan \alpha \implies$$
  

$$-dx' = y \left(1 + \tan^2 \alpha\right) d\alpha = \frac{y}{\cos^2 \alpha} d\alpha$$
  

$$x' = -L/2 \rightarrow L/2 \implies$$
  

$$\alpha = \arctan\left(\frac{x + L/2}{y}\right) \rightarrow \arctan\left(\frac{x - L/2}{y}\right) = \alpha_+ \rightarrow \alpha_-$$

The denominator under the integral becomes

$$\left( \left( x - x' \right)^2 + y^2 \right)^{3/2} = \left( y^2 \tan^2 \alpha + y^2 \right)^{3/2} = \frac{y^3}{\cos^3 \alpha}$$

so 
$$E_{y} = -\frac{1}{4\pi\varepsilon_{0}}\lambda y \int_{\alpha_{+}}^{\alpha_{-}} \frac{\cos^{3}\alpha}{y^{3}} \frac{yd\alpha}{\cos^{2}\alpha} = +\frac{\lambda}{4\pi\varepsilon_{0}} \frac{1}{y} \int_{\alpha_{-}}^{\alpha_{+}} \cos\alpha d\alpha = \frac{\lambda}{4\pi\varepsilon_{0}} \frac{1}{y^{2}} \sin\alpha \Big|_{\alpha_{-}}^{\alpha_{+}}$$
$$= \frac{\lambda}{4\pi\varepsilon_{0}} \frac{1}{y} \left[ \sin\left(\arctan\left[\frac{x+L/2}{y}\right]\right) - \sin\left(\arctan\left[\frac{x-L/2}{y}\right]\right) \right]$$
$$= \frac{\lambda}{4\pi\varepsilon_{0}} \frac{1}{y} \left[ \frac{x+L/2}{\sqrt{(x+L/2)^{2}+y^{2}}} - \frac{x-L/2}{\sqrt{(x-L/2)^{2}+y^{2}}} \right].$$

(Recall that if  $\tan \theta = a/b$ , then  $\sin \theta = a/\sqrt{a^2 + b^2}$  and  $\cos \theta = b/\sqrt{a^2 + b^2}$ .)

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And

And we get 
$$E_{y} = \frac{\lambda}{4\pi\varepsilon_{0}} \frac{1}{y} \left[ \frac{x + L/2}{\sqrt{(x + L/2)^{2} + y^{2}}} - \frac{x - L/2}{\sqrt{(x - L/2)^{2} + y^{2}}} \right]$$

$$\square \text{ Note that if } x = 0, E_{x} = 0, \text{ and thus}$$

$$E = \frac{\lambda}{4\pi\varepsilon_{0}} \frac{1}{y} \frac{L}{\sqrt{(L/2)^{2} + y^{2}}} \hat{y}$$
as found in workshop, and as you might have expected from the symmetry.

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## **Other useful consequences of our recent Examples**

#### **1.** Far away from a line charge

If the length *L* is very small compared to *y*, then the field in the plane through the midpoint looks even simpler:

$$E(0,y) = \frac{1}{4\pi\varepsilon_0} \frac{\chi L}{y\sqrt{(L/2)^2 + y^2}} \hat{y} \cong \frac{1}{4\pi\varepsilon_0} \frac{Q}{y^2} \hat{y}$$
  
neglect, compared to  $y^2$ .

just as it's supposed to from a place far enough away that the line looks like a point (Coulomb's law).

Examining a limit in which you know what the answer is, is a good way of checking your answer.

# Other useful consequences of our recent Examples (continued)

2. **Close to a line charge** (or anywhere, in the presence of an infinitely-long line charge)

If on the other hand *L* is much larger than *y*,

$$E(0,y) = \frac{1}{4\pi\varepsilon_0} \frac{\lambda L}{y\sqrt{(L/2)^2 + y^2}} \hat{y} \cong \frac{1}{2\pi\varepsilon_0} \frac{\lambda}{y} \hat{y}$$
  
neglect, compared to  $(L/2)^2$ 

*E* points perpendicularly away from a long line charge, and its magnitude drops off with increasing distance as 1/y.

## Other useful consequences of our recent Examples (continued)

3. Far from a charged disk.

For those familiar with the binomial approximation,  $(1+x)^n \cong 1+nx$  if  $|x| \ll 1$ :

$$E(z) = \frac{\sigma}{2\varepsilon_0} \hat{z} \left( 1 - \frac{z}{\sqrt{z^2 + R^2}} \right) = \frac{\sigma}{2\varepsilon_0} \hat{z} \left( 1 - \frac{1}{\sqrt{1 + (R/z)^2}} \right)$$
$$\approx \frac{\sigma}{2\varepsilon_0} \hat{z} \left( 1 - \left[ 1 - \frac{1}{2} \left( \frac{R}{z} \right)^2 \right] \right) = \frac{1}{4\pi\varepsilon_0} \frac{\pi R^2 \sigma}{z^2} \hat{z}$$
$$= \frac{1}{4\pi\varepsilon_0} \frac{Q}{z^2} \hat{z} \quad , \text{ again recovering Coulomb's law.}$$

## Other useful consequences of our recent Examples (continued)

4. **Near a charged disk**, or in the presence of an infinite charged disk

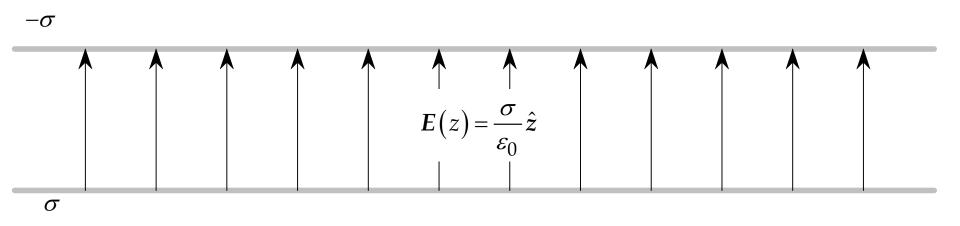
If the radius *R* of the charged disk we considered last time is very large compared to *z*, then the field along its axis winds up not depending upon *z* at all: neglect, compared to 1

$$E(z) = \frac{\sigma}{2\varepsilon_0} \hat{z} \left( 1 - \frac{z}{\sqrt{z^2 + R^2}} \right) \cong \frac{\sigma}{2\varepsilon_0} \hat{z}$$
  
neglect, compared to  $R^2$ 

-- that is, a **uniform** electric field.

## How to make a uniform electric field

This last result suggests a nice way to make a uniform electric field, should you ever need one in the lab: two uniformly-charged parallel planes, with distance small compared to length or width. Seen edge on:

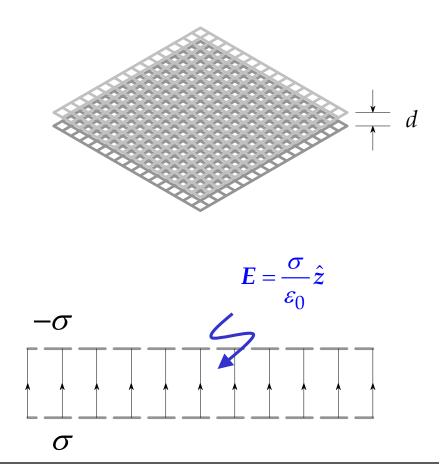


That is, just superpose the fields from two oppositely-charged planes. They add in between the planes, and cancel outside.

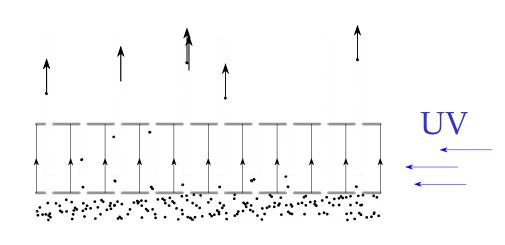
# **Application: ion drive**

Even simple, uniform *E* can be quite useful, as in the **ion drive**, the latest wrinkle in spacecraft propulsion.

- Charge up two close parallel plates that have holes in them, much smaller than the spacing between the plates but large enough for atoms to leak through.
- ❑ As we just learned, the resulting *E* is uniform and perpendicular to the plates, and zero outside.



- □ Now let some heavy, easily-ionized gas, like **xenon**, leak in from the positively-charged side.
- □ Ionize the leaked atoms, e.g. with UV light.
- □ Then the xenon ions positively charged are accelerated by *E*, and escape through the holes when they reach the other side.
  - The electrons from ionization move the opposite direction in *E*.
- The larger is *E*, the faster the ions are going when they escape.



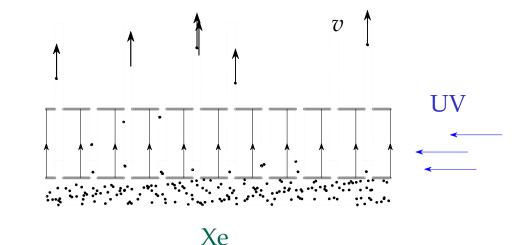
Xe

The speed of the ions can easily be calculated from what you learned in PHY 121:

$$F = QE = ma \implies a = eE/m_{Xe}; \quad d = \frac{1}{2}at^2 \implies t = \sqrt{\frac{2dm_{Xe}}{eE}}$$
$$v = at = \sqrt{\frac{2eEd}{m_{Xe}}} \qquad m_{Xe} = 2.2 \times 10^{-25} \text{ kg} \cong 124m_{\text{H}}$$

□ If *N* ions per unit time escape, they carry away momentum at the rate

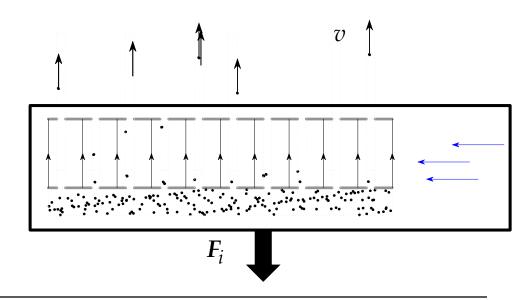
$$\frac{d\boldsymbol{p}}{dt} = N\boldsymbol{v} = N\sqrt{\frac{2eEd}{m_{Xe}}}\hat{z}$$



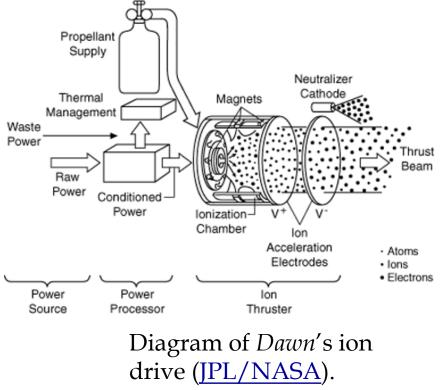
□ But momentum is conserved, so the assembly containing the plates, gas supply, UV light, *etc.* experiences a **thrust**,

$$F_i = -\frac{dp}{dt} = -N\sqrt{\frac{2eEd}{m_{Xe}}}\hat{z}$$

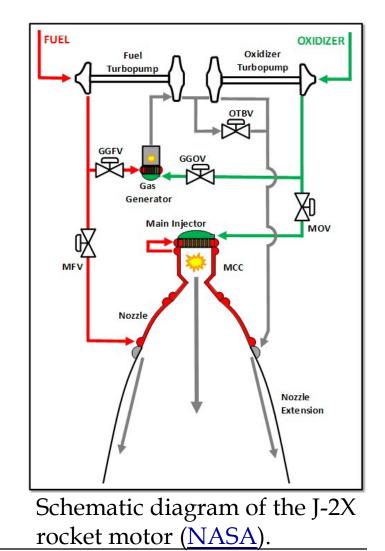
□ ...and thus the assembly (mass *M*) is propelled in the direction opposite that of the escaping ions, at a constant acceleration  $F_i/M$  which is larger, the larger is *E*.



- This is the principle behind NASA's NSTAR ion drives, which have been used on the *Deep Space 1* and *Dawn* missions to asteroids and comets 1.3-2.6 times further away from the Sun than Earth (*r* = 1.3-2.6 AU).
- The drives are powered by solar-panel generators.
- □ Typically NSTAR drives consume <u>2.3 kW of</u> <u>electrical power and</u> <u>produce 0.02 lb of thrust</u>.
- Doesn't sound like much? Read on...



- Ordinary rockets work by heating their fuel to high temperature and pressure, and letting some of this escape to produce thrust.
- Compared to ion drives, rockets are inefficient, since the ignited fuel atoms move in random directions in the ignition chamber.
- Thus an ion drive can reach higher speeds than a rocket of the same mass, though it may take longer to get to high speeds.



## Ion drive vs. rockets: tale of the tape

Let's compare the best rocket with the best ion drive, for taking a 1000 kg satellite out of the Solar system.

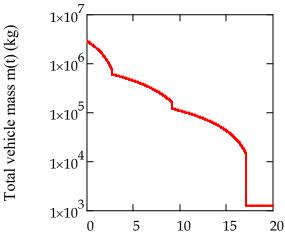
- Best rocket: the mighty Saturn V, famous from the Apollo missions.
  - Total mass of satellite plus three-stage rocket  $= 2.9 \times 10^6$  kg.
  - Initial thrust =  $3.9 \times 10^7$  N =  $8.7 \times 10^6$  lb. 1.7 x Falcon Heavy; 6 x Delta IV H; 40x Delta II H.
  - Also 1.2 x Energia; 2.8 x Ariane 5; 3.9 x Proton-M, 4.0 x Long March CZ-5, 7.6 x GSLV-III.
- Best ion drive: SAFE-400 nuclear reactor (100 kWe),
   43 NSTAR drives; total mass 600 kg.



Launch of Apollo 11, 1969, on a Saturn V booster (NASA).

Start the two vehicles with the same mass – adding Xe to the ion-drive vehicle til its mass is the same as the other – and accelerate as long as they are producing thrust. Their mass decreases as they go along, and use up fuel, so the acceleration is a function of time:

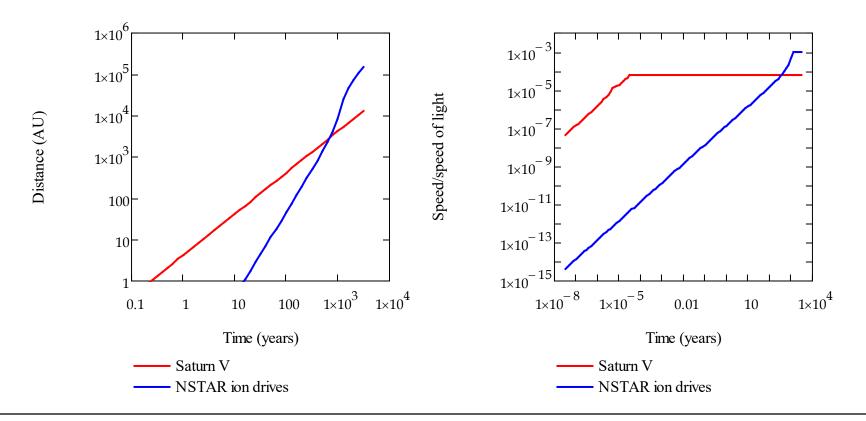
$$a(t) = F_{\text{thrust}} / m(t)$$
$$v(t) = \int_{0}^{t} a(t') dt'$$
$$x(t) = \int_{0}^{t} v(t') dt'$$



Time since start (minutes)

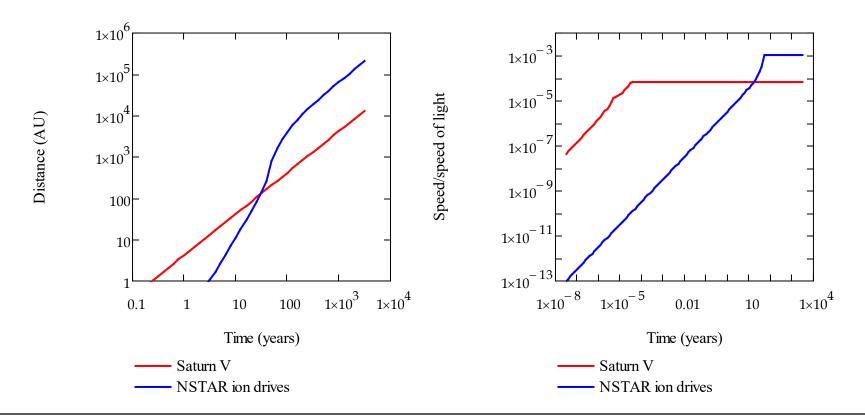
Mass of the rocket-powered vehicle, as the Saturn V expends and ejects its three stages.

Result: the Saturn V can get anywhere within the Solar system faster than *current* NSTAR ion drives, but eventually the ion drive goes farther and flies faster.



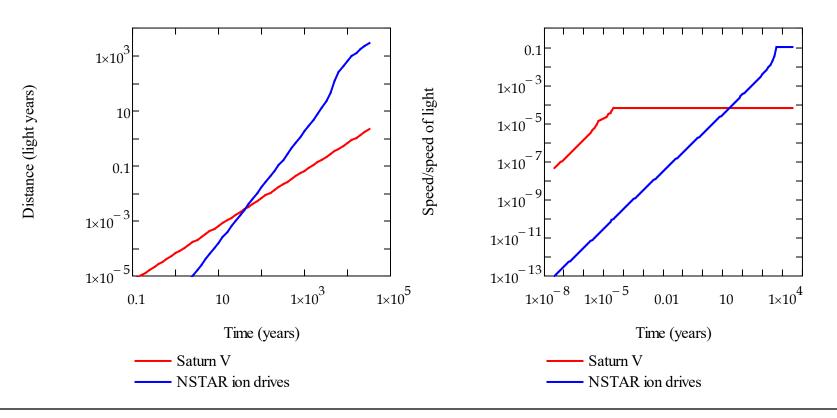
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Ion drives have much more room for improvement than rockets. Here are the results for NSTAR drives that can run on 100 W of electrical power:



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If in addition *E* could be made 10000 times larger, the ion drive's top speed begins to look respectable, at 10% the speed of light, and gets to the nearest stars in about 1500 years.



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