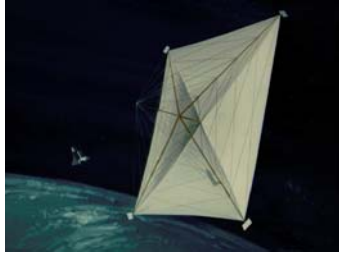


**Today in Physics 218: back to electromagnetic waves**

- ❑ Impedance
- ❑ Plane electromagnetic waves
- ❑ Energy and momentum in plane electromagnetic waves
- ❑ Radiation pressure



Artist's conception of a "solar sail": a spacecraft propelled by solar radiation pressure. (Benjamin Diedrich, Caltech.)

30 January 2004

Physics 218, Spring 2004

1

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**Reflection, transmission and impedance**

Consider a simple transverse wave again:

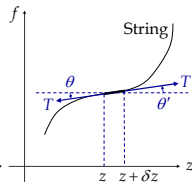
$$f(z, t) = Ae^{i(kx - \omega t)}$$

for which  $\frac{\partial f}{\partial z} = ikAe^{i(kx - \omega t)}$  ,  $\frac{\partial f}{\partial t} = -i\omega Ae^{i(kx - \omega t)}$  .

Consider again the force diagram for the string:

$$F_f(z) = T \sin \theta \cong T \tan \theta = T \frac{\partial f}{\partial z}$$

$$= T \left( -\frac{k}{\omega} \right) \frac{\partial f}{\partial t} = -\frac{T}{v} \frac{\partial f}{\partial t} \equiv -Z \frac{\partial f}{\partial t}$$



$Z = T/v = \sqrt{T\mu}$  is called the **impedance**.

30 January 2004

Physics 218, Spring 2004

2

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**Reflection, transmission and impedance (continued)**

In these terms, the two halves of our string have different impedances. We can case the reflected and transmitted wave amplitudes in this form:

$$\tilde{A}_R = \frac{v_2 - v_1}{v_2 + v_1} \tilde{A}_I = \frac{\frac{T}{Z_2} - \frac{T}{Z_1}}{\frac{T}{Z_2} + \frac{T}{Z_1}} \tilde{A}_I = \frac{Z_1 - Z_2}{Z_1 + Z_2} \tilde{A}_I$$

Similarly,  $\tilde{A}_T = \frac{2Z_1}{Z_1 + Z_2} \tilde{A}_I$  .

Changes in impedance are associated with reflection. What does the impedance "mean"? Consider the power carried by the wave past some point  $z$ :

30 January 2004

Physics 218, Spring 2004

3

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**Reflection, transmission and impedance  
(continued)**

$$P = \frac{d}{dt}(\mathbf{F} \cdot \boldsymbol{\ell}) = \mathbf{F} \cdot \frac{d\boldsymbol{\ell}}{dt} \quad (\text{since the wave is transverse})$$

$$= -F_f \frac{\partial f}{\partial t} = -F_f \left( -v \frac{\partial f}{\partial z} \right) = -\frac{v}{T} F_f \left( -T \frac{\partial f}{\partial z} \right) = \frac{1}{Z} F_f^2 \quad ,$$

or

$$= \left( T \frac{\partial f}{\partial z} \right) \frac{\partial f}{\partial t} = \frac{T}{v} \left( \frac{\partial f}{\partial t} \right)^2 = Z \left( \frac{\partial f}{\partial t} \right)^2 \quad .$$

Compare  $P = \frac{1}{Z} F_f^2 = Z \left( \frac{\partial f}{\partial t} \right)^2$  to some familiar electrical quantities:

$$P = \frac{1}{R} V^2 = RI^2 \quad (\text{DC}), \quad P = \text{Re} \left( \frac{1}{Z} V^2 \right) = \text{Re} (ZI^2) \quad (\text{AC}).$$


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30 January 2004                      Physics 218, Spring 2004                      4

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**Back to electromagnetic waves**

With strings providing an introduction, we will now return to electromagnetic waves, which you will recall involve fields that, in vacuum, obey classical wave equations:

$$\nabla^2 \mathbf{E} = \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} \quad , \quad \nabla^2 \mathbf{B} = \frac{1}{c^2} \frac{\partial^2 \mathbf{B}}{\partial t^2} \quad .$$

As with strings, the simplest solutions are the sinusoidal ones, called plane waves:

$$\tilde{\mathbf{E}}(\mathbf{r}, t) = \tilde{\mathbf{E}}_0 e^{-(kz - \omega t)} \quad , \quad \tilde{\mathbf{B}}(\mathbf{r}, t) = \tilde{\mathbf{B}}_0 e^{-(kz - \omega t)} \quad ,$$

where  $\tilde{\mathbf{E}}_0$  and  $\tilde{\mathbf{B}}_0$  are complex constants. First, a few remarks about the nature of electromagnetic plane waves in vacuum are in order:

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30 January 2004                      Physics 218, Spring 2004                      5

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**Relationship of  $\mathbf{E}$  and  $\mathbf{B}$**

□ Because the divergences of the fields are both zero in vacuum, we have, for instance,

$$\nabla \cdot \tilde{\mathbf{E}} = \frac{\partial}{\partial z} \tilde{E}_{0z} e^{i(kz - \omega t)} \quad \text{No } x \text{ or } y \text{ dependence}$$

$$= ik \tilde{E}_{0z} e^{i(kz - \omega t)} = 0 \quad \Rightarrow \quad \tilde{E}_{0z} = 0.$$

Similarly,  $\tilde{B}_{0z} = 0$ . Electromagnetic waves in vacuum have no field components along their direction of propagation; they're transverse, just like waves on a string.

□ We learn something from the curls, too:

$$\nabla \times \tilde{\mathbf{E}} = -\frac{\partial \tilde{E}_y}{\partial z} \hat{x} + \frac{\partial \tilde{E}_x}{\partial z} \hat{y} \quad \text{All other terms vanish}$$


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30 January 2004                      Physics 218, Spring 2004                      6

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**Relationship of E and B (continued)**

$$\nabla \times \tilde{\mathbf{E}} = -ik\tilde{E}_{0y}e^{i(kz-\omega t)}\hat{x} + ik\tilde{E}_{0x}e^{i(kz-\omega t)}\hat{y}$$

$$= -\frac{1}{c}\frac{\partial \tilde{\mathbf{B}}}{\partial t} = \frac{i\omega}{c}\tilde{\mathbf{B}}_0e^{i(kz-\omega t)}$$

$$= \frac{i\omega}{c}\tilde{B}_{0x}e^{i(kz-\omega t)}\hat{x} + ik\tilde{B}_{0y}e^{i(kz-\omega t)}\hat{y}.$$

But  $k = \omega/c$ , so

$$\tilde{E}_{0x} = \tilde{B}_{0y}, \quad \tilde{E}_{0y} = -\tilde{B}_{0x}$$

which has interesting implications for the relative magnitude, phase, and orientation of the fields:

30 January 2004 Physics 218, Spring 2004 7

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**Relationship of E and B (continued)**

the field amplitudes have the **same magnitude**, as

$$|\tilde{\mathbf{E}}_0| = \sqrt{E_{0x}^2 + E_{0y}^2} = \sqrt{B_{0y}^2 + B_{0x}^2} = |\tilde{\mathbf{B}}_0| ;$$

are **perpendicular** to one another, as

$$\tan \phi = \frac{\tilde{E}_{0y}}{\tilde{E}_{0x}}, \quad \tan \phi' = \frac{\tilde{B}_{0y}}{\tilde{B}_{0x}} = -\frac{\tilde{E}_{0x}}{\tilde{E}_{0y}} = -\tan \phi = \tan\left(\phi + \frac{\pi}{2}\right);$$

and are **in phase**, as we see by combining the last three results:

$$\tilde{\mathbf{B}}_0 = \hat{z} \times \tilde{\mathbf{E}}_0 .$$

In MKS, we'd get  $|\tilde{\mathbf{E}}_0| = c|\tilde{\mathbf{B}}_0|$ , but the fields would still be perpendicular and in phase.

30 January 2004 Physics 218, Spring 2004 8

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**Relationship of E and B (continued)**

Snapshot of the fields - in principle impossible to do, but if you could this is what you'd see)

30 January 2004 Physics 218, Spring 2004 9

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**Energy and momentum in electromagnetic waves**

More generally, if the wave propagates in a direction other than  $z$  (i.e. the wavenumber is a vector,  $\hat{k}$ ), then

$$\vec{E} = \vec{E}_0 e^{i(\hat{k}\cdot\vec{r} - \omega t)} \quad , \quad \vec{B} = \hat{k} \times \vec{E} \quad .$$

Taking  $E$  and  $B$  to be the real parts of the plane-wave fields  $\vec{E}$  and  $\vec{B}$ , we get the following for the energy density, and energy flux:

$$u = \frac{1}{8\pi} (E^2 + B^2) = \frac{E^2}{4\pi} = \frac{B^2}{4\pi} = \epsilon_0 E^2 = \frac{B^2}{\mu_0} \quad \text{in MKS.}$$

$$S = \frac{c}{4\pi} E \times B = \frac{c}{4\pi} E \times (\hat{k} \times E) \quad \text{Use Product Rule \#2:}$$

$$= \frac{c}{4\pi} \hat{k} (E \cdot E) - \frac{c}{4\pi} E (\hat{k} \cdot E) = \frac{cE^2}{4\pi} \hat{k} = cu\hat{k} \quad . \quad \text{Last one same in MKS.}$$

30 January 2004      Physics 218, Spring 2004      10

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**Energy and momentum in electromagnetic waves (continued)**

For the momentum density,

$$\vec{g} = \frac{S}{c^2} = \frac{E^2}{4\pi c} \hat{k} = \frac{u}{c} \hat{k} \quad . \quad \text{Last one same in MKS.}$$

So, unsurprisingly, an electromagnetic wave carries energy and momentum in its direction of motion.

In the case of light, the periods of oscillation are so short that time averages of these quantities is useful. Recall that the average over a period for the square of a cosine or a sine is

$$\langle \cos^2 \omega t \rangle \equiv \frac{\omega}{2\pi} \int_0^{2\pi/\omega} \cos^2 \omega t dt = \frac{1}{2} = \langle \sin^2 \omega t \rangle \quad .$$

30 January 2004      Physics 218, Spring 2004      11

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**Flashback (part e, problem 7.54)**

Remember the orthonormality of sines and cosines?

$$\int_0^{2\pi} \cos mx \cos nx dx = \int_0^{2\pi} \sin mx \sin nx dx = \pi \delta_{mn} \quad ,$$

$$\int_0^{2\pi} \cos mx \sin nx dx = 0 \quad , \quad \text{so}$$

$$\langle \cos^2 \omega t \rangle = \frac{\omega}{2\pi} \int_0^{2\pi/\omega} \cos^2 \omega t dt = \frac{1}{2\pi} \int_0^{2\pi} \cos^2 x dx = \frac{1}{2} \quad ,$$

and, similarly,

$$\langle \sin^2 \omega t \rangle = \frac{1}{2} \quad \text{and} \quad \langle \cos \omega t \sin \omega t \rangle = 0 \quad .$$

30 January 2004      Physics 218, Spring 2004      12

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**Energy and momentum in electromagnetic waves  
(continued)**

Thus,  $\langle u \rangle = \left\langle \frac{E^2}{4\pi} \right\rangle = \frac{E_0^2}{4\pi} \langle \cos^2(k \cdot r - \omega t) \rangle = \frac{E_0^2}{8\pi}$  .

Similarly,

$$\langle \mathbf{S} \rangle = \langle c u \hat{\mathbf{k}} \rangle = \frac{c E_0^2}{8\pi} \hat{\mathbf{k}} \equiv I \hat{\mathbf{k}} \quad , \quad I = \text{intensity}$$

$$\langle \mathbf{g} \rangle = \left\langle \frac{u}{c} \hat{\mathbf{k}} \right\rangle = \frac{E_0^2}{8\pi c} \hat{\mathbf{k}} \quad .$$


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30 January 2004 Physics 218, Spring 2004 13

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**Radiation pressure**

When light falls encounters a surface bounding a medium, it can be absorbed or reflect, in either case imparting its momentum to the medium and thus exerting a pressure. The momentum per unit area crossing an area  $A$  during a time  $\Delta t$  is

$$\Delta p = \langle \mathbf{g} \rangle c \Delta t \quad .$$

If the medium is a perfect absorber, the pressure is therefore

$$P_{\text{abs}} = \frac{1}{A} \frac{dp}{dt} = \frac{1}{A} \frac{\Delta p}{\Delta t} = c \langle \mathbf{g} \rangle = \frac{E_0^2}{8\pi} = u = \frac{I}{c} \quad .$$

If the medium's surface is a perfect reflector, the light rebounds elastically and the pressure is twice as high:

$$P_{\text{abs}} = 2P_{\text{ref}} = \frac{E_0^2}{4\pi} \quad .$$


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30 January 2004 Physics 218, Spring 2004 14

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**Example: problem 9.10**

*The intensity of sunlight hitting the Earth is about 1300 W/m<sup>2</sup>. If sunlight strikes a perfect absorber, what pressure does it exert? How about a perfect reflector? What fraction of atmospheric pressure does this amount to?*

In cgs units, that's  $1.3 \times 10^6$  erg sec<sup>-1</sup> cm<sup>-2</sup> :

$$P_{\text{abs}} = \frac{I}{c} = 4.3 \times 10^{-5} \text{ dyne cm}^{-2} \quad .$$

It's twice that much for a reflector. Either way it's not much compared to atmospheric pressure,

$$P_{\text{atm}} = 1.013 \times 10^6 \text{ dyne cm}^{-2}; \quad \frac{P_{\text{abs}}}{P_{\text{atm}}} = 4.3 \times 10^{-11} \quad .$$


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30 January 2004 Physics 218, Spring 2004 15

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