

Physics 113 - October 16, 2012

Midterm class survey

Exam 2 - Oct. 25, during regular lecture slot in Hoyt

From Newton's laws w/ friction to start of Momentum Cons.

9/25 lect to end of gravity in this lecture

chapters 5, 6, 7, 8 → Not sections

Probsets 5, 6

5-5, 5-6, 6-5, 6-7, 6-8

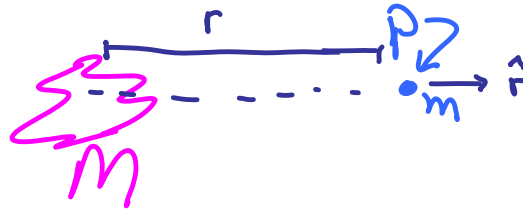
Workshops 4-6

No lecture in Hoyt Tues., Oct. 23

↳ Slides + Audio will be posted

Gravitation

Last Time —

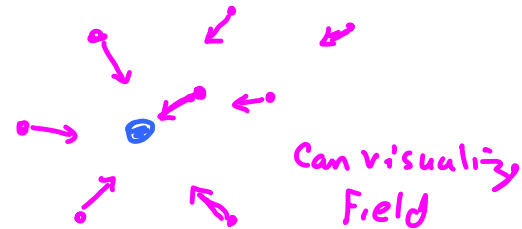


$$\vec{F}_{\text{of } M \text{ on } m} = -\frac{GMm}{r^2} \hat{r}$$

gravitational field

$$\vec{g}(p) = \vec{F}/m = -\frac{GM}{r^2} \hat{r}$$

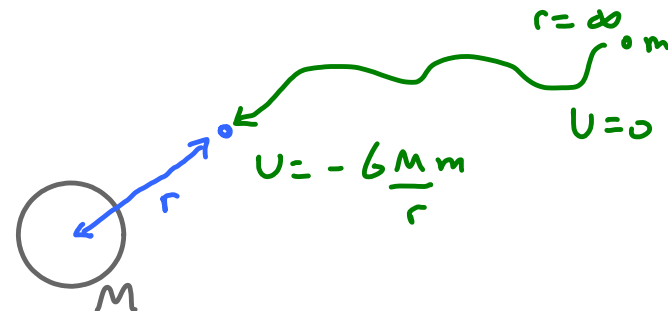
Vector

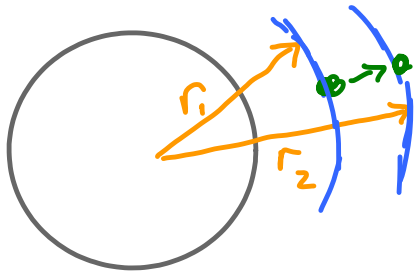


for mass brought in from ∞ to point P

$$PE_{\text{grav}} = -\frac{GMm}{r}$$

Scalar

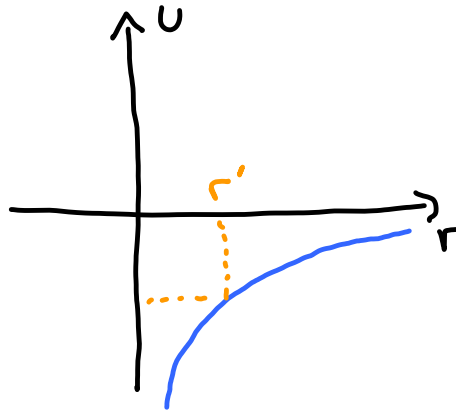
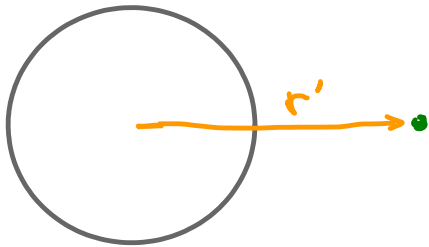




$$\Delta PE = -\frac{GMm}{r_2} - \left(-\frac{GMm}{r_1}\right)$$



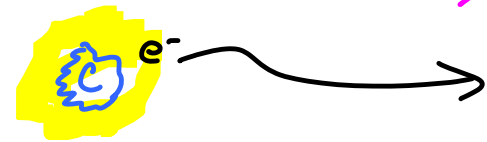
$\approx mg\Delta r \approx mgh$ for small h near Surf. of Earth

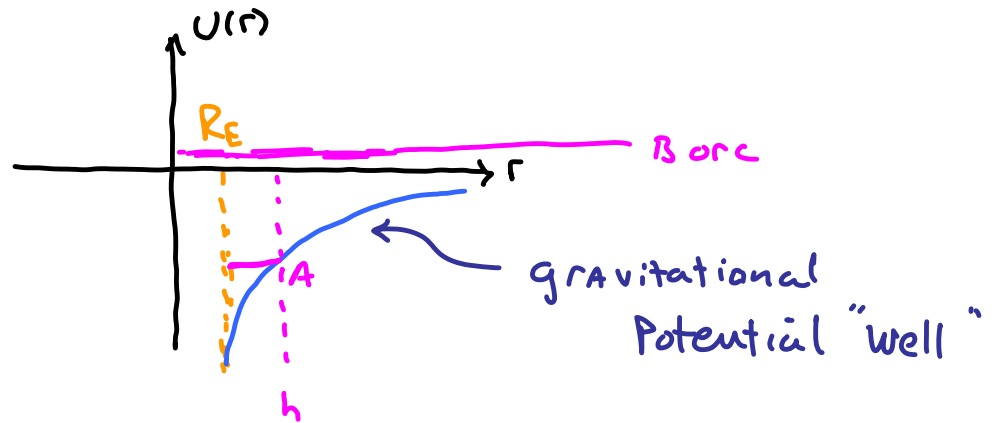
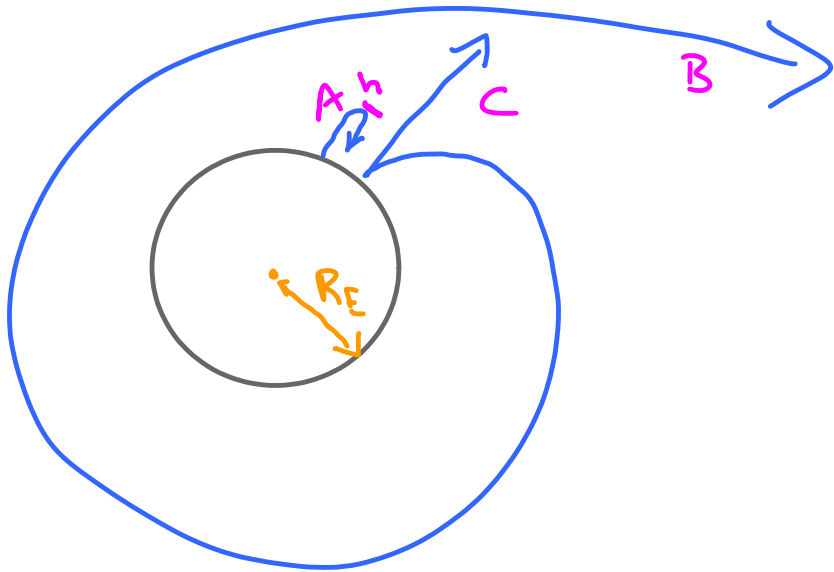


What does it mean that U is negative?

must put in $\frac{GMm}{r}$ of energy to free particle

Analogous to ionization energy in chemistry





Escape Velocity, V_{es}

$$\frac{1}{2} m V_{es}^2 = \frac{GMm}{r}$$

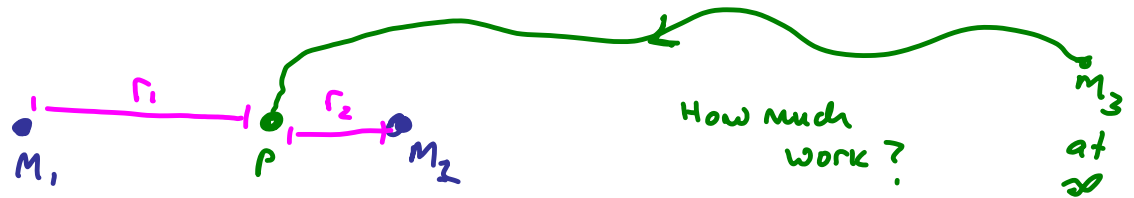
KE
 PE of "bound" state

From Surface of Earth

$$\frac{1}{2} m V_{es}^2 = \frac{GM_E m}{R_E}$$

$$V_{es} = \sqrt{\frac{2GM_E}{R_E}}$$

Why potential?



$$\text{Energy to move } M_3 \text{ to point } P = \sum U_i = -\frac{GM_1 M_3}{r_1} - \frac{GM_2 M_3}{r_2}$$

Easier to calculate than \vec{F}_g

$$F_s = -\frac{dU_s}{ds}$$

Very imp't in Electricity + Magnetism "Electrical potential"
"potential difference"

End of Material for EXAM 2

$$\vec{F} = m\vec{a} = m \frac{d\vec{v}}{dt} = \frac{d(m\vec{v})}{dt}$$

$$\underbrace{\frac{d(m\vec{v})}{dt}} = m \frac{d\vec{v}}{dt} + \underbrace{\frac{dm}{dt} \vec{v}}_0 \quad \text{not zero if mass } \Delta'es$$

$$\frac{d(m\vec{v})}{dt}$$

$$m\vec{v} \equiv \vec{p} \equiv \text{momentum}$$

$$\rightarrow \text{vector} \quad \vec{p} = m\vec{v}$$

$$p_x = mv_x$$

$$p_y = mv_y$$

$$p_z = mv_z$$

$$\vec{F} = \frac{d\vec{p}}{dt}$$

$$d\vec{p} = \vec{F} dt$$

$$\int d\vec{p} = \int \vec{F} dt$$

$$\Delta\vec{p} = \vec{p}_2 - \vec{p}_1 = \int_{t_1}^{t_2} \vec{F} dt$$

Impulse



$$\Delta \vec{P}_1 = \int_{t_i}^{t_f} \vec{F}_{2 \text{ on } 1} dt$$

$$\Delta \vec{P}_2 = \int_{t_i}^{t_f} \vec{F}_{1 \text{ on } 2} dt$$

$$\vec{F}_{2 \text{ on } 1} = -\vec{F}_{1 \text{ on } 2}$$

$$\Delta \vec{P}_1 = -\Delta \vec{P}_2$$

$$\Delta \vec{p}_1 + \Delta \vec{p}_2 = 0$$

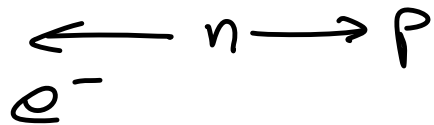
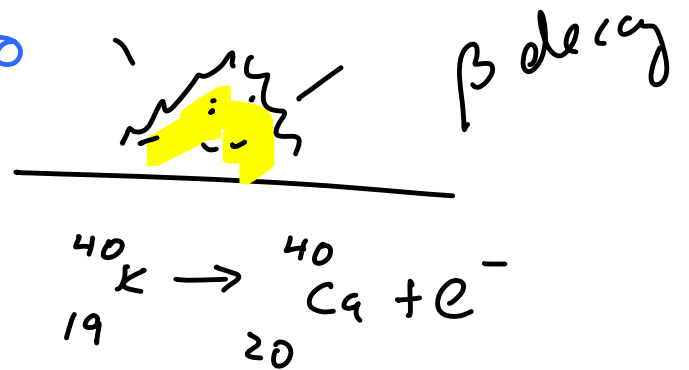
Sum of changes in $\vec{p} = 0$

MOMENTUM CONSERVATION

If there are no external forces acting on a system, the Total Momentum of the system is conserved

$$\sum \vec{p}_i = \sum \vec{p}_f$$

The discovery of the neutrino



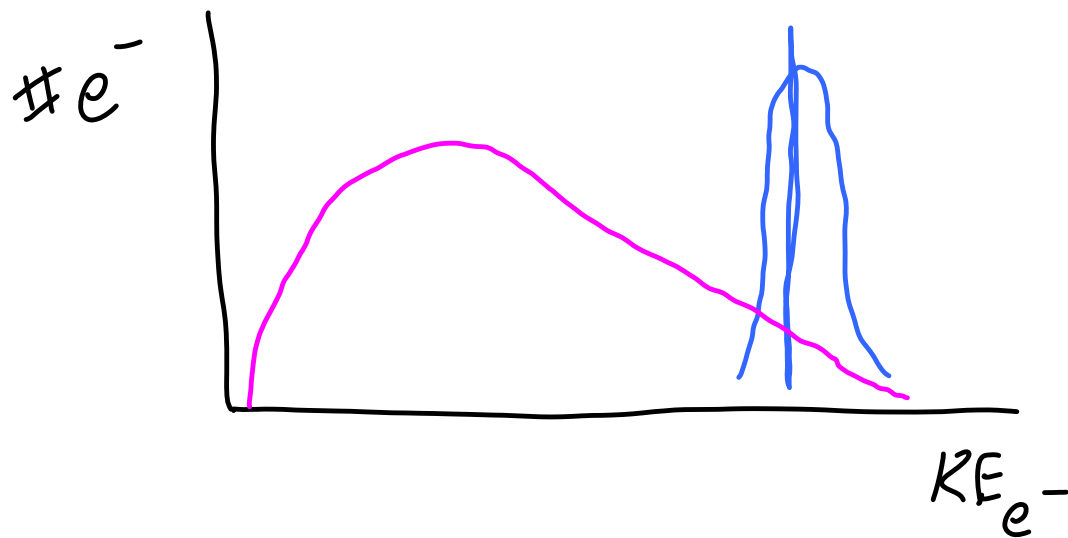
Cons

$$m_e v_e = -m_p v_p \rightarrow v_p = -\frac{m_e}{m_p} v_e$$

$$\text{Total KE} \rightarrow Q = KE_e + KE_p = \frac{1}{2} m_e v_e^2 + \frac{1}{2} m_p v_p^2$$

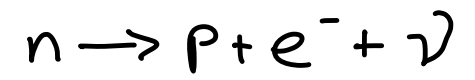
$$Q = KE_e \left[\frac{m_e}{m_p} + 1 \right]$$

$$KE_e = \frac{Q M_p}{M_p + M_e} \rightarrow \text{fixed \#}$$



1930-31

Pauli



neutrino

1959 Reines

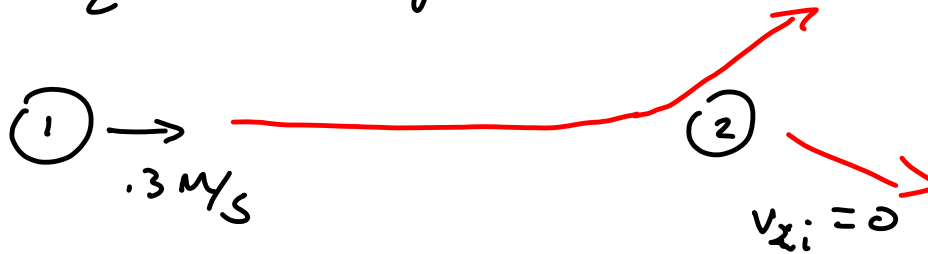
Elastic collision \rightarrow \vec{p} conserved, KE conserved

inelastic collision \rightarrow \vec{p} conserved

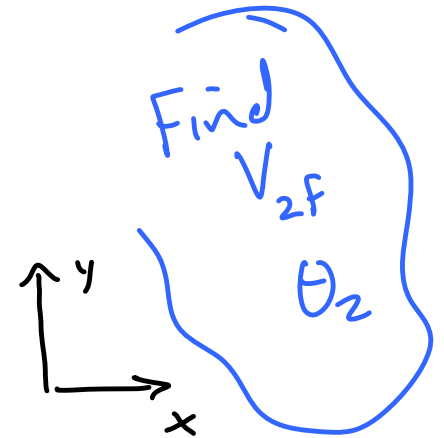
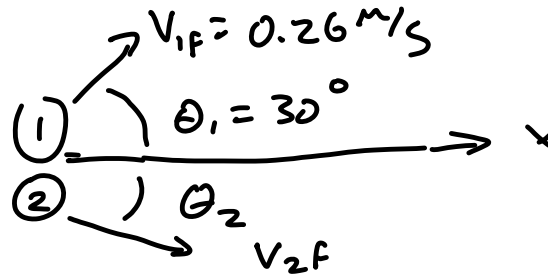
KE NOT conserved

$$M_1 = M_2 = 0.1 \text{ kg}$$

initial



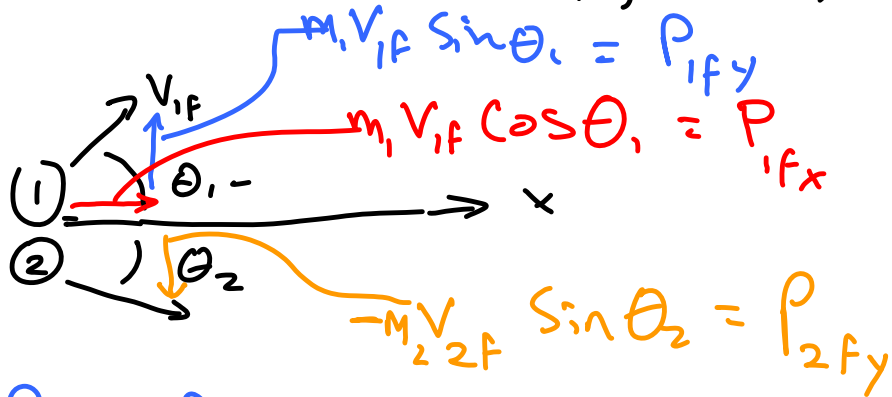
final state



$$\sum \vec{P}_i = \sum \vec{P}_f \longrightarrow \sum P_{ix} = \sum P_{fx}$$

$$\sum P_{iy} = \sum P_{fy}$$

x eqn
=



$$P_{1xi} + P_{2xi} = P_{1xf} + P_{2xf}$$

$$m_2 V_{2f} \cos \theta_2 = P_{2fx}$$

$$m_1 V_{1i} = m_1 V_{1f} \cos \theta_1 + m_2 V_{2f} \cos \theta_2$$

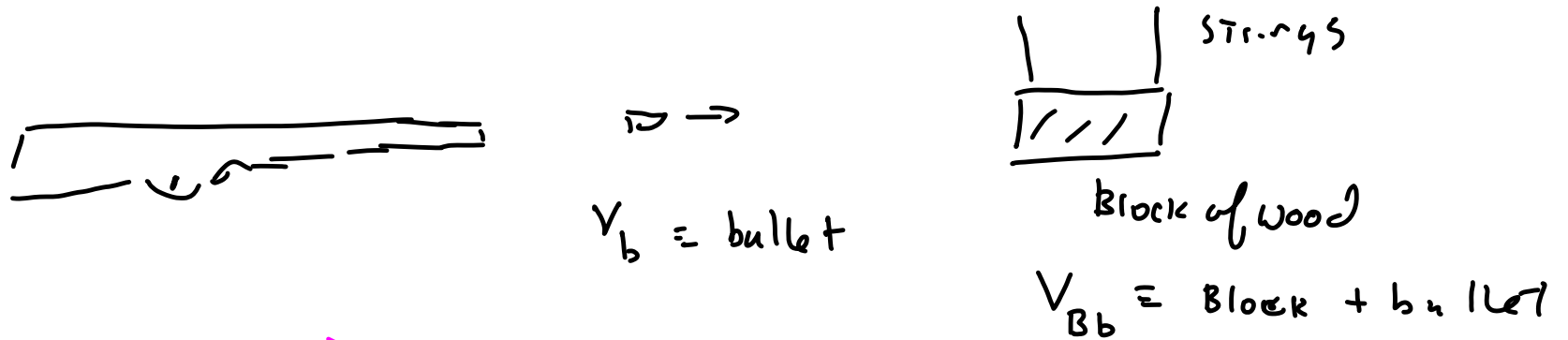
y eqn

2 eqns
2 unknowns

$$0 = m_1 v_{1f} \sin \theta_1 - m_2 v_{2f} \sin \theta_2$$

$$|v_{2f}| = 0.15 \text{ m/s}$$

$$\theta_2 = 60^\circ \text{ down from } +x \text{ axis}$$



inelastic
collision

~~$$\frac{1}{2} m_b v_b^2 = \frac{1}{2} (m_b + m_B) v_{Bb}^2$$~~

\vec{p} conservation is useful

$$m_b v_b = (m_b + m_B) v_{Bb}$$

