

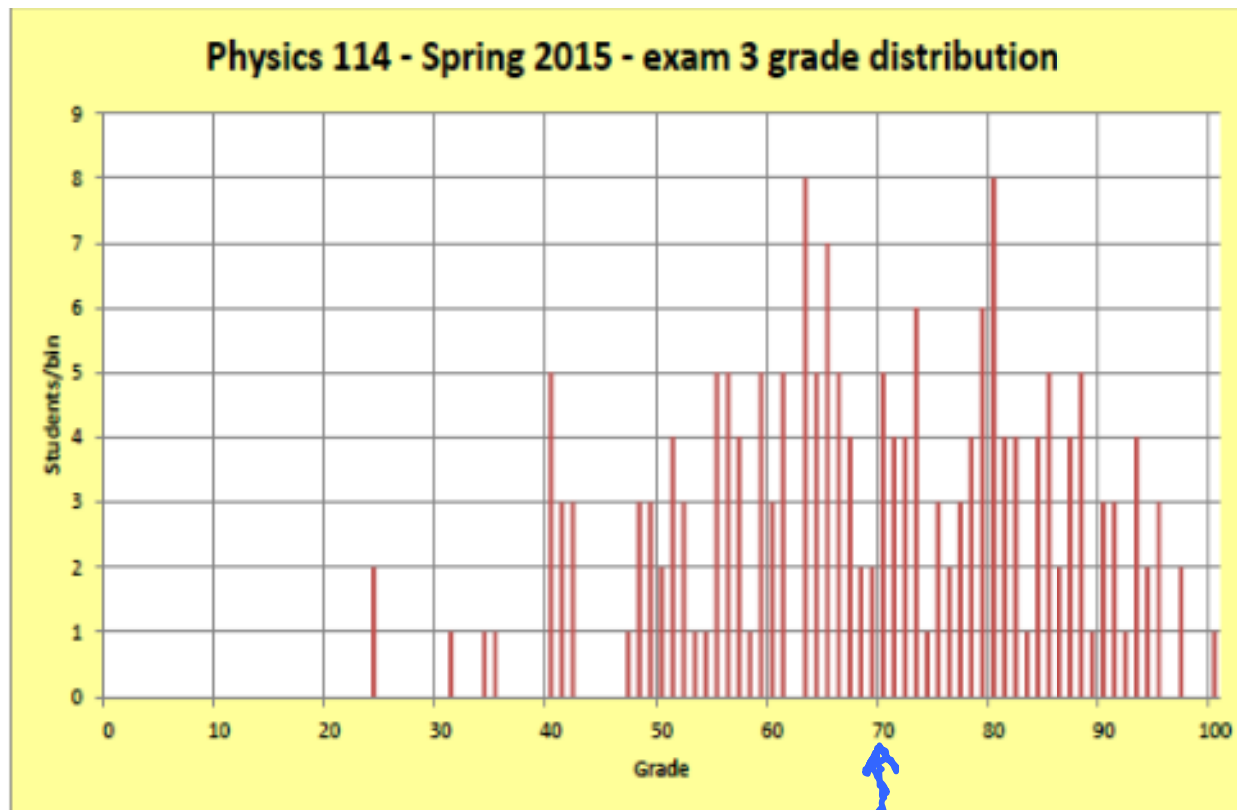
Physics 114- April 23, 2015

Note Title

4/20/2010

- Last class next Tuesday
- Last problem set to hand in
This Friday
- Will give you selected problems/solns
in last set of topics - Not for
handing in.
- EXAM 3 graded

Exam 3 - Mean 69, Median 70



Quantum Mechanics

1-d time independent Schrödinger equation

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi(x)}{dx^2} + U \psi(x) = E \psi(x)$$

Potential energy function

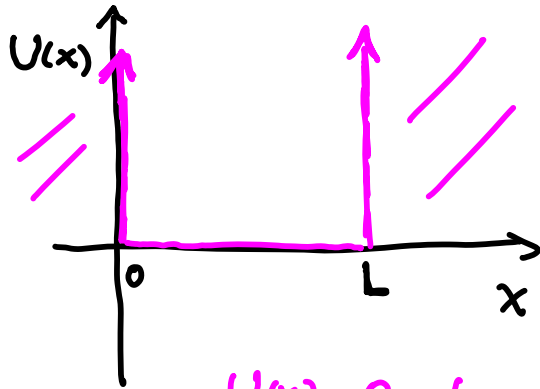
Plug in U and solve for E, ψ

$|\psi|^2 \approx$ Probability distribution for particle

$E \equiv$ Allowed energy states for particle

discrete if potential is negative
(force is attractive)

∞ Square well potential



$$\frac{d^2 \psi(x)}{dx^2} = -\frac{2mE}{\hbar^2} \psi(x)$$

$\equiv k^2$

$$U(x) = 0 \text{ for } 0 < x < L$$

$$U(x) = \infty \text{ for } x < 0, L < x$$

Solve eqn $\rightarrow \psi(x) = A \sin kx + B \cos kx$

Boundary conditions $\rightarrow \psi(0) = 0, \psi(L) = 0$

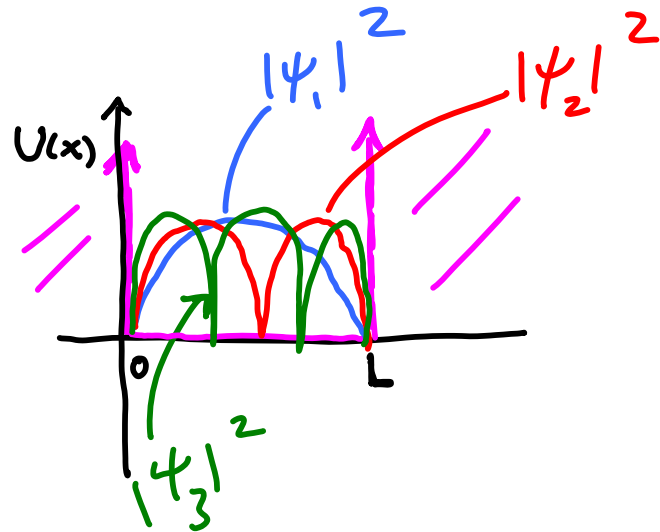
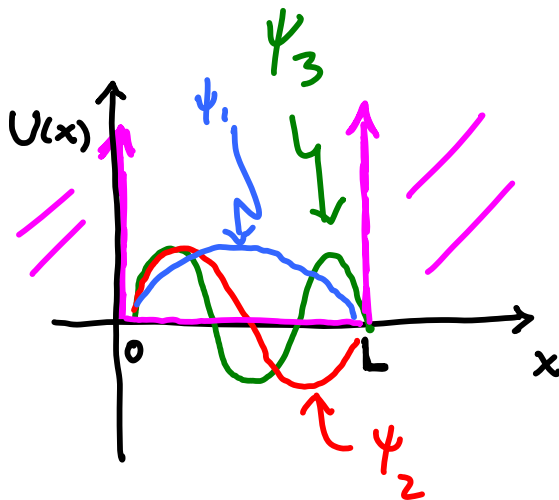
$\psi(x) = A \sin kx$

$k_n = \frac{n\pi}{L}$

Quantization happens here

$E_n = \frac{n^2 \pi^2 \hbar^2}{L^2 2m} \quad n = 1, 2, 3, \dots$

$\psi_n(x) = A \sin\left(\frac{n\pi}{L}x\right) \quad n = 1, 2, 3, \dots$



Attractive Potentials \rightsquigarrow discrete
STATES + Energies
Allowed



lead to quantization!

Go to QM and Atoms Slides -

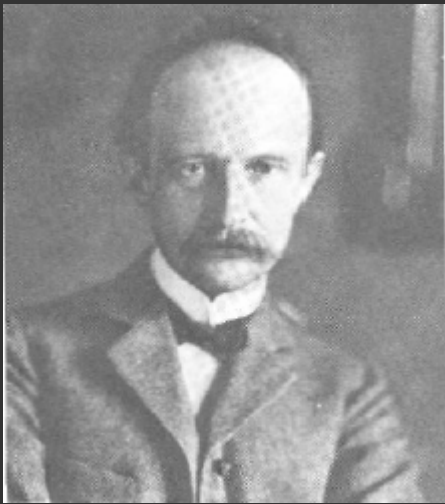
A bit on quantum mechanics and atomic physics

Physics 114



References and photo sources:

K. Krane, *Modern Physics*, John Wiley and Sons, 1983



**Max Planck (1858-1947) – 1918 Nobel Prize
for work on spectral distribution of
radiation (blackbody radiation)**



**Louis deBroglie (1892-1987)
First suggested matter has
wavelike properties**

**Three of the
players**



**Erwin Schrodinger (1887-1961) –
Developed mathematical theory of wave
mechanics that permitted the calculation
of physical systems**

**Earnest Rutherford (1871-1937)
nuclear “plantetary” model of atom**

**Niels Bohr (1885-1962) developed a
semi-classical nuclear model of the
single electron atom**



Time-independent Schrodinger equation

$$\frac{-\hbar^2}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} + V(x) \psi(x) = E \psi(x)$$

← KE Term ← PE TERM ← TOT E

$\psi(x) \equiv$ Wave function of particle

What is $\psi(x)$?

$|\psi(x)|^2 dv =$ prob. of finding particle
in volume dv

$$\int_{\text{All SPACE}} |\psi(x)|^2 dv = 1 \quad \text{particle is someplace}$$

Sub in V as appropriate + solve

for H atom

Must be generalized to

3d, spherical coordinates



$$V(r) \rightarrow \frac{1}{4\pi\epsilon_0} \frac{|q|^2}{r^2} + \text{Solve}$$

$$\frac{-\hbar^2}{2\mu} \left[\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \psi(r)}{\partial r} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \psi(r)}{\partial \phi^2} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial \psi(r)}{\partial \theta} \right) \right]$$

$$+ \frac{1}{4\pi\epsilon_0} \frac{|e|^2}{r^2} \psi(r) = E \psi(r)$$

Now
Solve



Separates into r, θ, ϕ eqns

Energy or principal quantum number

$n = 1, 2, 3 \dots$

Orbital quantum number

$l = 0, 1, \dots n-1$

Magnetic quantum number

$-l, -|l-1|, \dots, 0, 1, \dots l-1, l$

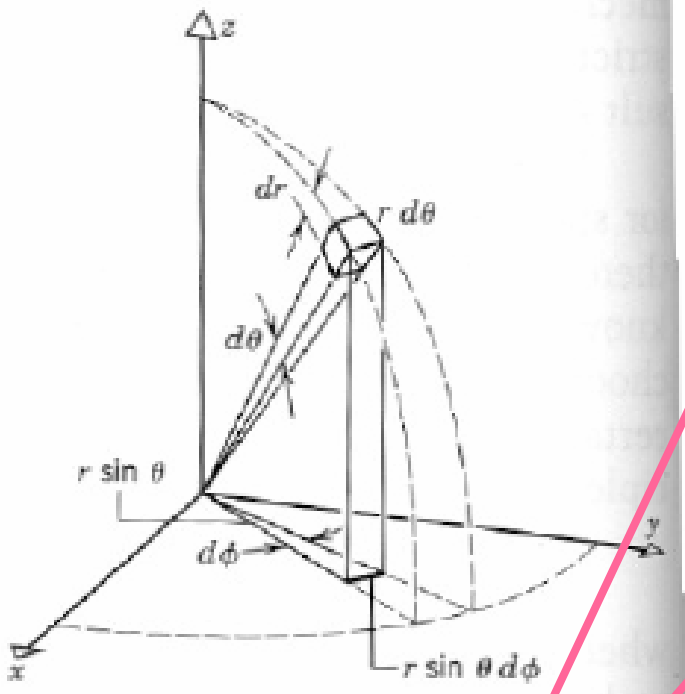
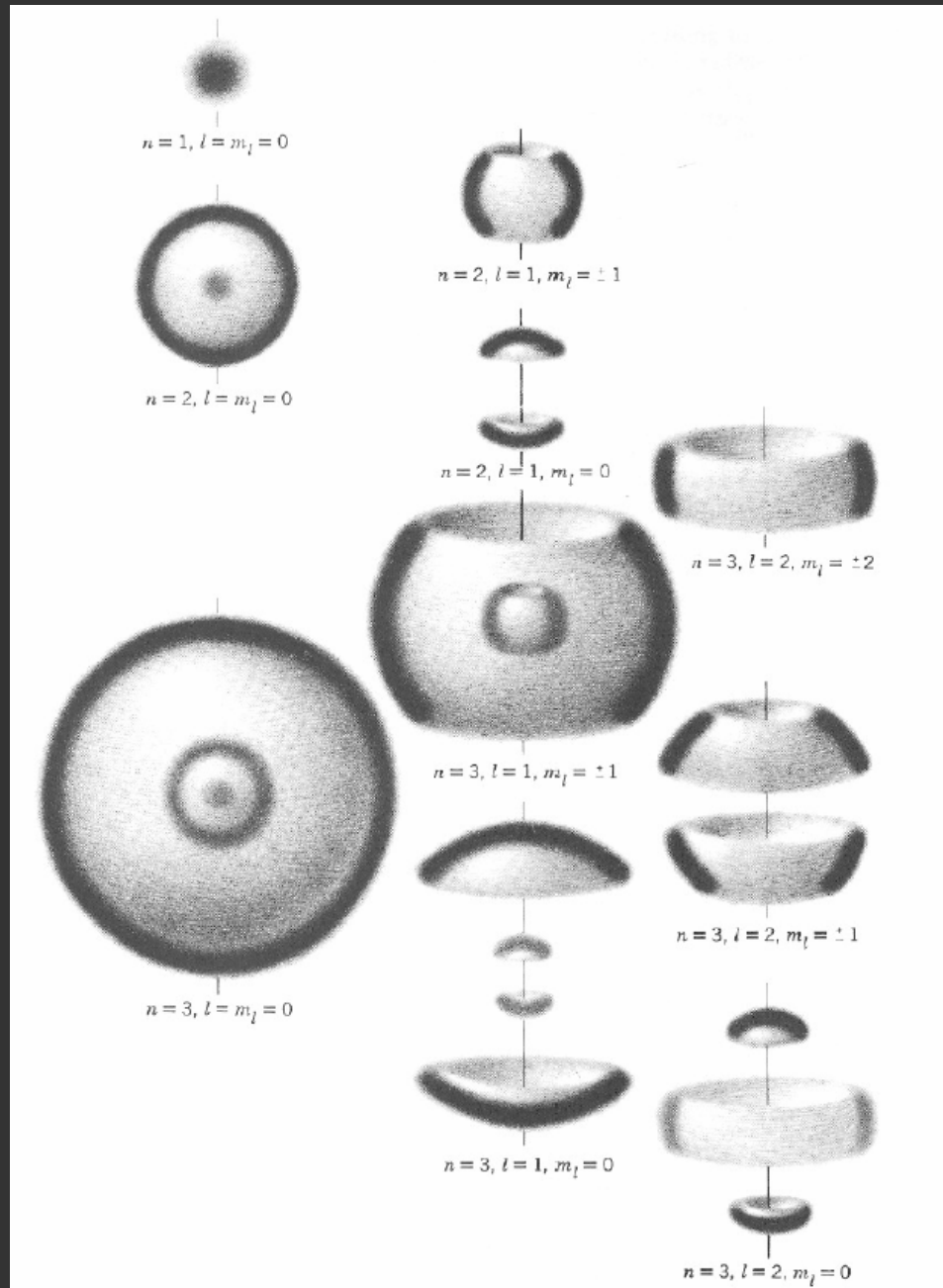


Table 7.1 Some Hydrogen Atom Wave Functions

n	l	m_l	$R(r)$	$\Theta(\theta)$	$\Phi(\phi)$
1	0	0	$\frac{2}{a_0^{3/2}} e^{-r/a_0}$	$\frac{1}{\sqrt{2}}$	$\frac{1}{\sqrt{2\pi}}$
2	0	0	$\frac{1}{(2a_0)^{3/2}} \left(2 - \frac{r}{a_0}\right) e^{-r/2a_0}$	$\frac{1}{\sqrt{2}}$	$\frac{1}{\sqrt{2\pi}}$
2	1	0	$\frac{1}{\sqrt{3}(2a_0)^{3/2}} \frac{r}{a_0} e^{-r/2a_0}$	$\sqrt{\frac{3}{2}} \cos \theta$	$\frac{1}{\sqrt{2\pi}}$
2	1	± 1	$\frac{1}{\sqrt{3}(2a_0)^{3/2}} \frac{r}{a_0} e^{-r/2a_0}$	$\frac{\sqrt{3}}{2} \sin \theta$	$\frac{1}{\sqrt{2\pi}} e^{\pm i\phi}$

Probability distributions for several allowed atomic states for the 1-electron atom

Increasing n adds new radial layers, $l=0$ give spherical symmetry, l not 0 brings in angular dependence



General Quant. Mech. result regarding force on magnetic dipole in a non-uniform magnetic field

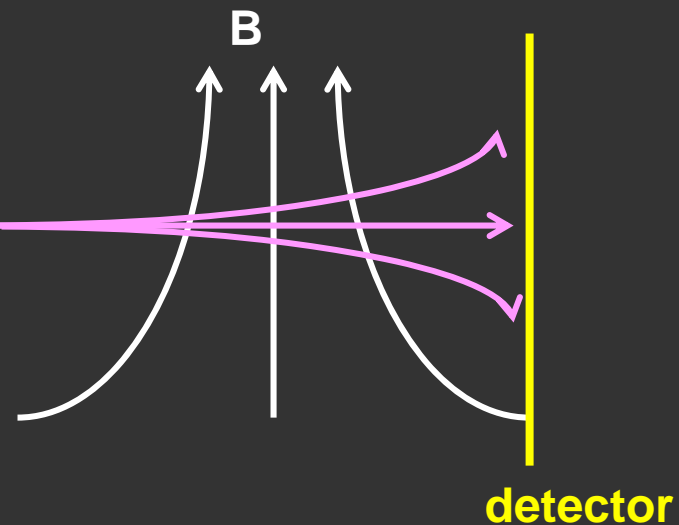
$$\vec{F}_z = \frac{\partial B_z}{\partial z} |\vec{\mu}_z| = \frac{\partial B_z}{\partial z} m$$

Stern-Gerlach experiment

e- beam in $l=1$ state

has $m=1,0,-1$ components

expect to see this



General Quant. Mech. result regarding force on magnetic dipole in a non-uniform magnetic field

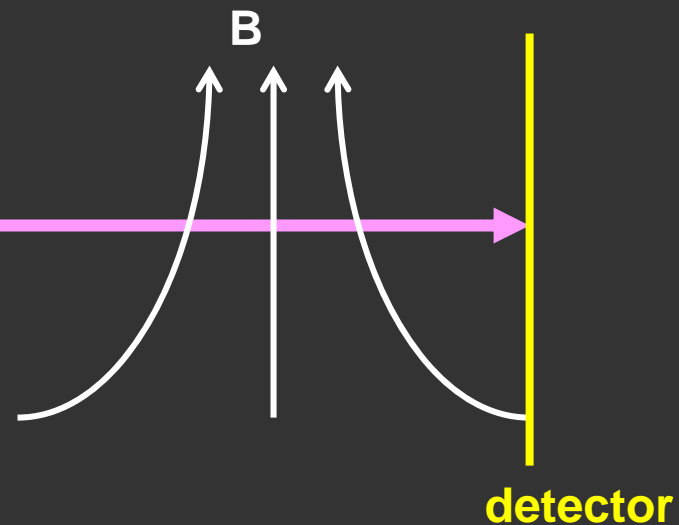
$$\vec{F}_z = \frac{\partial B_z}{\partial z} |\vec{\mu}_z| = \frac{\partial B_z}{\partial z} m$$

Stern-Gerlach experiment

e- beam in $l=0$ state

Has $m=0$ component only

expect to see this



SURPRISE! ... fundamental particles have an intrinsic magnetic moment. Call it spin.

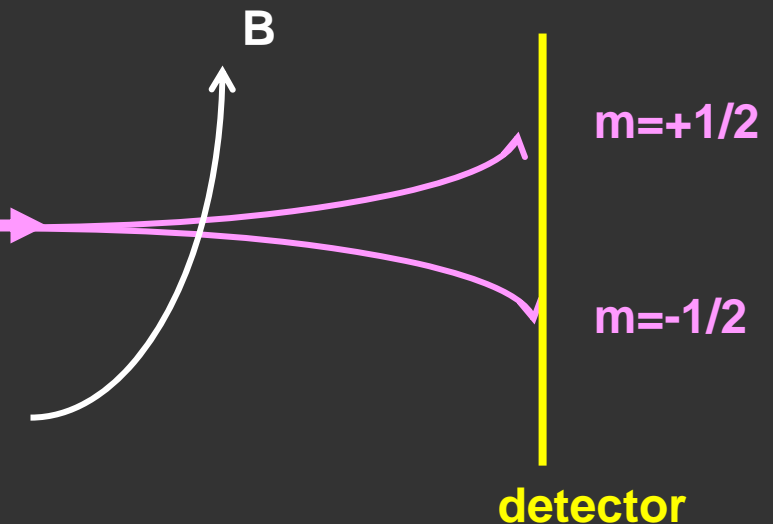
$$\vec{F}_z = \frac{\partial B_z}{\partial z} |\vec{\mu}_z| = \frac{\partial B_z}{\partial z} m$$

Stern-Gerlach experiment

e- beam in $l=0$ state

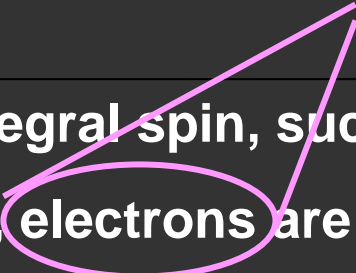
Has $m=0$ component only

Actually see this



Intrinsic spin - two varieties

Huge effect on
multi-electron
atoms



Fermions = half integral spin, such as $1/2, 3/2, 5/2, \dots, 73/2 \dots$
protons, neutrons, **electrons** are all fermions ($s=1/2$)
no two fermions can occupy the same exact quantum state

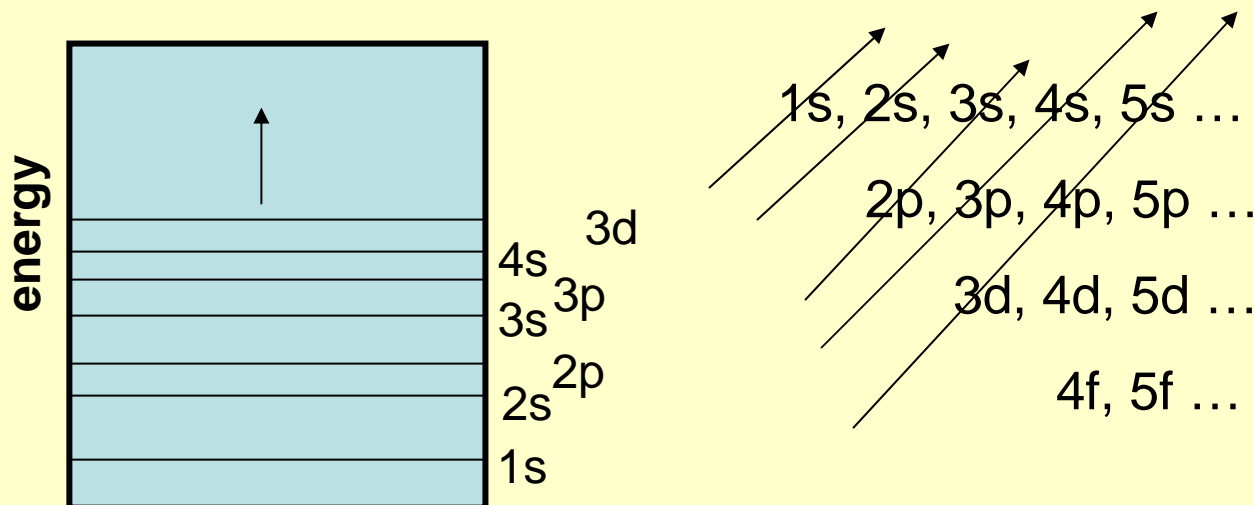
Bosons = integral spin, such as $0, 1, 2 \dots$
photons ($s=1$) and pions ($s=0$) are examples of bosons
bosons can occupy the same exact quantum state

Rules for Filling of state for multi-electron atom

n, l, m_l, m_s

Spectroscopic notation - s: $l=0$, p: $l=1$, d: $l=2$, f: $l=3$, ...

- No two electrons in same state (Pauli exclusion)
- Electrons go into the state with the lowest possible energy (Aufbau)
- Within a sublevel, electrons will have their spin unpaired as much as possible (due to spin-spin interaction contribution to energy)



Chemistry now "solved"

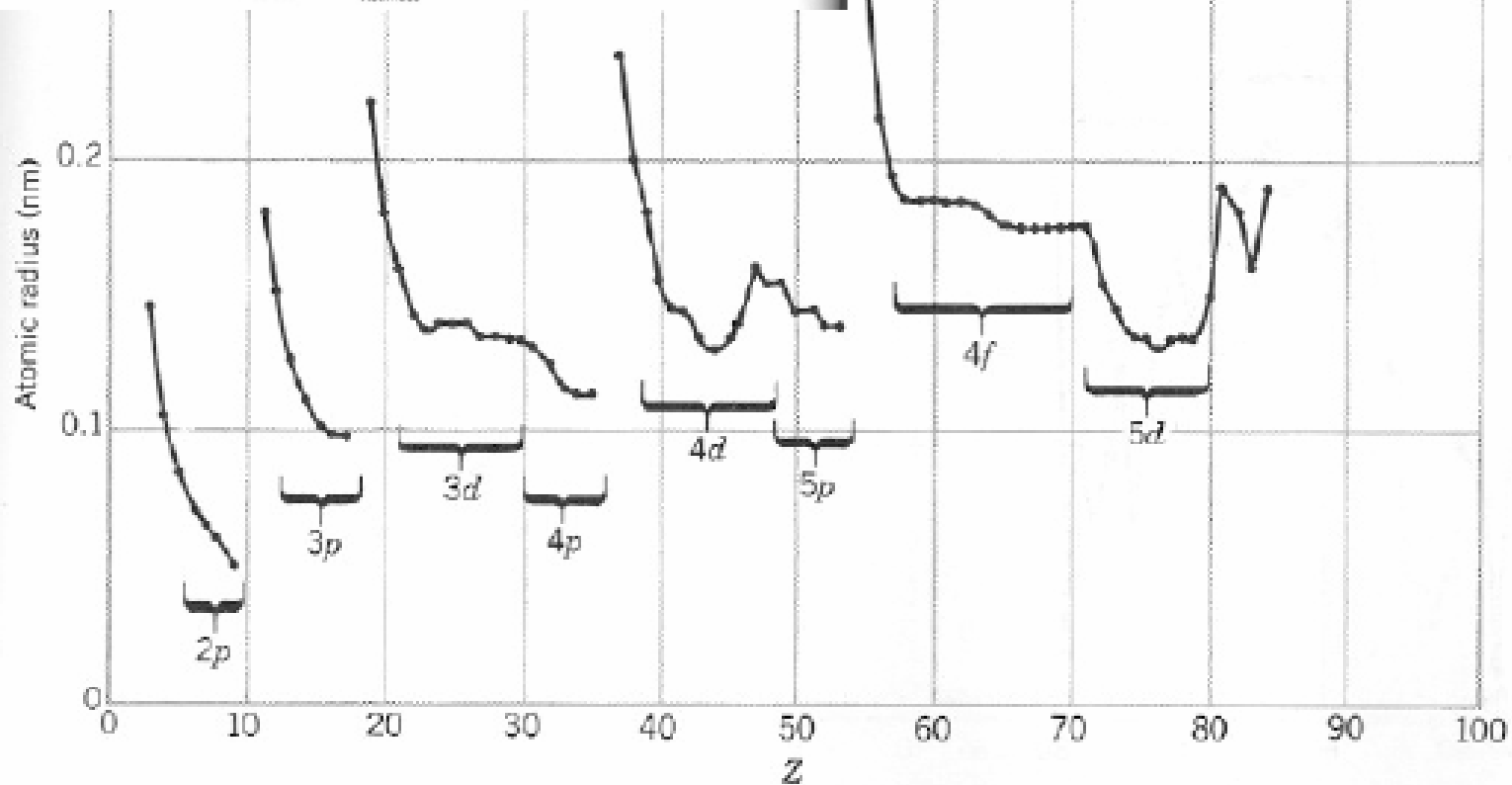
Alkalis		Transition metals										Halogens					Inert gases									
1s	1 H																2 He									
2s	3 Li 4 Be																									
3s	11 Na 12 Mg																									
4s	19 K 20 Ca	3d	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	3p	13 Al	14 Si	15 P	16 S	17 Cl	18 Ar							
5s	37 Rb 38 Sr	4d	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	4p	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr							
6s	55 Cs 56 Ba	5d	71 Lu	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	5p	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe							
7s	87 Fr 88 Ra	6d	103 Lr	104	105	106											6p	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn			
		Lanthanides (rare earths)																								
		4f	57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb										
		5f	89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Mv	102 No										
		Actinides																								

Alkalis	1 H	Alkaline earths	2 He	Inert gases
1s	3 Li	4 Be	5 B	6 C
2s	11 Na	12 Mg	7 N	8 O
3s	19 K	20 Ca	9 F	10 Ne
4s	37 Rb	38 Sr	13 Al	14 Si
5s	55 Cs	56 Ba	15 P	16 S
6s	87 Fr	88 Ra	17 Cl	18 Ar
7s			19 K	20 Ca
			31 Ga	32 Ge
			49 In	50 Sn
			81 Tl	82 Pb
			83 Bi	84 Po
			85 At	86 Rn

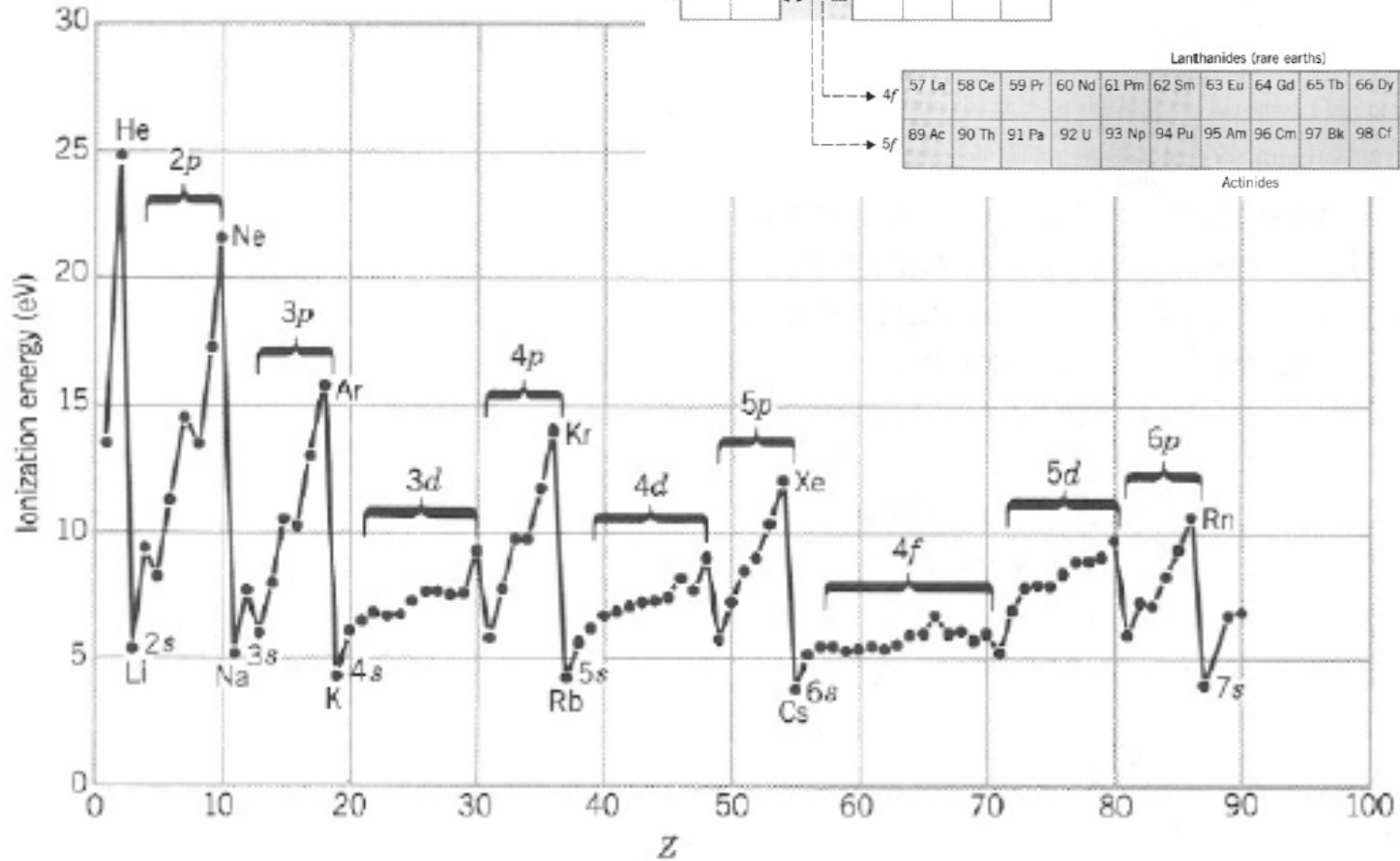
Transition metals										
3d	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn
4d	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd
5d	71 Lu	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg
6d	103 Lr	104	105	106						

Lanthanides (rare earths)														
4f	57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb
5f	89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Mv	102 No

Actinides														
	89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Mv	102 No

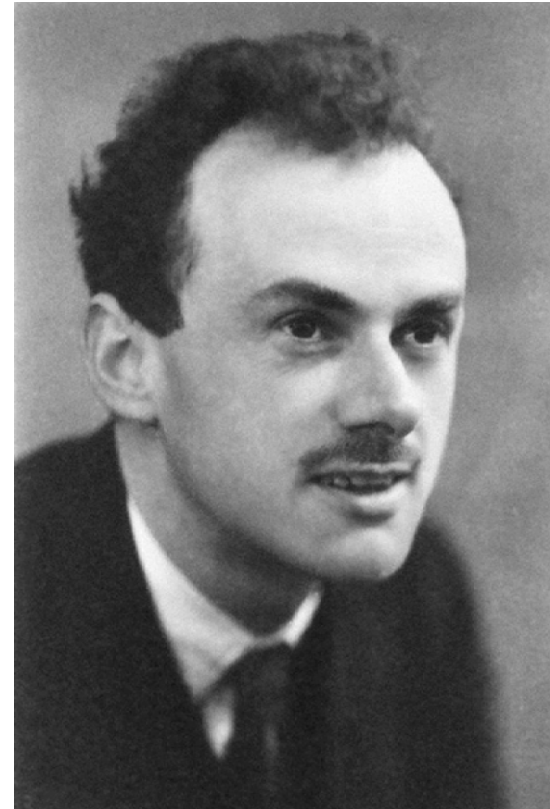


Alkalis										Inert gases														
1s	1 H		Alkaline earths												2 He									
2s	3 Li		4 Be												10 Ne									
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4s	19 K		20 Ca		21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr				
5s	37 Rb		38 Sr		39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe				
6s	55 Cs		56 Ba		71 Lu	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn				
7s	87 Fr		88 Ra		103 Lr	104	105	106																
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										Actinides														



The underlying physical laws necessary for the mathematical theory of a large part of physics and the whole of [chemistry](#) are thus completely known, and the difficulty is only that the exact application of these laws leads to equations much too complicated to be soluble. It therefore becomes desirable that approximate practical methods of applying [quantum mechanics](#) should be developed, which can lead to an explanation of the main features of complex atomic systems without too much computation.

- [*Proceedings of the Royal Society of London. Series A, Containing Papers of a Mathematical and Physical Character, Vol. 123, No. 792*](#) (6 April 1929)



Magnetic Resonance

Consider a current loop in a \vec{B} field



$$\vec{\tau} = \vec{\mu} \times \vec{B}$$

$$\text{PE of system} = -\vec{\mu} \cdot \vec{B}$$



in QM $\left\{ \begin{array}{l} \text{Spin} \\ \text{Spin component w.r. respect to an axis} \end{array} \right\}$ quantized

\rightarrow Could be orbital spin
or intrinsic spin

If we define \vec{B} to be along \hat{z}

$U \equiv$ energy of interaction of $\vec{\mu}$ w/ \vec{B}

$$U = -\mu_z B$$

$$\vec{\mu} = -\frac{1}{2} \frac{e}{m} \vec{L}$$

How magnetic moment
of e^- in atom
depends on \vec{L}
(orbital angular momentum)

For e^- in atom $l_z = m_l \hbar$

if $l = 1$ $m_l = -1, 0, +1$

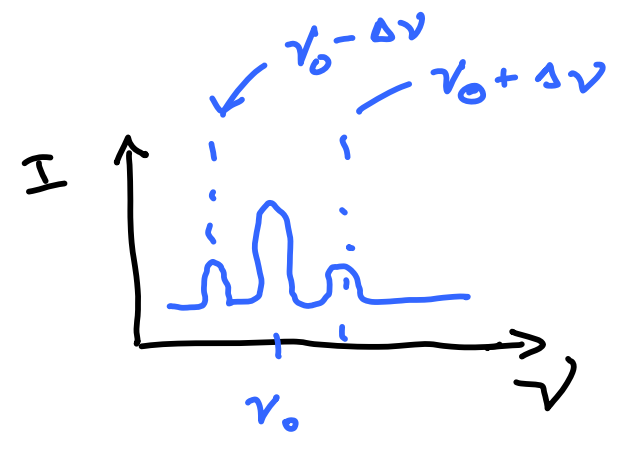
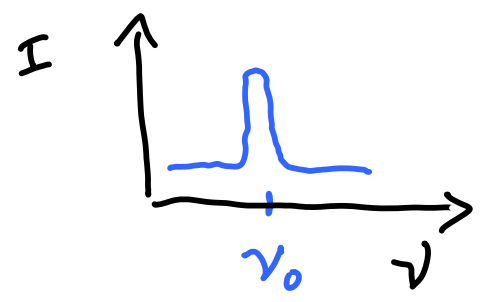
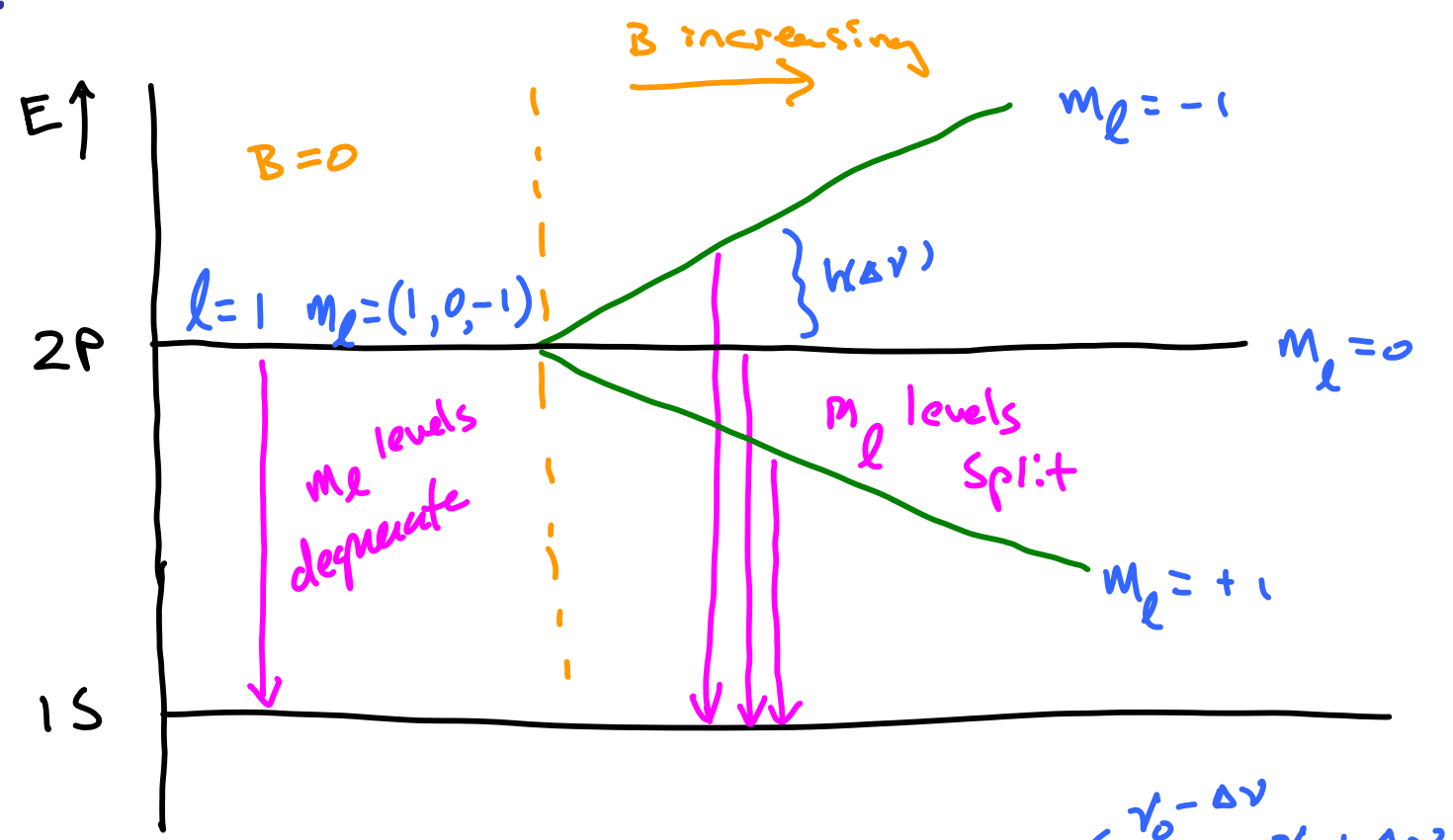
$$l_z = -\hbar, 0, +\hbar$$

$$\mu_z = \left(\frac{e\hbar}{2m} \right) m_l$$

Bohr magneton $\equiv \mu_B$

Zeeman effect

$$U = -\mu_B m_l B \quad \text{For } l=1 \quad m_l = -1, 0, +1$$



Electron Spin Resonance

$\vec{B}_{\text{external}}$

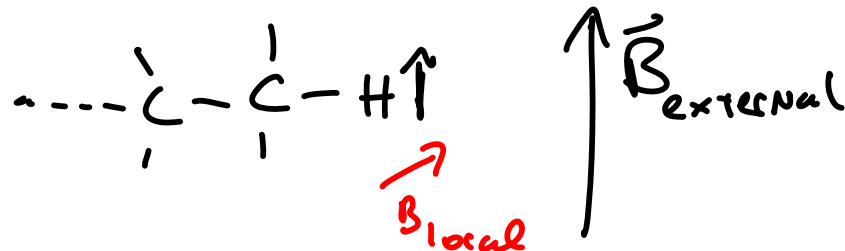
Chemical
crud

\downarrow \uparrow unpaired
 \vec{B}_{local} electron

Scan field

...

Nuclear Magnetic Resonance



or Scan frequency

Max Born German (1882-1970)

Note Title

2/26/2007

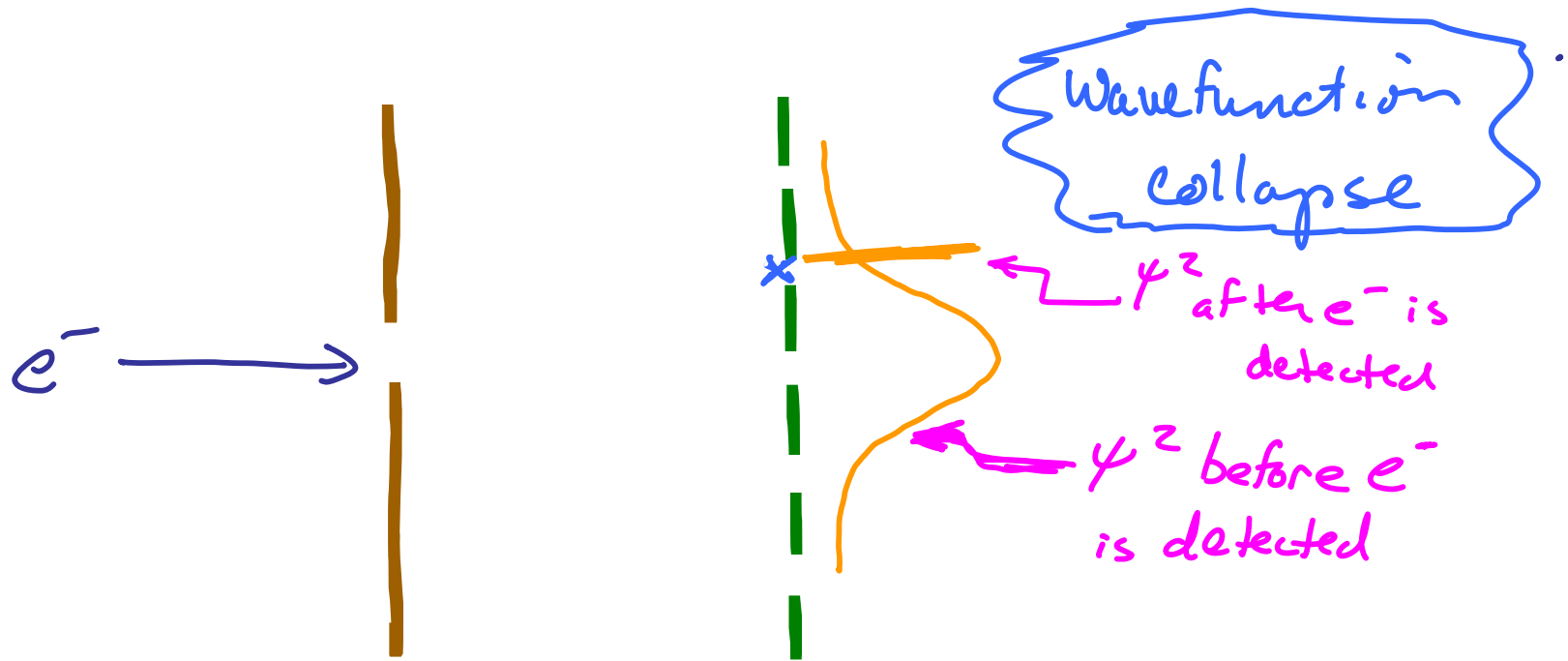


1954 Nobel Prize in physics

"For his fundamental research
in quantum mechanics,
especially for his statistical
interpretation of the
wavefunction"

$\psi(x)$ wave function

$\psi^2(x) \sim$ probability of finding particle
in region of space



Once electron hits the film/detector we know with 100% certainty where the electron hits
- So wavefunction has to "collapse"