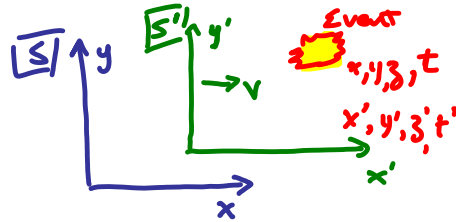


Physics 123 - January 30, 2013

Last Time

Lorentz Transformations



$S \rightarrow S'$

$$x' = \gamma(x - vt)$$
$$y' = y$$
$$z' = z$$
$$t' = \gamma\left(t - \frac{v}{c^2}x\right)$$

$S' \rightarrow S$

$$x = \gamma(x' + vt')$$
$$y = y'$$
$$z = z'$$
$$t = \gamma\left(t' + \frac{v}{c^2}x'\right)$$

you've seen it ...
Will see much more if you
keep taking physics

That said ... we won't use it ...

can be written as

$$X^\nu = L^{\mu\nu} X_\mu$$

4-vector index

Lorentz
matrix
index

repeated indices imply summation
"Einstein Summation notation"

Define spacetime 4-vector

$$x_0 = ct \quad x_1 = x \quad x_2 = y \quad x_3 = z$$

$$\begin{pmatrix} x'_0 \\ x'_1 \\ x'_2 \\ x'_3 \end{pmatrix} = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x_0 \\ x_1 \\ x_2 \\ x_3 \end{pmatrix}$$



$$\begin{cases} x'_0 = \gamma x_0 - \beta\gamma x_1 \\ x'_1 = -\beta\gamma x_0 + \gamma x_1 \\ x'_2 = x_2 \\ x'_3 = x_3 \end{cases}$$



Lorentz Transformation matrix

Dot product between 4-vectors is Lorentz invariant

Meaning What?

$$a^\mu b_\mu \quad \text{or} \quad a \cdot b = -a_0 b_0 + a_1 b_1 + a_2 b_2 + a_3 b_3$$

where a and b
are 4-vectors

Lorentz invariant

Suppose we have events a and b
described by 4-vectors a and b

$$\Delta x_\mu = a - b \quad \therefore \begin{pmatrix} a_0 - b_0 \\ a_1 - b_1 \\ a_2 - b_2 \\ a_3 - b_3 \end{pmatrix}$$

← displacement 4 vector

$$\Delta x \cdot \Delta x = -c^2 (\Delta t)^2 + \underbrace{\Delta x^2 + \Delta y^2 + \Delta z^2}_{d^2}$$

invariant interval

(spatial distance d^2 bet. evnts)²

Proper velocity 4-vector

$$\eta_0 = c \frac{dt}{d\tau} = c\gamma$$

$$\eta_1 = \frac{dx}{d\tau} = \gamma v_x$$

$$\eta_2 = \frac{dy}{d\tau} = \gamma v_y$$

$$\eta_3 = \frac{dz}{d\tau} = \gamma v_z$$

or u_x, u_y, u_z

Momentum + Energy conservation
hold across Lorentz transformations
if they are Defined
through the

Energy-Momentum 4-vector

$$P_0 = m\eta_0 = m\gamma c = \frac{m\gamma c^2}{c} = \frac{E}{c}$$

$$P_1 = m\eta_1$$

$$P_2 = m\eta_2$$

$$P_3 = m\eta_3$$

where

$$E = m\gamma c^2$$

= Relativistic
Energy

$$P^\mu P_\mu = P \cdot P = -\frac{E^2}{c^2} + \frac{m^2 u^2}{1 - \frac{v^2}{c^2}}$$

$$= -\frac{m^2 c^4}{c^2 (1 - \frac{v^2}{c^2})} + \frac{m^2 u^2}{(1 - \frac{v^2}{c^2})} = \frac{m^2 u^2 c^2 - m^2 c^4}{(1 - \frac{v^2}{c^2})} = \frac{m^2 c^2 (u^2 - c^2)}{c^2 - u^2}$$

$$= -m^2 c^2$$

$$\frac{(m \gamma c^2)^2}{c^2}$$

invariant mass

Lorentz invariant

$$P \cdot P = -m^2 c^2 = -\frac{E^2}{c^2} + m^2 u^2 \gamma^2$$

$$p^2$$

$$E^2 = m^2 c^4 + p^2 c^2$$

$$E = \gamma mc^2 = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$\frac{E}{c^2} = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{1}{c^2}$$

$$E = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Taylor Series

$$(1-x)^{-1/2} = 1 - \frac{1}{2}x + \frac{1 \cdot 3}{2 \cdot 4}x^2 - \frac{1 \cdot 3 \cdot 5}{2 \cdot 4 \cdot 6}x^3 + \dots$$

$$\left(1 - \frac{v^2}{c^2}\right) = 1 + \frac{1}{2} \frac{v^2}{c^2} + \frac{3}{8} \left(\frac{v^2}{c^2}\right)^2 - \dots$$

$$E = mc^2 + \frac{1}{2}mv^2 + \frac{3}{8}m \frac{v^4}{c^4} - \dots$$

rest
Energy

kinetic
Energy

higher order terms

Small unless v is close to c
we will ignore these

mass of Proton is $1.67 \times 10^{-27} \text{ kg} \sim 1 \text{ amu}$

Moves at $0.7c$

$$E = (1.4) m c^2 = 1.4 (1.67 \times 10^{-27} \text{ kg}) (3 \times 10^8 \text{ m/s})^2$$
$$= 2.1 \times 10^{-10} \text{ kg} \frac{\text{m}^2}{\text{s}^2} \rightsquigarrow \text{J}$$

Joules are not a convenient unit for Atomic + Subatomic particles

1 eV \equiv electron-Volt \equiv energy gained by one electron moving thru potential diff of 1 volt

$$\text{Energy} = m c^2$$

\uparrow
eV

unit of
Momentum $\sim \frac{\text{eV}}{c}$

unit of
Mass $\sim \frac{\text{eV}}{c^2}$

$$E^2 = p^2 c^2 + m^2 c^4$$

$$M_{\text{electron}} = 0.511 \text{ MeV}/c^2$$

$$1000 \text{ MeV} = \text{GeV}$$

$$M_{\text{proton}} = 938 \text{ MeV}/c^2$$

1000 MeV proton passing by ...

Total energy

$$M_p = 1.67 \times 10^{-27} \text{ kg} = 1.5 \times 10^{-10} \frac{\text{J}}{c^2} \frac{1 \text{ eV}}{1.6 \times 10^{-19} \text{ J}} = 938 \frac{\text{MeV}}{c^2}$$

$$E_{\text{TOT}} = E_0 + KE = 938 \frac{\text{MeV}}{c^2} c^2 + 1000 = 1938 \text{ MeV}$$

\downarrow
 mc^2

Momentum of that proton

$$E^2 = pc^2 + E_0^2$$

$$p = \frac{\sqrt{E^2 - E_0^2}}{c} = \frac{\sqrt{1938^2 - 938^2}}{c} = 1696 \frac{\text{MeV}}{c}$$

velocity in lab?

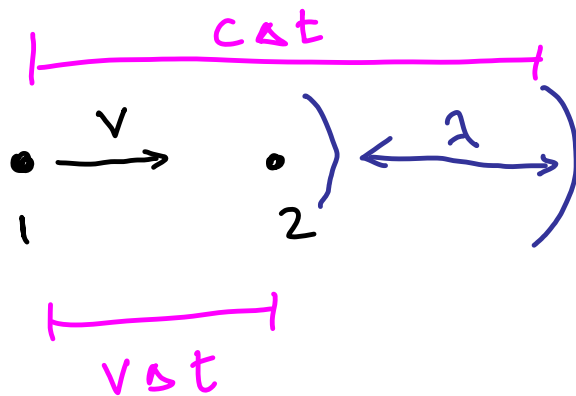
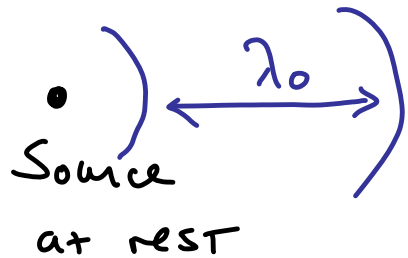
$$E = \gamma mc^2 =$$

$$mc^2 = \sqrt{1 - \left(\frac{v}{c}\right)^2} E$$

$$v = c \sqrt{1 - \left(\frac{E_0}{E}\right)^2}$$

$$v = c \sqrt{1 - \left(\frac{938}{1938}\right)^2} = 0.87 c$$

Doppler shift of light



Follows Giacomini's discussion

• Observer at rest

$$\lambda = c \Delta t - v \Delta t$$

• Observer

Δt is O frame

Δt_0 is S frame
 proper frame

$$\Delta t = \Delta t_0 \gamma$$

$\nu = \text{frequency}$

$$\lambda = (c - v) \delta \Delta t_0 = \frac{c - v}{\sqrt{1 - v^2/c^2}} \Delta t_0 = \frac{c - v}{\sqrt{c^2 - v^2}} \overbrace{\Delta t_0 c}^{\lambda_0}$$

$$\nu = \frac{c}{\lambda}$$

$$\lambda_{\text{observer}} = \lambda_0 \sqrt{\frac{c - v}{c + v}}$$

↑
Source

$$\nu = \nu_0 \sqrt{\frac{c + v}{c - v}}$$

Source
+
Observer
Moving toward
each other

$$\lambda_{\text{obs}} = \lambda_0 \sqrt{\frac{c + v}{c - v}}$$

$$\nu = \nu_0 \sqrt{\frac{c - v}{c + v}}$$

Moving
Away
from
each
other



Vesto Slipher (1875-1969) Lowell Observatory discovers a strange thing in 1912 ...

**Most nearby galaxies are moving away from us
Made use of the Doppler shift in atomic spectra**



Hubble

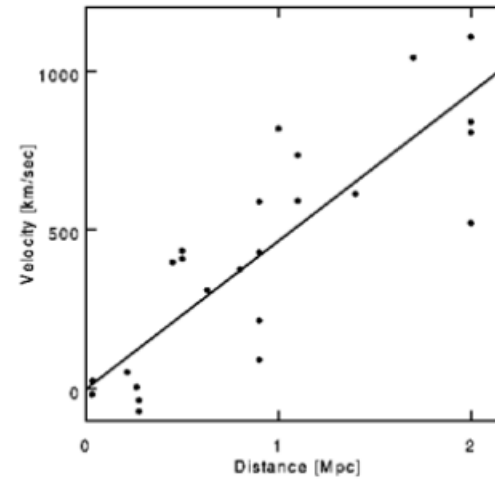


Humason

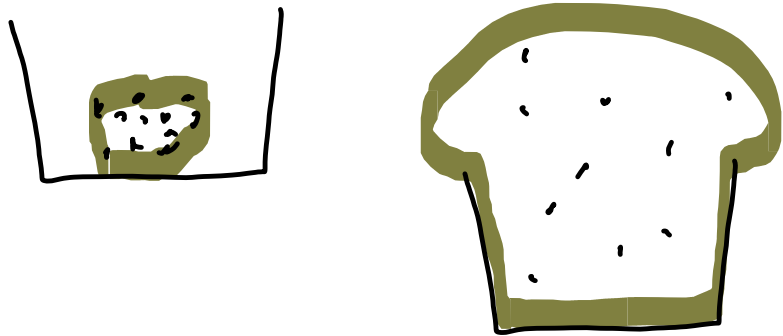
Edwin Hubble (1889-1953) and Milton Humason (1891-1972) at Mount Wilson Observatory combine Hubble's distance measurements (Cepheid variable stars) with Slipher's redshift information and discover ...

Galaxies that are further away are moving away from us faster

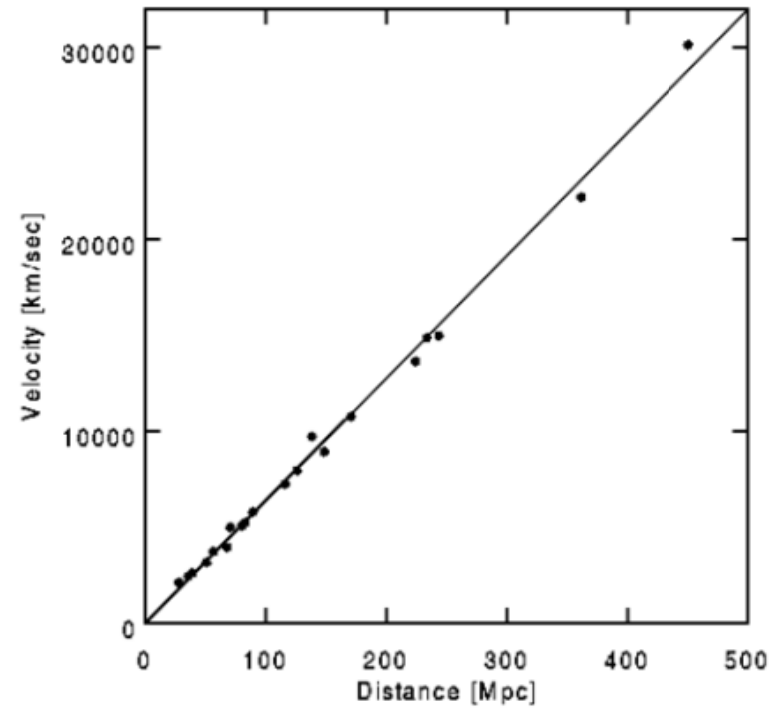
Hubble's Law $V=Hd$



Welcome to the
Expanding
Universe!



Type Ia SNe from Riess, Press and Kirshner (1996)



Light travels from NYC to San Francisco in 1/100 second and it
travels 1 Mpc in 3 million years