

where $\vec{m}(\omega) \equiv \frac{1}{2} \int d^3r' (\vec{r}' \times \vec{j}(\vec{r}', \omega))$ is magnetic dipole moment

$$\vec{Q}'_{ij}(\omega) = \int d^3r' 3 \vec{r}'_i \vec{r}'_j \rho(\vec{r}', \omega)$$

looks very close to electric quadrupole tensor

$$\vec{Q}_{ij} = \int d^3r' (3 \vec{r}'_i \vec{r}'_j - r'^2 \delta_{ij}) \rho(\vec{r}', \omega)$$

$$\vec{Q}'_{ij} = \vec{Q}_{ij} + \delta_{ij} \int d^3r' r'^2 \rho(\vec{r}', \omega)$$

$$\vec{I}_2 = -\hat{r} \times \vec{m}(\omega) - \frac{i\omega}{6} \hat{r} \cdot \vec{Q}(\omega) - \frac{i\omega}{6} \hat{r} \int d^3r' r'^2 \rho(\vec{r}', \omega)$$

call this $\mathcal{C}(\omega)$
a scalar

plug back into $\vec{A}(\vec{r}, \omega)$

$$\vec{A}(\vec{r}, \omega) = \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \left\{ \vec{I}_1 + \left(\frac{1}{r} - ik\right) \vec{I}_2 \right\}$$

$$= \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \left\{ \begin{array}{l} -i\omega \vec{p} \quad \uparrow \quad \text{electric dipole contribution} \\ -(\frac{1}{r} - ik) \left(\hat{r} \times \vec{m} \quad \uparrow \quad \text{magnetic dipole contribution} \right. \\ \left. + \frac{i\omega}{6} \hat{r} \cdot \vec{Q} \quad \uparrow \quad \text{electric quadrupole contribution} \right. \\ \left. + \frac{i\omega}{6} \hat{r} \mathcal{C} \right) \end{array} \right\}$$

The last piece which contributes to \vec{A} , i.e. $\frac{i\omega}{c} \hat{r} \frac{e^{-ikr}}{r}$ is unimportant - it does not affect the \vec{E} or \vec{B} fields since

$$\vec{B} = \vec{\nabla} \times \vec{A} \quad \text{and} \quad \vec{\nabla} \times [f(r) \hat{r}] = 0$$

similarly, away from sources, where $\vec{j} = 0$, Ampere's law gives

$$\mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t} = \vec{\nabla} \times \vec{B} = \vec{\nabla} \times (\vec{\nabla} \times \vec{A})$$

$$-i\omega \mu_0 \epsilon_0 \vec{E}(\vec{r}, \omega) = \vec{\nabla} \times (\vec{\nabla} \times \vec{A})$$

since last term doesn't contribute to \vec{B} , it doesn't contribute to \vec{E} . Formally, we could remove it by making a gauge transformation. Less formally, we will just drop it!

$$\vec{A}(\vec{r}, \omega) = \frac{\mu_0}{4\pi} \frac{e^{ikr}}{r} \left\{ -i\omega \vec{p} - \left(\frac{1}{r} - ik \right) (\hat{r} \times \vec{m} + \frac{i\omega}{c} \hat{r} \cdot \vec{Q}) \right\}$$

note: $\left(\frac{1}{r} - ik \right) = -\left(1 + \frac{i}{kr} \right) ik$

lets look at relative strengths of the different terms

far from sources, $\frac{1}{r}$ will be small compared to k , $\frac{1}{r} \ll k \Rightarrow \frac{1}{kr} \ll 1$
radiation zone: just consider those terms in \vec{A} that decrease as slowest powers of $\left(\frac{1}{r} \right)^n$. This will be the $\frac{1}{r}$ terms

Approx ① $d \ll r$

Approx ② $d \ll \lambda$

Radiation zone ③ $\lambda \ll r$ so $kr \gg 1$

Combine: $d \ll \lambda \ll r$ is RZ

electric dipole term $\vec{p} \approx qd$ q is typical charge in source
 d is size of source region

magnetic dipole term $\vec{m} = \frac{1}{2} \int d^3r \vec{r} \times \vec{j}$ $\vec{j} \sim qv$ where v is typical velocity
 $\approx d j \approx d v q$ $v \sim \frac{d\omega}{c} \sim d\omega$
 $\approx q d^2 \omega \sim q c d^2 k$ $\sim dck$

electric quadrupole term $\vec{Q} \sim \int d^3r \vec{r} \vec{r} \rho$
 $\sim q d^2$

so electric dipole contrib to \vec{A} goes as $\omega \vec{p} \sim q \omega d = qc(kd)$
 magnetic dipole contrib to \vec{A} goes as $k \vec{m} \sim q \omega k d^2 = qc(kd)^2$
 electric quadrupole contrib to \vec{A} goes as $k \omega \vec{Q} \sim q \omega k d^2 = qc(kd)^2$

Since approx (2) assumed (kd) was small
 (non relativistic approx: $kd \approx v/c$)

we have an expansion for \vec{A} in powers of (kd)

leading term is the electric dipole term.

next order terms are { magnetic dipole } \leftarrow these are comparable
 { electric quadrupole } in strength.

If we kept higher order terms in our expansion,
 the next terms would be the magnetic quadrupole
 and electric octopole, both of order $qc(kd)^3$.

Consider now the leading term, the electric dipole term

$$\vec{A}_{E1} = \frac{\mu_0}{4\pi} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} (-i\omega) \vec{p}(\omega) \quad \text{"E1" = electric dipole term}$$

magnetic field

$$\vec{B}_{E1}(\vec{r}, \omega) = \vec{\nabla} \times \vec{A}_{E1}(\vec{r}, \omega) = \frac{-i\omega\mu_0}{4\pi} \vec{\nabla} \times \left(\frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} \vec{p}(\omega) \right)$$

use $\vec{\nabla} \times (f\vec{g}) = (\vec{\nabla} f) \times \vec{g} + f \vec{\nabla} \times \vec{g}$ with $f = e^{i\mathbf{k}\cdot\mathbf{r}}$, $\vec{g} = \frac{\vec{p}(\omega)}{r}$

$$\vec{\nabla} e^{i\mathbf{k}\cdot\mathbf{r}} = \hat{r} \frac{\partial}{\partial r} (e^{i\mathbf{k}\cdot\mathbf{r}}) = e^{i\mathbf{k}\cdot\mathbf{r}} i\mathbf{k} \hat{r} \quad \text{in spherical coordinates}$$

$$\vec{\nabla} \times \left(\frac{\vec{p}}{r} \right) = \left(\vec{\nabla} \frac{1}{r} \right) \times \vec{p} + \frac{1}{r} \vec{\nabla} \times \vec{p} = -\frac{1}{r^2} \hat{r} \times \vec{p}$$

$\underbrace{\vec{\nabla} \times \vec{p}}_{=0 \text{ since } \vec{p} \text{ is constant}}$

$$\vec{B}_{E1}(\vec{r}, \omega) = \frac{-i\omega\mu_0}{4\pi} \left[e^{i\mathbf{k}\cdot\mathbf{r}} i\mathbf{k} \hat{r} \times \frac{\vec{p}(\omega)}{r} - e^{i\mathbf{k}\cdot\mathbf{r}} \frac{\hat{r} \times \vec{p}(\omega)}{r^2} \right]$$

use $\omega = ck$ $\vec{B}_{E1} = -\frac{c\mu_0 k^2}{4\pi} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} \left(1 + \frac{i}{kr} \right) \vec{p}(\omega) \times \hat{r}$ use $\hat{r} \times \vec{p} = -\vec{p} \times \hat{r}$

↑ small compared to 1 when $kr \gg 1$

We define the Radiation Zone limit when $r \gg \lambda \Rightarrow kr \gg 1$

far away on the scale of the wavelength of the radiated wave

In this limit the 2nd term is small compared to the first

$(1 + i/kr) \approx 1$ and we have

in R.Z. $\vec{B}_{E1}(\vec{r}, \omega) = -\frac{c\mu_0 k^2}{4\pi} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} \vec{p}(\omega) \times \hat{r}$

Electric field

from Ampere, $\mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t} = \vec{\nabla} \times \vec{B}$ since $\vec{j} = 0$
for from source

$$\Rightarrow \vec{E}_{EI} = \frac{i}{\omega \mu_0 \epsilon_0} \vec{\nabla} \times \vec{B}_{EI} \quad \text{since } \frac{\partial \vec{E}}{\partial t} = -i\omega \vec{E}$$

$$= \frac{-i}{\omega \mu_0 \epsilon_0} \frac{c \mu_0}{4\pi} k^2 \vec{\nabla} \times \left(\frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r} \right)$$

$$\omega = ck$$

to evaluate the $\vec{\nabla} \times (\)$ term, we $\vec{\nabla} \times (f\vec{g}) = f\vec{\nabla} \times \vec{g} + \vec{\nabla} f \times \vec{g}$
with $f = e^{ikr}$, $\vec{g} = \frac{1}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r}$

$$\text{then } \vec{\nabla} \times (\) = \left(\vec{\nabla} e^{ikr} \right) \times \left(\frac{1}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r} \right)$$

$$+ e^{ikr} \vec{\nabla} \times \left(\frac{1}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r} \right)$$

But in the radiation zone we can ignore the second term, since

$$\vec{\nabla} \times \left(\frac{1}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r} \right) \sim \frac{1}{r^2}$$

we see this by noting that

$$\vec{\nabla} \frac{1}{r} \sim \frac{1}{r^2}, \quad \vec{\nabla} \frac{1}{r^2} \sim \frac{1}{r^3}$$

$$\vec{\nabla} \hat{r} \sim \frac{\partial \hat{r}}{\partial x} \frac{1}{|\hat{r}|} = \frac{\hat{x}}{|\hat{r}|} - \frac{\hat{r}}{|\hat{r}|^2} \frac{x}{|\hat{r}|} = \frac{\hat{x}}{|\hat{r}|} - \frac{x\hat{r}}{|\hat{r}|^2} \sim 0 \left(\frac{1}{r}\right)$$

So keep only 1st term in Radiation Zone and we get

$$\vec{E}_{EI} = \frac{-i}{\omega \mu_0 \epsilon_0} \frac{c \mu_0}{4\pi} k^2 \left(i k e^{ikr} \hat{r} \right) \times \left(\frac{1}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r} \right)$$

use $\omega = ck$

$$= \vec{\nabla} e^{ikr}$$

\vec{E} ignore in RZ

$$\vec{E}_{EI} = \frac{k^2}{4\pi \epsilon_0} \frac{e^{ikr}}{r} \hat{r} \times (\vec{p} \times \hat{r})$$

electric field - full calculation without making Radiation Zone approximation

from Ampere, $\mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t} = \nabla \times \vec{B}$, since $\vec{j} = 0$ far from source

$$\Rightarrow \vec{E}_{E1} = \frac{i}{\omega \mu_0 \epsilon_0} \nabla \times \vec{B}_{E1} \quad \text{since } \frac{\partial \vec{E}}{\partial t} = -i\omega \vec{E}$$

$$= \frac{-i}{\omega \mu_0 \epsilon_0} \frac{c \mu_0}{4\pi} k^2 \nabla \times \left(\frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right) \vec{p} \times \hat{r} \right)$$

$\omega = ck$

to evaluate $\nabla \times ()$, use $\nabla \times (f\vec{g}) = f \nabla \times \vec{g} + \nabla f \times \vec{g}$
with $f = \frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right)$ and $\vec{g} = \vec{p} \times \hat{r}$

$$\nabla \times () = \frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right) \nabla \times (\vec{p} \times \hat{r}) + \nabla \left(\frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right) \right) \times (\vec{p} \times \hat{r})$$

evaluate second term: $\nabla \left(\frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right) \right) = \frac{\partial}{\partial r} \left(\frac{e^{ikr}}{r} \left(1 + \frac{i}{kr}\right) \right) \hat{r}$ in spherical coords

$$= e^{ikr} \left[ik \left(\frac{1}{r} + \frac{i}{kr^2} \right) - \frac{1}{r^2} - \frac{2i}{kr^3} \right] \hat{r}$$

$$= \frac{e^{ikr}}{r} \left[ik - \frac{2}{r} - \frac{2i}{kr^2} \right] \hat{r}$$

evaluate first term:

$$\nabla \times (\vec{p} \times \hat{r}) = \vec{p} (\nabla \cdot \hat{r}) - (\vec{p} \cdot \nabla) \hat{r}$$

$$\text{where } \nabla \cdot \hat{r} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2) = \frac{2}{r}$$

cover Griffiths evaluating in spherical coordinates

$$\text{and } (\vec{p} \cdot \nabla) \hat{r} = \sum_k p_k \frac{\partial \hat{r}}{\partial r_k}$$

unit vector in k direction

$$\text{where } \frac{\partial \hat{r}}{\partial r_k} = \frac{\partial}{\partial r_k} \left(\frac{\vec{r}}{r} \right) = \vec{r} \left(-\frac{1}{r^2} \frac{\partial r}{\partial r_k} \right) + \frac{\hat{e}_k}{r}$$

$$= \vec{r} \left(-\frac{1}{r^2} \frac{r_k}{r} \right) + \frac{\hat{e}_k}{r} \quad \text{as } \frac{\partial r}{\partial r_k} = \frac{r_k}{r}$$

$$\begin{aligned}
 \text{so } \vec{\nabla} \times (\vec{p} \times \hat{r}) &= \frac{2\vec{p}}{r} - \sum_k p_k \left(-\frac{\hat{r}}{r^3} r_k + \frac{\hat{e}_k}{r} \right) \\
 &= \frac{2\vec{p}}{r} + \frac{\hat{r}}{r^3} \vec{p} \cdot \hat{r} - \frac{\vec{p}}{r} \\
 &= \frac{\vec{p} + \hat{r}(\vec{p} \cdot \hat{r})}{r} \quad \text{using } \hat{r} = \frac{\vec{r}}{r}
 \end{aligned}$$

putting all the pieces together

$$\begin{aligned}
 \vec{E}_{E1} &= \frac{-ik}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \left[(1 + \frac{i}{kr}) \frac{\vec{p} + \hat{r}(\vec{p} \cdot \hat{r})}{r} \right. \\
 &\quad \left. + \left(ik - \frac{2}{r} - \frac{2i}{kr^2} \right) \frac{\hat{r} \times (\vec{p} \times \hat{r})}{\vec{p} - \hat{r}(\vec{p} \cdot \hat{r})} \right]
 \end{aligned}$$

order by powers of $\frac{1}{r}$

$$\begin{aligned}
 &= \frac{-ik}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \left[ik(\vec{p} - \hat{r}(\vec{p} \cdot \hat{r})) + \frac{1}{r} \left(1 + \frac{i}{kr} \right) (\vec{p} + \hat{r}(\vec{p} \cdot \hat{r})) \right. \\
 &\quad \left. - \frac{2}{r} \left(1 + \frac{i}{kr} \right) (\vec{p} - \hat{r}(\vec{p} \cdot \hat{r})) \right] \\
 &= \frac{k^2}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \left[\vec{p} - \hat{r}(\vec{p} \cdot \hat{r}) - \frac{i}{kr} \left(1 + \frac{i}{kr} \right) (\vec{p} + \hat{r}(\vec{p} \cdot \hat{r})) \right. \\
 &\quad \left. - 2\vec{p} + 2\hat{r}(\vec{p} \cdot \hat{r}) \right]
 \end{aligned}$$

$$\boxed{\vec{E}_{E1} = \frac{k^2}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \left[\vec{p} - \hat{r}(\vec{p} \cdot \hat{r}) - \frac{i}{kr} \left(1 + \frac{i}{kr} \right) (3\hat{r}(\vec{p} \cdot \hat{r}) - \vec{p}) \right]}$$

radiation zone approx $kr \gg 1$ keep only terms of order $\frac{1}{r}$

$$\boxed{\vec{E}_{E1} = \frac{k^2}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \left[\underbrace{\vec{p} - \hat{r}(\vec{p} \cdot \hat{r})}_{\hat{r} \times (\vec{p} \times \hat{r})} \right]}$$

Radiation zone limit

$$\vec{E}_{EI} = \frac{k^2}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \hat{r} \times (\vec{p} \times \hat{r})$$

$$\vec{B}_{EI} = -\frac{c\mu_0}{4\pi} k^2 \frac{e^{ikr}}{r} \vec{p} \times \hat{r}$$

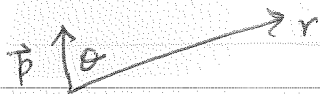
radiation zone fields
in electric dipole approx

\vec{E} and \vec{B} are outwards traveling spherical waves
 $\sim \frac{e^{ikr}}{r}$

$$\frac{|\vec{B}_{EI}|}{|\vec{E}_{EI}|} = \frac{c\mu_0}{4\pi} \cdot 4\pi\epsilon_0 = c\mu_0\epsilon_0 \stackrel{\text{using } \mu_0\epsilon_0 = 1/c^2}{=} \frac{c}{c^2} = \frac{1}{c}$$

just as for plane waves in vacuum

if choose coordinates so that \vec{p} is along \hat{z} axis



$$\vec{p} \times \hat{r} = p \sin\theta \hat{\phi}$$

$$\hat{r} \times (\vec{p} \times \hat{r}) = p \sin\theta (\hat{r} \times \hat{\phi})$$

$$= -\hat{\theta} \sin\theta p$$

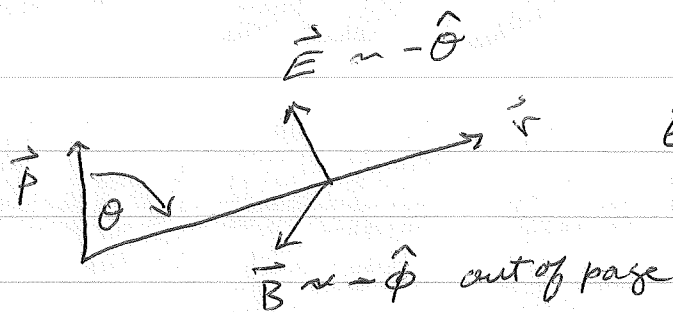
$$\vec{E}_{EI} = -\frac{k^2 p}{4\pi\epsilon_0} \frac{e^{ikr}}{r} \sin\theta \hat{\theta}$$

$$\vec{B}_{EI} = -\frac{c\mu_0}{4\pi} k^2 p \frac{e^{ikr}}{r} \sin\theta \hat{\phi}$$

$\hat{\theta}$ + $\hat{\phi}$ are
spherical coord
basis vectors

Note: Above assumes that \vec{p} is a real
valued vector. In general it is possible

\vec{p} may be complex, $\vec{p} = \vec{p}_1 + i\vec{p}_2$, and
that \vec{p}_1 and \vec{p}_2 may point in different directions!



\vec{E} is in the plane containing \vec{p} and \hat{r}
 \vec{B} is \perp to this plane

\vec{E}_{E1} and \vec{B}_{E1} are orthogonal, as in a plane wave,
 and both are orthogonal to the direction of propagation \hat{r}

\Rightarrow oscillating source emits spherical electromagnetic waves

What is Power emitted?

Poynting vector: $\vec{S}_{E1}(\vec{r}, t) = \frac{1}{\mu_0} \text{Re}[\vec{E}_{E1}(\vec{r}, t)] \times \text{Re}[\dot{\vec{B}}_{E1}(\vec{r}, t)]$

$$\text{Re}[\vec{E}_{E1}(\vec{r}, t)] = \text{Re}\left[-\frac{k^2 p}{4\pi\epsilon_0} \frac{e^{i(kr - \omega t)}}{r} \sin\theta \hat{\theta} e^{-i\omega t}\right]$$

$$= \frac{-k^2 p}{4\pi\epsilon_0} \frac{\cos(kr - \omega t)}{r} \sin\theta \hat{\theta}$$

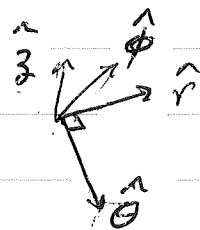
$$\text{Re}[\dot{\vec{B}}_{E1}(\vec{r}, t)] = \text{Re}\left[-\frac{c\mu_0}{4\pi} k^2 p \frac{e^{i(kr - \omega t)}}{r} \sin\theta \hat{\phi} e^{-i\omega t}\right]$$

$$= -\frac{c\mu_0 k^2 p}{4\pi} \frac{\cos(kr - \omega t)}{r} \sin\theta \hat{\phi}$$

 Assuming
 \vec{p} is a real
 valued vector

$$\vec{S}_{E1} = \frac{1}{\mu_0} \frac{k^2 p}{4\pi\epsilon_0} \frac{c\mu_0 k^2 p}{4\pi} \frac{\cos^2(kr - \omega t)}{r^2} \sin^2\theta (\hat{\theta} \times \hat{\phi})$$

$$= \frac{c k^4 p^2}{(4\pi)^2 \epsilon_0} \frac{\cos^2(kr - \omega t)}{r^2} \sin^2\theta \hat{r}$$



Average over one period of oscillation $\langle \cos^2(kr - \omega t) \rangle = \frac{1}{2}$

$$\langle \vec{S}_{EI} \rangle = \frac{ck^4 p^2}{2(4\pi)^2 \epsilon_0} \frac{\sin^2 \theta}{r^2} \hat{r}$$

Note, the $\frac{1}{r^2}$ is important for energy conservation.

If we integrate $\langle \vec{S}_{EI} \rangle \cdot \hat{r}$ over the surface of a sphere of radius r , the result is independent of r .

average energy flux flowing through an element of area at spherical angles θ, ϕ is

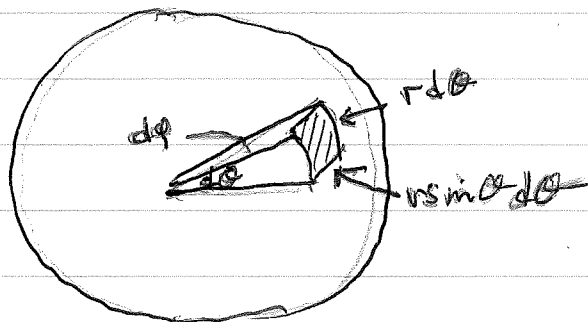
$$\text{power } dP_{EI} = \hat{r} \cdot \langle \vec{S}_{EI} \rangle \underbrace{r^2 \sin \theta d\theta d\phi}_{\text{differential area on surface of sphere spanned by } d\theta \text{ and } d\phi = r^2 d\Omega}$$

differential area on surface of sphere spanned by $d\theta$ and $d\phi$
 $= r^2 d\Omega$

$d\Omega = \sin \theta d\theta d\phi$ differential solid angle

$$dP_{EI} = \hat{r} \cdot \langle \vec{S}_{EI} \rangle \cdot r^2 d\Omega$$

$$\frac{dP_{EI}}{d\Omega} = \hat{r} \cdot \langle \vec{S}_{EI} \rangle r^2 = \frac{ck^4 p^2}{2(4\pi)^2 \epsilon_0} \sin^2 \theta$$



$$da = r^2 \sin \theta d\theta d\phi = r^2 d\Omega$$