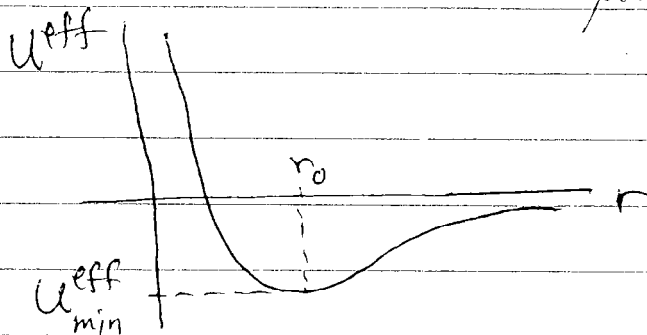


# Stability of circular orbits

$$U_{\text{eff}}(r) = U(r) + \frac{l^2}{2\mu r^2}$$

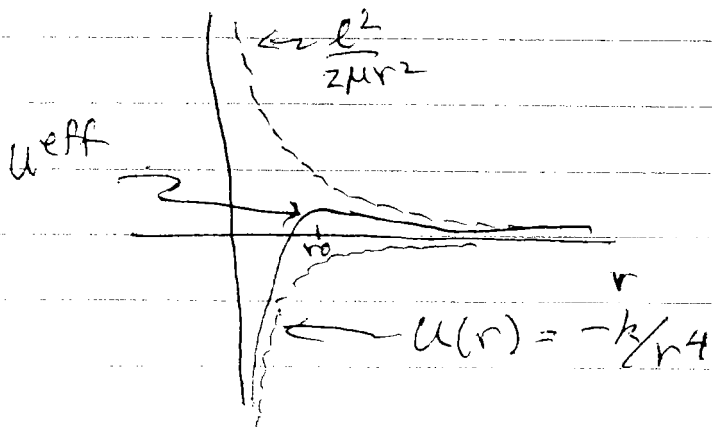


← for example when  $U(r) = -\frac{k}{r}$

There will be a stable ~~one~~ circular orbit at  $r = r_0$  if  $\left. \frac{dU^{\text{eff}}}{dr} \right|_{r=r_0} = 0$  (ie  $U^{\text{eff}}$  is extremal at  $r_0$ )

and  $\left. \frac{d^2U^{\text{eff}}}{dr^2} \right|_{r=r_0} > 0$  (ie  $r_0$  is a minimum)

We could have unstable orbits



If  $|U(r)|$  decreases faster than  $\frac{l^2}{2\mu r^2}$

we get a maximum in  $U^{\text{eff}}(r)$

Here  $\left. \frac{dU^{\text{eff}}}{dr} \right|_{r=r_0} = 0$

but  $\left. \frac{d^2U^{\text{eff}}}{dr^2} \right|_{r=r_0} < 0$  unstable

The two conditions for stable orbit at  $r_0$  are therefore

$$\textcircled{1} \quad \left. \frac{dU^{\text{eff}}}{dr} \right|_{r_0} = \left. \frac{dU}{dr} \right|_{r_0} - \frac{l^2}{\mu r_0^3} = 0$$

$$\text{since } F = -\frac{dU}{dr} \Rightarrow F = \frac{-l^2}{\mu r_0^3}$$

$$\textcircled{2} \quad \left. \frac{d^2U^{\text{eff}}}{dr^2} \right|_{r_0} = \left. \frac{d^2U}{dr^2} \right|_{r_0} + \frac{3l^2}{\mu r_0^4} > 0$$

$$\text{or } -\frac{dF}{dr} + \frac{3l^2}{\mu r_0^4} > 0$$

For a power law force

$$F = -\frac{k}{r^n}$$

$$\textcircled{1} \Rightarrow -\frac{k}{r_0^n} = \frac{-l^2}{\mu r_0^3} \Rightarrow r_0^{n-3} = \frac{\mu k}{l^2}$$

$$\textcircled{2} \Rightarrow -\frac{nk}{r_0^{n+1}} + \frac{3l^2}{\mu r_0^4} > 0$$

$$\Rightarrow -\frac{nk}{r_0^{n-3}} + \frac{3l^2}{\mu} > 0$$

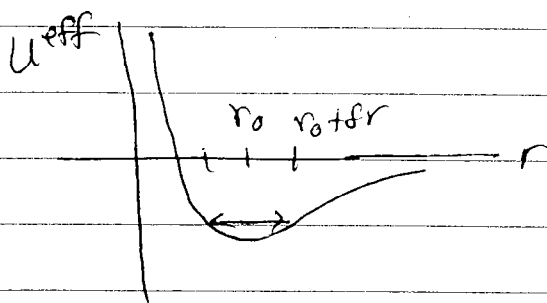
$$\Rightarrow \frac{-nk}{(\mu k/l^2)} + \frac{3l^2}{\mu} = \frac{l^2}{\mu} (3-n) > 0$$

$$\Rightarrow (3-n) > 0 \Rightarrow \boxed{n < 3} \quad \text{for stable circular orbit}$$

$n=2$  for gravity

prob 8-35

Suppose we have a stable periodic orbit,  
and the orbit radius is perturbed to  $r_0 + \delta r$



The particle will oscillate  
between radii  $r_0 + \delta r$   
and  $r_0 - \delta r$  with frequency  
of oscillation

$$\omega_0^2 = \left[ \left( \frac{d^2 U^{\text{eff}}}{dr^2} \right)_{r=r_0} \frac{1}{\mu} \right] = \frac{\text{curvature}}{\text{mass}}$$

From previous page, for  $F = -\frac{k}{r^n}$

$$\left. \frac{d^2 U^{\text{eff}}}{dr^2} \right|_{r_0} = \frac{l^2}{\mu r_0^4} (3-n)$$

$$\omega_0 = \sqrt{\frac{l^2}{\mu^2 r_0^4} (3-n)} = \frac{l}{\mu r_0^2} \sqrt{3-n}$$

$$\text{since } l = \mu r_0^2 \dot{\theta}$$

$$\omega_0 = \dot{\theta} \sqrt{3-n}$$

for gravitational force,  $n=2$

$$\Rightarrow \omega_0 = \dot{\theta}$$

period of oscillation between  $r_{\text{min}}$  and  
 $r_{\text{max}}$  is just period of revolution  
of elliptical orbit, as expected.

More generally,

$$\omega_0 = \dot{\theta} \sqrt{3-n}$$

$$\Rightarrow \text{period of oscillation } T = \frac{2\pi}{\omega_0} = \frac{2\pi}{\dot{\theta} \sqrt{3-n}}$$

$$\Rightarrow \frac{d\theta}{dt} = \frac{2\pi}{T\sqrt{3-n}} \Rightarrow \int_{\theta_1}^{\theta_2} d\theta = \int_0^T dt \frac{2\pi}{T\sqrt{3-n}}$$

$\Rightarrow$  change in angle of rotation  $\Delta\theta$ , after one period  $T$ ,

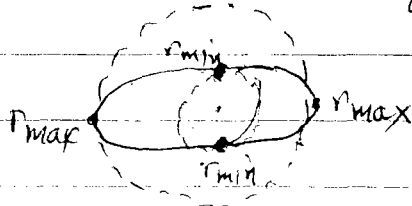
$$\Delta\theta = \frac{2\pi}{\sqrt{3-n}}$$

for  $n=2$ ,  $\Delta\theta = 2\pi$ , one complete revolution of origin in each period of oscillation of  $r$ .

$n=1$ ,  $\Delta\theta = \frac{2\pi}{\sqrt{2}}$ , less than one complete revolution of origin in time  $T$ .  $\frac{\Delta\theta}{2\pi}$  is irrational  $\Rightarrow$  orbit is not closed

$n=-1$ ,  $\Delta\theta = \frac{2\pi}{2} = \pi$ , one half revolution each period orbit closes after two periods

$F = -kr$   
harmonic force



example: screened Coulomb potential

$$U(r) = -\frac{k}{r} e^{-r/a} \quad \sim -\frac{1}{r} \text{ for } r \ll a$$
$$\sim e^{-r/a} \text{ for } r \gg a$$

expect ~~effect~~ that, depending on values of  $a$  and  $l$ , we might have stable or unstable circular orbit  
expect stable orbit when  $r_0 \lesssim a$ , unstable when  $r_0 \gtrsim a$

$$U^{\text{eff}} = U(r) + \frac{l^2}{2\mu r^2} = -\frac{k}{r} e^{-r/a} + \frac{l^2}{2\mu r^2}$$

$$\textcircled{1} \quad \frac{dU^{\text{eff}}}{dr} = \left( \frac{k}{r^2} + \frac{k}{ar} \right) e^{-r/a} - \frac{l^2}{\mu r^3} = 0$$

$$\textcircled{2} \quad \frac{d^2 U^{\text{eff}}}{dr^2} = \left( -\frac{2k}{r^3} - \frac{k}{ar^2} - \frac{k}{a^2 r} - \frac{k}{a^2 r} \right) e^{-r/a} + \frac{3l^2}{\mu r^4} > 0$$
$$= -k \left( \frac{2}{r^3} + \frac{2}{ar^2} + \frac{1}{a^2 r} \right) e^{-r/a} + \frac{3l^2}{\mu r^4} > 0$$

$$\text{from } \textcircled{1}: \left( kr + \frac{kr^2}{a} \right) e^{-r/a} = \frac{l^2}{\mu}$$

Solve for  $r$  to get radius of orbit as function of angular momentum  $l$ . - no analytical solution in general - need to do numerically.

$$\text{But still we can write } e^{-r/a} = \frac{l^2}{\mu k} \frac{1}{(r + r^2/a)}$$

substitute into (2)

$$-k \left( \frac{2}{r^3} + \frac{2}{ar^2} + \frac{1}{a^2 r} \right) \frac{l^2}{\mu k} \frac{1}{(r + r^2/a)} + \frac{3l^2}{\mu r^4} > 0$$

multiply by  $\frac{\mu r^4}{l^2}$

$$\Rightarrow 3 - \frac{(2r + 2\frac{r^2}{a} + \frac{r^3}{a^2})}{r + r^2/a} > 0$$

$$\Rightarrow 3 - \frac{2 + 2(r/a) + (r/a)^2}{1 + (r/a)} > 0$$

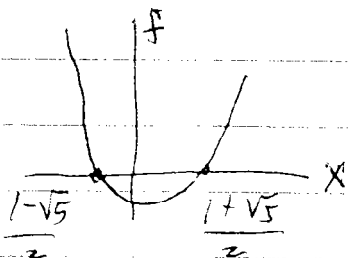
$$\Rightarrow 3 + \frac{3r}{a} - 2 - 2\left(\frac{r}{a}\right) - \left(\frac{r}{a}\right)^2 > 0$$

$$\Rightarrow 1 + \frac{r}{a} - \left(\frac{r}{a}\right)^2 > 0$$

$$\text{or } \left(\frac{r}{a}\right)^2 - \frac{r}{a} - 1 < 0$$

$$\text{define } x = \frac{r}{a}, f(x) = x^2 - x - 1 < 0$$

$$f(x) \text{ has zeros at } x_0 = \frac{1}{2} \pm \sqrt{\frac{1}{4} + 1} = \frac{1}{2} \pm \sqrt{\frac{5}{4}} \\ = \frac{1 \pm \sqrt{5}}{2}$$



So  $f(x)$  is negative between the two roots. Since  $x = r/a$  must be positive, we have physical

solution for all  $0 < \frac{r}{a} < \frac{1 + \sqrt{5}}{2}$

or  $r/a < 1.618$  ← for all  $r$  in this range, there is an  $l$  that gives stable orbit.

If we compare the screened Coulomb potential

$$U(r) = -\frac{k}{r} e^{-r/a}$$

to the ordinary Coulomb (or gravitational) potential

$$U(r) = -\frac{k}{r}$$

We found that for  $U^e$ , there is always a stable circular orbit for any angular momentum  $l$ . This orbit is at radius

$$r_0 = \frac{l^2}{\mu k}$$

as  $l$  increases, the radius  $r_0$  increases, and can in principle become arbitrarily large - one never "escapes" the Coulomb force - there can be bound states at arbitrarily large radius.

For  $U^{sc}$  however, there are stable circular orbits only out to a distance  $r_{\max} = 1.618 a$ , set by the "screening length"  $a$ . No stable circular orbits exist for radius  $r > r_{\max}$ . The relation between the radius  $r_0$  of the stable circular orbit and the angular momentum  $l$ , is given by the solution to 
$$kr \left(1 + \frac{r}{a}\right) e^{-r/a} = \frac{l^2}{\mu}$$

Although we have no analytic solution, it is clear that as  $l$  increases,  $r_0$  increases. Therefore there is a maximum angular momentum  $l_{\max}$  for a stable circular orbit.

When one treats such problems quantum mechanically, one finds that angular momentum is quantized into discrete values

$$l^2 = \hbar^2 \ell(\ell+1) \text{ for integer } \ell.$$

The conclusion for  $U^C$ , that there is a stable circular orbit for any value of  $l$  no matter how big, translates into the quantum mechanical conclusion that there are ~~an infinite~~ a countably infinite number of bound states in a Coulomb potential.  
- i.e. for the hydrogen atom, the principle quantum number  $n$  can be arbitrarily large.

The conclusion for  $U^{\text{sc}}$  that there are stable circular orbits only out to a particular ~~max~~ maximum value  $l_{\text{max}}$  becomes the quantum mechanical conclusion that the screened Coulomb potential has only a finite number of bound states. Depending on the value of  $\mu ka$ , there might be, for example, only one, or even none, bound states. The larger  $\mu ka$ , the more bound states.

$$\text{Since } \left(\frac{r}{a}\right) \left(1 + \left(\frac{r}{a}\right)\right) e^{-r/a} = \frac{l^2}{\mu ka}$$

$\Rightarrow \mu ka$  sets the scale of the max value of  $l$ .



Consider the equation that ~~is~~ determines the extremal values of  $U^{\text{eff}}$

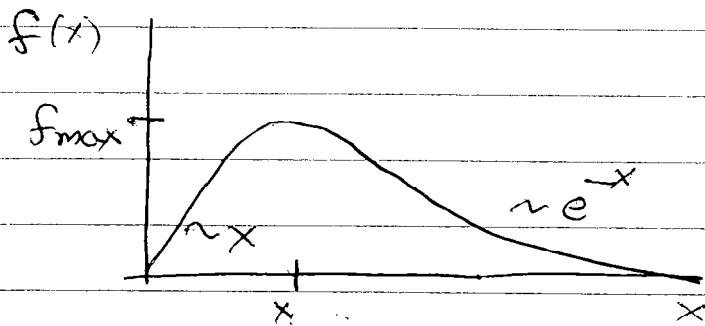
$$\left(\frac{r}{a}\right)\left(1 + \frac{r}{a}\right)e^{-r/a} = \frac{l^2}{\mu ka}$$

let  $x = \frac{r}{a}$ ,  $y = \frac{l^2}{\mu ka}$

$$f(x) = x(1+x)e^{-x}$$

For given value of  $y$ , there are extrema of  $U^{\text{eff}}$  at  $x$  such that

$$f(x) = y$$



to determine  $x_0$

$$\frac{df}{dx} = 0$$

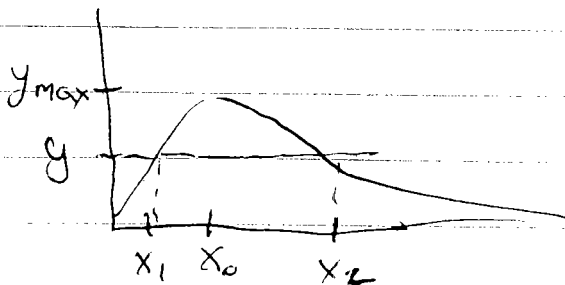
$$\Rightarrow (1+x+x-x(1+x))e^{-x} = 0$$

$$\Rightarrow 1+x-x^2 = 0$$

$$\Rightarrow x_0 = \frac{1+\sqrt{5}}{2} = 1.618$$

let  $y_{\text{max}} \equiv f(x_0)$

For  $y < y_{\text{max}}$ , there are two solutions to  $f(x) = y$ .



only  $x_1$  is stable since

$$x_2 = \frac{r_2}{a} > 1.618 = x_0$$

as  $y$  increases above  $y_{\max}$  there are no solutions

$$y_{\max} = \frac{l_{\max}^2}{\mu ka} = f(1.618) \quad \text{determines } l_{\max}$$

for  $l > l_{\max}$  there are no stable solutions

$$l_{\max}^2 = \mu ka f(1.618)$$

By QM, there is a minimum possible <sup>(non-zero)</sup> value to  $l_{\min}^2 = \hbar^2$

if  $l_{\min} > l_{\max}$  there will be no bound states

if  $\hbar^2 > \mu ka f(1.618)$  no bound states

⇒ decrease # bound states by decreasing parameter  $\mu ka$

as  $\mu ka$  decreases,  $l_{\max}^2$  increases therefore fewer values of  $p$  such that

$$l^2 = p(1+p)\hbar^2 < l_{\max}^2$$

i.e. fewer bound states