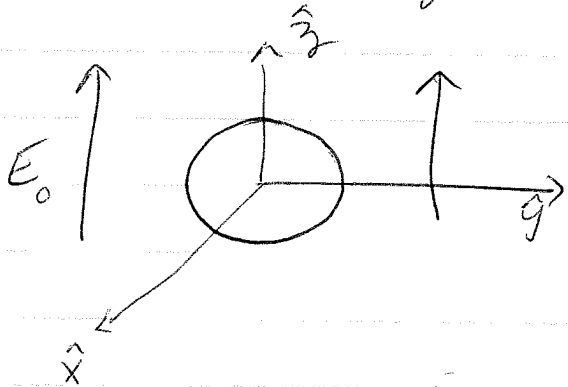


③ Grounded
 Conducting sphere in uniform electric field $\vec{E} = E_0 \hat{z}$

as $r \rightarrow \infty$ far from sphere, $\vec{E} = E_0 \hat{z} \Rightarrow \phi = -E_0 z$



boundary conditions $= -E_0 r \cos \theta$

$$\begin{cases} \phi(R, \theta) = 0 \\ \phi(r \rightarrow \infty, \theta) = -E_0 r \cos \theta \end{cases}$$

solution outside sphere has the form

$$\phi(r, \theta) = \sum_{l=0}^{\infty} \left[A_l r^l + \frac{B_l}{r^{l+1}} \right] P_l(\cos \theta)$$

From boundary condition as $r \rightarrow \infty$ we have

$$A_l = 0 \quad \text{all } l \neq 1$$

$$A_1 = -E_0 \quad \text{since } P_1(\cos \theta) = \cos \theta$$

$$\phi(r, \theta) = -E_0 r \cos \theta + \sum_{l=0}^{\infty} \frac{B_l}{r^{l+1}} P_l(\cos \theta)$$

From $\phi(R, \theta) = 0$ we have

$$0 = -E_0 R \cos \theta + \sum_{l=0}^{\infty} \frac{B_l}{R^{l+1}} P_l(\cos \theta)$$

$$\Rightarrow B_l = 0 \quad \text{all } l \neq 1$$

$$\frac{B_1}{R^2} = E_0 R \Rightarrow B_1 = +E_0 R^3$$

So

$$\phi(r, \theta) = -E_0 \left(r - \frac{R^3}{r^2} \right) \cos \theta$$

1st term is just potential $-E_0 r \cos \theta$ of the uniform applied electric field.

2nd term is potential due to the induced surface charge on the surface - it is a dipole field

Induced charge density is

$$4\pi\sigma(\theta) = -\left. \frac{\partial\phi}{\partial r} \right|_{r=R} = E_0 \left(1 + \frac{2R^3}{R^3} \right) \cos \theta \\ = 3E_0 \cos \theta$$

$$\sigma(\theta) = \frac{3}{4\pi} E_0 \cos \theta \quad \text{like uniformly polarized sphere} \quad k = \frac{3E_0}{4\pi}$$

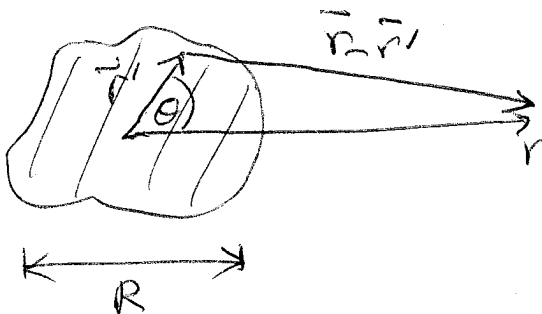
from ② we know that the field inside the sphere due to this σ is just $-\frac{4}{3}\pi k \hat{z} = -\frac{4}{3}\pi \frac{3E_0}{4\pi} \hat{z}$

$= -E_0 \hat{z}$. This is just what is required so that the total field in the conducting sphere vanishes,

Can check that outside the sphere, $\vec{E} = -\vec{\nabla}\phi$ is normal to surface of sphere at $r=R$.

Multipole Expansion

region with $\rho \neq 0$



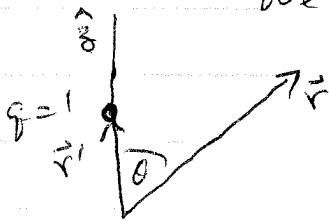
We want to find the potential ϕ for an arbitrary localized distribution of charge ρ , at distances far away $r \gg R$.

$$\phi(\vec{r}) = \int d^3r' \frac{\rho(\vec{r}')}{|\vec{r} - \vec{r}'|}$$

General Coulomb formula

We want an expansion of $\frac{1}{|\vec{r} - \vec{r}'|}$ in powers of $\left(\frac{r'}{r}\right)$ for $r \gg r'$

$\frac{1}{|\vec{r} - \vec{r}'|}$ view this as the potential at \vec{r} due to a unit point charge located at position \vec{r}' . We take \vec{r}' on the \hat{z} axis.



The problem has azimuthal symmetry $\Rightarrow \phi$ depends only on r and θ , so we can express it as an expansion in Legendre polynomials.

For $r > r'$,

$$\phi(r, \theta) = \sum_{l=0}^{\infty} \frac{B_l}{r^{l+1}} P_l(\cos\theta)$$

all $A_l = 0$

as need $\phi \rightarrow 0$

as $r \rightarrow \infty$

$$= \frac{1}{r} \sum_{l=0}^{\infty} \frac{B_l}{r^l} P_l(\cos\theta)$$

We know $\phi(r, \theta=0) = \frac{1}{r-r'}$ (for $r > r'$)

↳ scalars here since when $\theta=0$, \vec{r} and \vec{r}' are both on \hat{z} axis

$$\Rightarrow \phi(r, 0) = \frac{1}{r} \sum_e \frac{B_e}{r^e} P_e(1)$$

$$= \frac{1}{r} \sum_{e=0}^{\infty} \frac{B_e}{r^e} \quad \text{as } P_e(1) = 1$$

$$= \frac{1}{r} \frac{1}{(1-r'/r)} \leftarrow \text{exact result from Coulomb}$$

Now Taylor expansion $\frac{1}{1-e} = 1 + e + e^2 + e^3 + e^4 + \dots$

$$\Rightarrow \frac{1}{r} \sum_{e=0}^{\infty} \frac{B_e}{r^e} = \frac{1}{r} \left(1 + \frac{r'}{r} + \left(\frac{r'}{r}\right)^2 + \left(\frac{r'}{r}\right)^3 + \dots \right)$$

$$\Rightarrow B_e = (r')^e \text{ is solution}$$

So for $r > r'$

$$\boxed{\frac{1}{|\vec{r}-\vec{r}'|} = \frac{1}{r} \sum_{l=0}^{\infty} \left(\frac{r'}{r}\right)^l P_l(\cos \theta)}$$

So for the charge distribution ρ ,

$$\begin{aligned} \phi(\vec{r}) &= \int d^3r' \frac{\rho(\vec{r}')}{|\vec{r}-\vec{r}'|} = \int d^3r' \frac{\rho(\vec{r}')}{r} \sum_{l=0}^{\infty} \left(\frac{r'}{r}\right)^l P_l(\cos \theta) \\ &= \sum_{l=0}^{\infty} \frac{1}{r^{l+1}} \int d^3r' \rho(\vec{r}') (r')^l P_l(\cos \theta) \end{aligned}$$

where θ is the angle between the fixed observation point \vec{r} and the integration variable \vec{r}' .

This is the multipole expansion, which expresses the potential far from a localized source as a power series in (r'/r) . It is exact provided one adds all the infinite l terms. In practice, one generally approximates by summing only up to some finite l .

Note: in doing the integrals

$$\int d^3r' \rho(\vec{r}') (r')^l P_l(\cos\theta)$$

θ is defined as the angle of \vec{r}' with respect to observation point \vec{r} . We therefore in principle have to repeat this integration every time we change \vec{r} .

We will find a way around this by

- (i) just looking explicitly at the few lowest order terms
- (ii) a general method involving spherical harmonics $Y_{lm}(\theta, \phi)$

monopole: $l=0$ term

$$\phi^{(0)}(\vec{r}) = \frac{1}{r} \int d^3r' f(r')$$

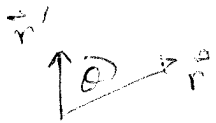
$$P_0(\cos\theta) = 1$$

$$= \frac{q}{r} \quad \text{where } q \equiv \int d^3r' f(r') \text{ is}$$

total charge

Dipole: $l=1$ term

$$\phi^{(1)}(\vec{r}) = \frac{1}{r^2} \int d^3r' f(\vec{r}') r' P_1(\cos\theta)$$



$$= \frac{1}{r^2} \int d^3r' f(\vec{r}') r' \cos\theta$$

Now $\hat{r} \cdot \vec{r}' = r r' \cos\theta \Rightarrow \hat{r} \cdot \vec{r}' = r' \cos\theta$

$$\phi^{(1)}(\vec{r}) = \frac{1}{r^2} \hat{r} \cdot \int d^3r' f(\vec{r}') \vec{r}'$$

$$= \frac{\vec{p} \cdot \hat{r}}{r^2} \quad \text{where } \vec{p} \equiv \int d^3r' f(\vec{r}') \vec{r}'$$

is the dipole moment

For a set of point charges q_i at \vec{r}_i ,

$$\vec{p} = \sum_i q_i \vec{r}_i$$

quadrupole: $l=2$ Term

$$\begin{aligned}\phi^{(2)}(\vec{r}) &= \frac{1}{r^3} \int d^3r' \rho(\vec{r}') r'^2 P_2(\cos\theta) \\ &= \frac{1}{r^3} \int d^3r' \rho(\vec{r}') r'^2 \frac{1}{2} (3\cos^2\theta - 1)\end{aligned}$$

use $\cos\theta = \hat{r}' \cdot \hat{r}$

$$\begin{aligned}\phi^{(2)}(\vec{r}) &= \frac{1}{r^3} \int d^3r' \rho(\vec{r}') \frac{1}{2} (3(\vec{r}' \cdot \hat{r})^2 - (r')^2) \\ &= \frac{1}{r^3} \hat{r} \cdot \left[\int d^3r' \rho(\vec{r}') \frac{1}{2} (3\vec{r}'\vec{r}' - (r')^2 \overset{\leftrightarrow}{\mathbf{I}}) \right] \cdot \hat{r}\end{aligned}$$

where $\overset{\leftrightarrow}{\mathbf{I}}$ is the identity tensor such that for any two vectors \vec{v} and \vec{u} , $\vec{u} \cdot \overset{\leftrightarrow}{\mathbf{I}} \cdot \vec{v} = \vec{u} \cdot \vec{v}$.

and $\vec{r}'\vec{r}'$ is the tensor such that for any two vectors \vec{v} and \vec{u} , $\vec{u} \cdot [\vec{r}'\vec{r}'] \cdot \vec{v} = (\vec{u} \cdot \vec{r}')(\vec{r}' \cdot \vec{v})$

Define quadrupole tensor $\overset{\leftrightarrow}{\mathbf{Q}} \equiv \int d^3r' \rho(\vec{r}') (3\vec{r}'\vec{r}' - (r')^2 \overset{\leftrightarrow}{\mathbf{I}})$

$$\phi^{(2)}(\vec{r}) = \frac{1}{r^3} \frac{1}{2} \hat{r} \cdot \overset{\leftrightarrow}{\mathbf{Q}} \cdot \hat{r}$$

So to lowest three terms

$$\phi(\vec{r}) = \frac{q}{r} + \frac{\vec{p} \cdot \hat{r}}{r^2} + \frac{\hat{r} \cdot \overset{\leftrightarrow}{\mathbf{Q}} \cdot \hat{r}}{2r^3} + \dots$$

defined in terms of the moments q , \vec{p} , $\overset{\leftrightarrow}{\mathbf{Q}}$ of the charge distribution.

Note, the moments q , \vec{P} , \vec{Q} do not depend on the observation point \vec{r} - we can calculate them once and then use them to get $\phi(\vec{r})$ at all \vec{r} .

monopole: $q = \int d^3r \rho(\vec{r})$ scalar integral

dipole: $\vec{P} = \int d^3r \rho(\vec{r}) \vec{r}$ vector integral
 $\hat{e}_1 \equiv \hat{x}, \hat{e}_2 \equiv \hat{y}, \hat{e}_3 \equiv \hat{z}$

if we pick a coordinate system, we have to do 3 integrations to get the three components of \vec{P}

$$\hat{e}_i \cdot \vec{P} = P_i = \int d^3r \rho(\vec{r}) r_i$$

quadrupole: $\vec{Q} = \int d^3r \rho(\vec{r}) (3\vec{r} \vec{r} - r^2 \vec{I})$ tensor integral

if we pick a coord system x, y, z then

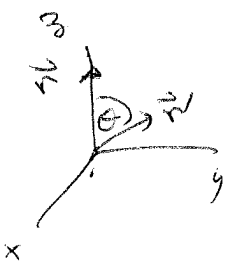
\vec{Q} is a matrix with components $\hat{e}_1 \equiv \hat{x}, \hat{e}_2 \equiv \hat{y}, \hat{e}_3 \equiv \hat{z}$

$$\hat{e}_i \cdot \vec{Q} \cdot \hat{e}_j = Q_{ij} = \int d^3r \rho(\vec{r}) [3r_i r_j - r^2 \delta_{ij}]$$

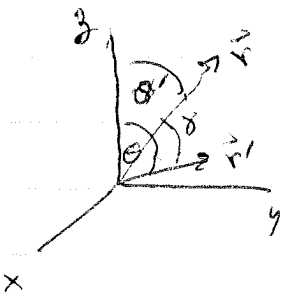
There are 9 elements of the 3×3 matrix (Q_{ij}) , but $Q_{ij} = Q_{ji}$ is symmetric so there are only 6 independent elements to compute.

General method

$$\phi(\vec{r}) = \sum_{l=0}^{\infty} \frac{1}{r^{l+1}} \int d^3r' \rho(\vec{r}') (r')^l P_l(\cos\theta)$$



in above, θ is angle between \vec{r} and \hat{z}
 if we think of $P_l \cos\theta$ as the spherical coord θ ,
 then in effect, above is choosing \vec{r} to be on
 \hat{z} axis. We would like a representation in
 which \vec{r} is positioned arbitrarily with respect
 to the axes used in describing ρ



Use the addition theorem for spherical harmonics
 - see Jackson 3.6 for discussion & proof

$$P_l(\cos\delta) = \frac{4\pi}{2l+1} \sum_{m=-l}^l Y_{lm}^*(\theta', \phi') Y_{lm}(\theta, \phi)$$

where (θ, ϕ) are the angles of \hat{r} , (θ', ϕ') are
 the angles of \hat{r}' , and δ is the angle
 between \hat{r} and \hat{r}' , i.e.

$$\cos\delta = \hat{r} \cdot \hat{r}'$$

$$\cos\theta = \hat{z} \cdot \hat{r}$$

$$\cos\theta' = \hat{z} \cdot \hat{r}'$$

\Rightarrow

$$\phi(\vec{r}) = \sum_{l=0}^{\infty} \frac{1}{r^{l+1}} \frac{4\pi}{2l+1} \sum_{m=-l}^l \int d^3r' \rho(\vec{r}') (r')^l Y_{lm}^*(\theta', \phi') Y_{lm}(\theta, \phi)$$

Define the moment

$$Q_{lm} \equiv \int d^3r' \rho(\vec{r}') (r')^l Y_{lm}^*(\theta', \phi')$$

independent of observation point

Then

$$\phi(\vec{r}) = 4\pi \sum_{l=0}^{\infty} \sum_{m=-l}^l \frac{f_{lm} Y_{lm}(\theta, \phi)}{(2l+1) r^{l+1}}$$

see Jackson eqn (4.4), (4.5), (4.6) to relate f_{lm} to q , \vec{p} , \vec{Q} .

$$\phi(\vec{r}) = \frac{q}{r} + \frac{\vec{p} \cdot \hat{r}}{r^2} + \frac{\hat{r} \cdot (\vec{Q} \times \hat{r})}{2r^3} + \dots$$

electric field $\vec{E} = -\vec{\nabla}\phi = -\frac{\partial\phi}{\partial r} \hat{r} + \frac{1}{r} \frac{\partial\phi}{\partial\theta} \hat{\theta} + \frac{1}{r\sin\theta} \frac{\partial\phi}{\partial\phi} \hat{\phi}$

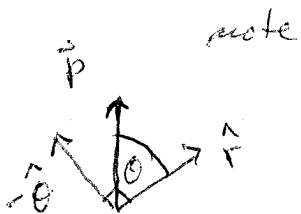
For the monopole term $\vec{E} = \frac{q}{r^2} \hat{r}$

For the dipole term, choose \vec{p} along \hat{z} axis so

$$\phi(\vec{r}) = \frac{p \cos\theta}{r^2}$$

$$\vec{E} = \frac{2p \cos\theta}{r^3} \hat{r} + \frac{p \sin\theta}{r^3} \hat{\theta}$$

$$\vec{E} = \frac{p}{r^3} (2\cos\theta \hat{r} + \sin\theta \hat{\theta})$$



$$p \cos\theta \hat{r} = (\vec{p} \cdot \hat{r}) \hat{r}$$

$$p \sin\theta \hat{\theta} = -(\vec{p} \cdot \hat{\theta}) \hat{\theta}$$

$$\text{Now } \vec{p} = (\vec{p} \cdot \hat{r}) \hat{r} + (\vec{p} \cdot \hat{\theta}) \hat{\theta}$$

$$\Rightarrow -(\vec{p} \cdot \hat{\theta}) \hat{\theta} = (\vec{p} \cdot \hat{r}) \hat{r} - \vec{p}$$

so

$$\vec{E} = \frac{1}{r^3} \left[2(\vec{p} \cdot \hat{r}) \hat{r} + (\vec{p} \cdot \hat{r}) \hat{r} - \vec{p} \right] = \frac{1}{r^3} \left[3(\vec{p} \cdot \hat{r}) \hat{r} - \vec{p} \right]$$

expresses \vec{E} in coord. form