

if do not make radiation zone approx, one gets

$$\vec{E}_{EI} = \frac{k^2 e^{ikr}}{r} \left[\vec{p}_\omega - \hat{r}(\vec{p}_\omega \cdot \hat{r}) - \frac{i}{kr} (1 + \frac{i}{kr}) (3\hat{r}(\vec{p}_\omega \cdot \hat{r}) - \vec{p}_\omega) \right]$$

Using radiation zone approx:

$$\vec{E}_{EI} = k^2 \frac{e^{ikr}}{r} \hat{r} \times (\vec{p}_\omega \times \hat{r})$$

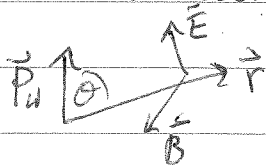
$$|\vec{E}_{EI}| = |\vec{B}_{EI}|$$

$$\vec{B}_{EI} = -k^2 \frac{e^{ikr}}{r} \vec{p}_\omega \times \hat{r}$$

$$\vec{E}_{EI} \perp \vec{B}_{EI}$$

If \vec{p}_ω is a real vector, then

If choose coordinates so that \vec{p}_ω is along \hat{z} axis, then



$$\vec{E}_{EI} = -k^2 p_\omega \frac{e^{ikr}}{r} \sin\theta \hat{\theta}$$

$$\vec{B}_{EI} = -k^2 p_\omega \frac{e^{ikr}}{r} \sin\theta \hat{\phi}$$

Emitted power

Pointing vector $\vec{S}_{EI}(\vec{r}, t) = \frac{c}{4\pi} \text{Re}\{\vec{E}_{EI}\} \times \text{Re}\{\vec{B}_{EI}\}$

need to take real parts of complex expressions before multiplying

$$\text{Re}\{\vec{E}_{EI}(\vec{r}, t)\} = -k^2 p_\omega \frac{\cos(kr - \omega t)}{r} \sin\theta \hat{\theta}$$

$$\text{Re}\{\vec{B}_{EI}(\vec{r}, t)\} = -k^2 p_\omega \frac{\cos(kr - \omega t)}{r} \sin\theta \hat{\phi}$$

$$\boxed{\vec{S}_{EI}(\vec{r}, t) = \frac{c}{4\pi} k^4 p_\omega^2 \frac{\cos^2(kr - \omega t)}{r^2} \sin^2\theta \hat{r}}$$

$\vec{S}_{EI} \sim \hat{r} \Rightarrow$ energy flows radially outwards

$\vec{S}_{EI} \sim \frac{1}{r^2} \Rightarrow$ energy conserved

$$\oint da \hat{n} \cdot \langle \vec{S}_{EI} \rangle = \text{constant for all } R$$

sphere
radius R

→ Question - what about the

time averaged energy current

$\frac{1}{r^n}, n > 2$, terms if we do not
make radiation zone approx?

$$\langle \vec{S}_{EI} \rangle = \frac{1}{T} \int_0^T dt \vec{S}_{EI}(\vec{r}, t)$$

$$T \text{ is period } T = \frac{2\pi}{\omega}$$

$$\langle \cos^2(\dots) \rangle = \frac{1}{2}$$

$$= \frac{c}{8\pi} k^4 p_\omega^2 \frac{\sin^2 \theta}{r^2} \hat{r}$$

average energy flowing through an element
of area at spherical angles θ, ϕ is

$$dP_{EI} = \hat{r} \cdot \langle \vec{S}_{EI} \rangle \underbrace{r^2 \sin \theta d\theta d\phi}_{\text{area of surface element}}$$

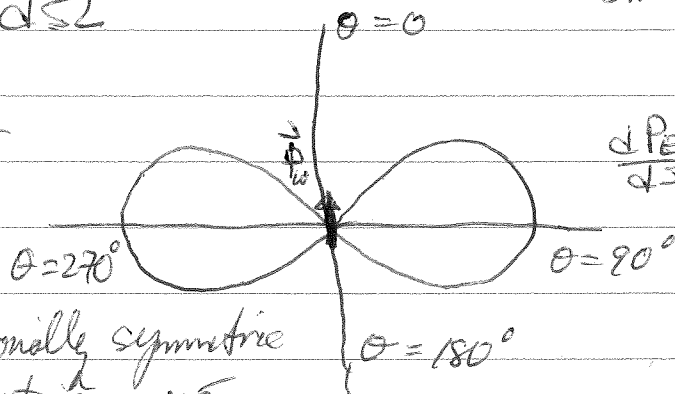
area of surface element

$$= r^2 d\Omega \quad \Omega \text{ is solid angle}$$

$$= \hat{r} \cdot \langle \vec{S}_{EI} \rangle r^2 d\Omega$$

$$\frac{dP_{EI}}{d\Omega} = \hat{r} \cdot \langle \vec{S}_{EI} \rangle r^2 = \frac{c}{8\pi} k^4 p_\omega^2 \sin^2 \theta \sim \omega^4 \sin^2 \theta$$

polar plot



rotationally symmetric
about \hat{z} axis

$$\frac{dP_{EI}}{d\Omega} \sim \sin^2 \theta$$

most of power is
directed outwards
into plane $\perp \vec{P}_\omega$
ie peaked about $\theta = 90$

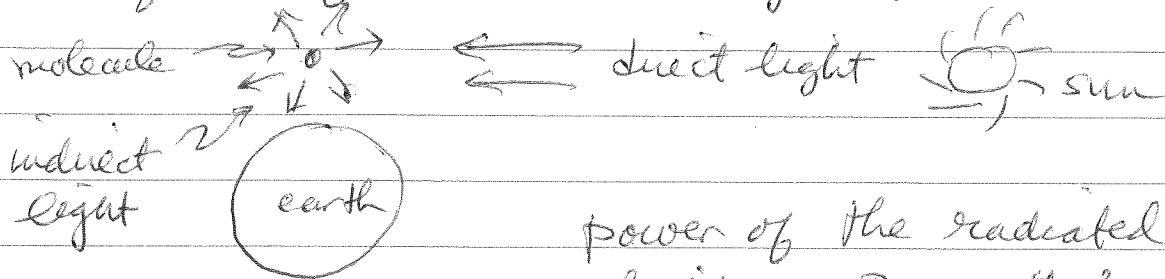
Total power radiated is

$$P_{EI} = \int \frac{dP_{EI}}{d\Omega} d\Omega = \frac{ck^4 p_w^2}{8\pi} \underbrace{2\pi \int_0^{\pi} \sin \theta \sin^2 \theta}_{4/3}$$

$$P_{EI} = \frac{ck^4 p_w^2}{3} = \frac{p_w^2 \omega^4}{3c^3} \sim \omega^4$$

why the sky is blue - Lord Rayleigh

when look up at sky, you are seeing the indirect light of the sun, which is the light emitted by the atoms and molecules of the atmosphere as they oscillate, and so radiate, due to the electric field of the direct light from the sun



power of the radiated indirect light is $P \sim \omega^4 p_w^2$

$$\vec{P} = \alpha \vec{E} \quad \alpha \sim \frac{e^2}{m} \frac{1}{\omega_0^2 - \omega^2 - i\omega\gamma}$$

For molecules in atmosphere (N_2 etc) ω_0 is typically at a freq higher than visible spectrum. Therefore, in visible spectrum $\alpha \sim \frac{e^2}{m\omega_0^2}$ indep of ω .

\Rightarrow power radiated is $P \sim \omega^4$

$P \sim \omega^4$ largest at high freq

Since light from sun is "white light"

it has components of all freqs. Of these freqs, the higher ones are scattered the most & make up the indirect light we see.

Since blue is the largest ω in visible spectrum, the sky is blue!

When we look at sunrise or sunset, we are looking at the direct rays of the sun. Since these rays are least scattered at low $\omega \Rightarrow$ sunset and sunrise are red!

Magnetic Dipole approx - Radiation Zone for $\gg 1$

$$\vec{A}_{MI} = \frac{e^{ikr}}{r} \left(\frac{1}{r} - ik \right) (-\hat{r} \times \vec{m}_{\omega})$$
$$\approx ik \hat{r} \times \vec{m}_{\omega} \frac{e^{ikr}}{r} \quad \text{in RZ}$$

$$\vec{B}_{MI} = \vec{\nabla} \times \vec{A}_{MI} = (\vec{\nabla} e^{ikr}) \times \left(\frac{ik \hat{r} \times \vec{m}_{\omega}}{r} \right)$$
$$+ e^{ikr} \vec{\nabla} \times \left(\frac{ik \hat{r} \times \vec{m}_{\omega}}{r} \right)$$

will give terms of $o\left(\frac{1}{r^2}\right)$
so ignore in RZ approx

$$\vec{B}_{MI} = -k^2 \frac{e^{ikr}}{r} \hat{r} \times (\hat{r} \times \vec{m}_{\omega})$$

From Ampere's Law

$$\vec{E}_{MI} = \frac{c}{k} \vec{\nabla} \times \vec{B}_{MI} = -ik (\vec{\nabla} e^{ikr}) \times \left(\frac{\hat{r} \times [\hat{r} \times \vec{m}_{\omega}]}{r} \right)$$
$$= ik e^{ikr} \vec{\nabla} \times \left(\frac{\hat{r} \times [\hat{r} \times \vec{m}_{\omega}]}{r} \right)$$

will give terms of $o\left(\frac{1}{r^2}\right)$
so ignore in RZ approx

$$\vec{E}_{MI} = k^2 \frac{e^{ikr}}{r} \hat{r} \times (\hat{r} \times (\hat{r} \times \vec{m}_{\omega}))$$

triple product rule

$$= k^2 \frac{e^{ikr}}{r} \left\{ \hat{r} [\hat{r} \cdot (\hat{r} \times \vec{m}_{\omega})] - (\hat{r} \times \vec{m}_{\omega}) [\hat{r} \cdot \hat{r}] \right\}$$

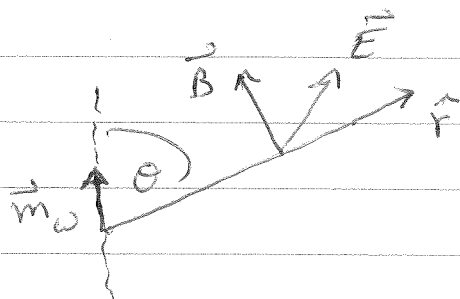
$$\vec{E}_{MI} = -\frac{k^2}{r} e^{ikr} (\hat{r} \times \vec{m}_{\omega})$$

If \vec{m}_ω is a real vector, then

If align axes so that $\vec{m}_\omega = m_\omega \hat{z}$ then

$$\vec{E}_{M1} = m_\omega \frac{k^2}{r} e^{i(kr - \omega t)} \sin\theta \hat{\phi}$$

$$\vec{B}_{M1} = -m_\omega \frac{k^2}{r} e^{i(kr - \omega t)} \sin\theta \hat{\theta}$$



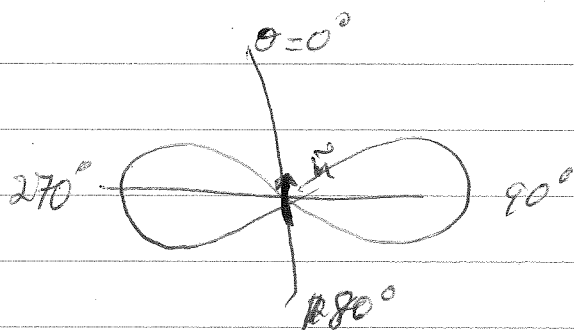
Poynting vector

$$\vec{S}_{M1} = \frac{c}{4\pi} \text{Re}\{\vec{E}_{M1}\} \times \text{Re}\{\vec{B}_{M1}\}$$

$$= \frac{c}{4\pi} \frac{k^4 m_\omega^2}{v^2} \cos^2(kr - \omega t) \sin^2\theta \hat{r}$$

$$\langle \vec{S}_{M1} \rangle = \frac{c}{8\pi} \frac{k^4 m_\omega^2}{v^2} \sin^2\theta \hat{r}$$

$$\frac{dP_{M1}}{d\Omega} = \hat{r} \cdot \langle \vec{S}_{M1} \rangle r^2 = \frac{c}{8\pi} k^4 m_\omega^2 \sin^2\theta \sim \omega^4 \sin^2\theta$$



rotationally symmetric
about \hat{z} axis

$$P_{M1} = \int d\Omega \frac{dP_{M1}}{d\Omega} = \frac{c k^4}{3} m_\omega^2 = \frac{m_\omega^2 \omega^4}{30^3}$$

$$\frac{P_{M1}}{P_{E1}} = \frac{m_\omega^2}{P_\omega^2} \sim \left(\frac{v}{c}\right)^2, \quad \left. \begin{array}{l} m_\omega \sim \frac{df}{c} \sim dg \frac{v}{c} \\ P_\omega \sim dg \end{array} \right\} \Rightarrow \frac{m_\omega}{P_\omega} \sim \frac{v}{c}$$

Electric Quadrupole radiation - radiation zone approx

$$\begin{aligned}\vec{A}_{E2} &= \frac{e^{i\vec{k}\cdot\vec{r}}}{r} \left(\frac{1}{r} - ik \right) \left(\frac{-i\omega}{6c} \hat{r} \cdot \vec{Q}_\omega \right) \\ &= -\frac{e^{i\vec{k}\cdot\vec{r}}}{r} \frac{k^2}{6} \hat{r} \cdot \vec{Q}_\omega \quad \text{in RZ approx}\end{aligned}$$

$$\begin{aligned}\vec{B}_{E2} &= \vec{\nabla} \times \vec{A}_{E2} = -(\vec{\nabla} e^{i\vec{k}\cdot\vec{r}}) \times \left[\frac{k^2 \hat{r} \cdot \vec{Q}_\omega}{6r} \right] \\ &= e^{i\vec{k}\cdot\vec{r}} \vec{\nabla} \times \left[\frac{k^2 \hat{r} \cdot \vec{Q}_\omega}{6r} \right]\end{aligned}$$

$\mathcal{O}\left(\frac{1}{r^2}\right)$ so ignore in RZ approx

$$\vec{B}_{E2} = -ik^3 \frac{e^{i\vec{k}\cdot\vec{r}}}{6r} \hat{r} \times (\hat{r} \cdot \vec{Q}_\omega)$$

$$\begin{aligned}\vec{E}_{E2} &= \frac{i}{k} \vec{\nabla} \times \vec{B}_{E2} = k^2 (\vec{\nabla} e^{i\vec{k}\cdot\vec{r}}) \times \left[\frac{\hat{r} \times (\hat{r} \cdot \vec{Q}_\omega)}{6r} \right] \\ &+ k^2 e^{i\vec{k}\cdot\vec{r}} \vec{\nabla} \times \left[\frac{\hat{r} \times (\hat{r} \cdot \vec{Q}_\omega)}{6r} \right]\end{aligned}$$

$\mathcal{O}\left(\frac{1}{r}\right)$ so ignore in RZ approx

$$\vec{E}_{E2} = ik^3 \frac{e^{i\vec{k}\cdot\vec{r}}}{6r} \hat{r} \times \left[\hat{r} \times (\hat{r} \cdot \vec{Q}_\omega) \right]$$

triple product rule

$$= ik^3 \frac{e^{i\vec{k}\cdot\vec{r}}}{6r} \left\{ \hat{r} \left[\hat{r} \cdot (\hat{r} \cdot \vec{Q}_\omega) \right] - (\hat{r} \cdot \vec{Q}_\omega) [\hat{r} \cdot \hat{r}] \right\}$$

$$\vec{E}_{E2} = \frac{ike^3 e^{i\vec{k}\cdot\vec{r}}}{6r} \left\{ \hat{r} (\hat{r} \cdot \vec{Q}_\omega \cdot \hat{r}) - (\hat{r} \cdot \vec{Q}_\omega) \right\}$$

Poynting vector

$$\vec{S}_{E2} = \frac{c}{4\pi} \operatorname{Re} \{ \vec{E}_{E2} \} \times \operatorname{Re} \{ \vec{B}_{E2} \} \quad \text{If } \vec{Q}_w \text{ is real then,}$$

$$= \frac{-c}{4\pi} \frac{k^6}{36r^2} \sin^2(kr - \omega t) \left\{ \hat{r} (\hat{r} \cdot \vec{Q}_w \cdot \hat{r}) - (\hat{r} \cdot \vec{Q}_w) \right\} \times \left[\hat{r} \times (\hat{r} \cdot \vec{Q}_w) \right]$$

$$= \frac{-c}{4\pi} \frac{k^6}{36r^2} \sin^2(kr - \omega t) \left\{ \hat{r} \left[\hat{r} (\hat{r} \cdot \vec{Q}_w \cdot \hat{r}) \cdot (\hat{r} \cdot \vec{Q}_w) \right] \right. \\ \left. - (\hat{r} \cdot \vec{Q}_w) \left[\hat{r} (\hat{r} \cdot \vec{Q}_w \cdot \hat{r}) \cdot \hat{r} \right] \right.$$

$$\left. - \hat{r} \left[(\hat{r} \cdot \vec{Q}_w) \cdot (\hat{r} \cdot \vec{Q}_w) \right] \right. \\ \left. + (\hat{r} \cdot \vec{Q}_w) \left[(\hat{r} \cdot \vec{Q}_w) \cdot \hat{r} \right] \right\}$$

$$= \frac{-c}{4\pi} \frac{k^6}{36r^2} \sin^2(kr - \omega t) \left\{ (\hat{r} \cdot \vec{Q}_w \cdot \hat{r})^2 \hat{r} - (\hat{r} \cdot \vec{Q}_w) (\hat{r} \cdot \vec{Q}_w \cdot \hat{r}) \right. \\ \left. - (\hat{r} \cdot \vec{Q}_w \cdot \vec{Q}_w \cdot \hat{r}) \hat{r} + (\hat{r} \cdot \vec{Q}_w) (\hat{r} \cdot \vec{Q}_w \cdot \hat{r}) \right\}$$

$$\vec{S}_{E2} = \frac{-c}{4\pi} \frac{k^6}{36r^2} \sin^2(kr - \omega t) \left\{ (\hat{r} \cdot \vec{Q}_w \cdot \hat{r})^2 - (\hat{r} \cdot \vec{Q}_w)^2 \right\} \hat{r}$$

$$\langle \vec{S}_{E2} \rangle = \frac{-ck^6}{4\pi 72r^2} \left\{ (\hat{r} \cdot \vec{Q}_w \cdot \hat{r})^2 - (\hat{r} \cdot \vec{Q}_w)^2 \right\} \hat{r}$$

$$\frac{dP_{E2}}{d\Omega} = \hat{r} \cdot \langle \vec{S}_{E2} \rangle r^2 = \frac{ck^6}{4\pi 72} \left\{ (\hat{r} \cdot \vec{Q}_w)^2 - (\hat{r} \cdot \vec{Q}_w \cdot \hat{r})^2 \right\}$$

angular dependence of $\frac{dP_{E2}}{d\Omega}$ depends
on specific form of the tensor \vec{Q}_w

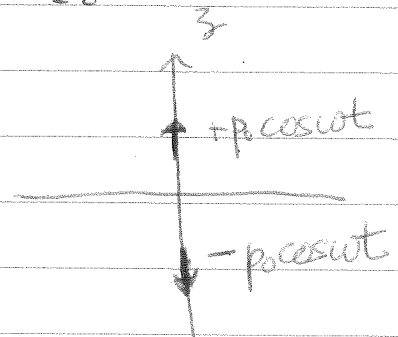
For example: suppose $Q_{ij} = 0$ except for Q_{zz}

$$\Rightarrow \vec{Q} \omega = Q_{zz} \hat{z} \hat{z}$$

$$(\hat{r} \cdot \vec{Q} \omega \cdot \hat{r})^2 = (Q_{zz} \cos^2 \theta)^2$$

$$(\hat{r} \cdot \vec{Q} \omega)^2 = Q_{zz}^2 \cos^2 \theta$$

ex



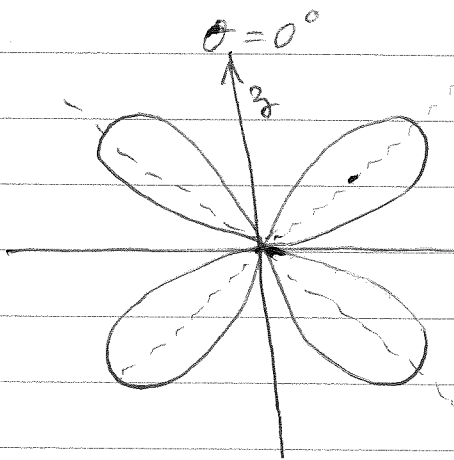
$$\frac{dP_{E2}}{d\Omega} = \frac{c k^6}{4\pi^2} Q_{zz}^2 [\cos^2 \theta - \cos^4 \theta]$$

$$= \frac{c k^6}{4\pi^2} Q_{zz}^2 \cos^2 \theta \sin^2 \theta$$

$$\cos \theta \sin \theta = \frac{1}{2} \sin 2\theta$$

$$= \frac{c k^6}{4\pi^2} Q_{zz}^2 \sin^2 2\theta$$

$\frac{dP_{E2}}{d\Omega}$



peak at 45°

rotationally invariant about \hat{z} axis

$$\frac{P_{E2}}{P_{E1}} \sim \frac{k^6 Q^2}{k^4 p^2} \sim \frac{k^2 (q d^2)^2}{(q d)^2} \sim k^2 d^2 \sim \left(\frac{v}{c}\right)^2$$

$$P_{E2} \sim P_{M1}$$

For more general case, choose axes so that $\vec{Q} \cdot \vec{w}$ is diagonal - can always do this since $\vec{Q} \cdot \vec{w}$ is symmetric

$$(\hat{r} \cdot \vec{Q} \cdot \hat{r}) = \hat{r} \cdot \begin{pmatrix} Q_{xx} & 0 & 0 \\ 0 & Q_{yy} & 0 \\ 0 & 0 & Q_{zz} \end{pmatrix} \cdot \hat{r}$$

$$= \hat{r} \cdot \begin{pmatrix} Q_{xx} \sin\theta \cos\varphi \\ Q_{yy} \sin\theta \sin\varphi \\ Q_{zz} \cos\theta \end{pmatrix} = Q_{xx} \sin^2\theta \cos^2\varphi + Q_{yy} \sin^2\theta \sin^2\varphi + Q_{zz} \cos^2\theta$$

$$(\hat{r} \cdot \vec{Q} \cdot \hat{r})^2 = Q_{xx}^2 \sin^4\theta \cos^4\varphi + Q_{yy}^2 \sin^4\theta \sin^4\varphi + Q_{zz}^2 \cos^4\theta$$

$$\frac{dP_{E2}}{d\Omega} = \frac{k^6}{72} \left\{ Q_{zz}^2 (\cos^2\theta - \cos^4\theta) + Q_{xx}^2 (\sin^2\theta \cos^2\varphi - \sin^4\theta \cos^4\varphi) + Q_{yy}^2 (\sin^2\theta \sin^2\varphi - \sin^4\theta \sin^4\varphi) \right\}$$

$$\frac{dP_{E2}}{d\Omega} = \frac{k^6}{72} \left\{ Q_{zz}^2 \cos^2\theta \sin^2\theta + Q_{xx}^2 \sin^2\theta \cos^2\varphi (1 - \sin^2\theta \cos^2\varphi) + Q_{yy}^2 \sin^2\theta \sin^2\varphi (1 - \sin^2\theta \sin^2\varphi) \right\}$$

no special symmetries - varies with θ and φ

For arbitrary charge distributions - not pure harmonic

For $\vec{p}_\omega e^{-i\omega t}$ pure harmonic oscillation, we found the radiated fields in electric dipole approx are

$$\vec{E} = \vec{E}_\omega e^{-i\omega t}, \quad \vec{B} = \vec{B}_\omega e^{-i\omega t}$$

$$\vec{E}_\omega = -k^2 \frac{e^{-i\omega r}}{r} \hat{r} \times (\hat{r} \times \vec{p}_\omega) = -\frac{\omega^2}{c^2} \frac{e^{i\omega r/c}}{r} \hat{r} \times (\hat{r} \times \vec{p}_\omega)$$

$$\vec{B}_\omega = k^2 \frac{e^{-i\omega r}}{r} (\hat{r} \times \vec{p}_\omega) = \frac{\omega^2}{c^2} \frac{e^{i\omega r/c}}{r} (\hat{r} \times \vec{p}_\omega)$$

$$\text{as } k = \frac{\omega}{c}$$

For an arbitrarily time varying charge distribution with electric dipole moment

$$\vec{p}(t) = \int \frac{d\omega}{2\pi} \vec{p}_\omega e^{-i\omega t}$$

then solution for fields given by superposition

$$\vec{E}(\vec{r}, t) = \int \frac{d\omega}{2\pi} \vec{E}_\omega e^{-i\omega t}$$

$$= - \int \frac{d\omega}{2\pi} \frac{e^{-i\omega(t-r/c)}}{r} \left(\frac{\omega^2}{c^2} \right) \hat{r} \times (\hat{r} \times \vec{p}_\omega)$$

$$= \frac{-1}{c^2 r} \hat{r} \times \left[\hat{r} \times \int \frac{d\omega}{2\pi} e^{-i\omega(t-r/c)} \vec{p}_\omega \omega^2 \right]$$

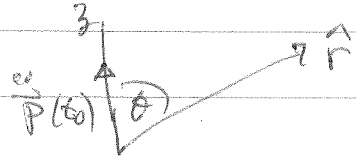
$$= \frac{1}{c^2 r} \hat{r} \times \left[\hat{r} \times \frac{\partial^2}{\partial t^2} \int \frac{d\omega}{2\pi} e^{-i\omega(t-r/c)} \vec{p}_\omega \right]$$

$$\boxed{\vec{E}(\vec{r}, t) = \frac{1}{c^2 r} \hat{r} \times \left[\hat{r} \times \ddot{\vec{p}}(t - r/c) \right]} \quad \ddot{\vec{p}} = \frac{d^2 \vec{p}}{dt^2}$$

define $t_0 \equiv t - r/c$ = "retarded time"

in spherical coords, if $\ddot{\vec{p}}(t_0)$ is along \hat{z}

$$\vec{E}(\vec{r}, t) = \frac{\ddot{p}(t_0)}{c^2 r} \sin \theta \hat{\theta}$$



Similarly

$$\vec{B}(\vec{r}, t) = \int \frac{d\omega}{2\pi} \vec{B}_\omega e^{-i\omega t}$$

$$= \int \frac{d\omega}{2\pi} \frac{e^{-i\omega(t - r/c)}}{r} \left(\frac{\omega^2}{c^2} \right) (\hat{r} \times \vec{p}_\omega)$$

$$= \frac{-1}{c^2 r} \hat{r} \times \frac{\partial^2}{\partial t^2} \int \frac{d\omega}{2\pi} e^{-i\omega(t - r/c)} \vec{p}_\omega$$

$$\boxed{\vec{B}(\vec{r}, t) = \frac{-1}{c^2 r} \hat{r} \times \ddot{\vec{p}}(t_0)}$$

$$\vec{B}(\vec{r}, t) = \frac{\ddot{p}(t_0)}{c^2 r} \sin \theta \hat{\phi} \quad \text{in spherical coords}$$

Poynting vector

$$\vec{S} = \frac{c}{4\pi} \vec{E} \times \vec{B} = \frac{c}{4\pi} \left(\frac{1}{c^2 r} \right)^2 \left[\ddot{p}(t_0) \right]^2 \sin^2 \theta \hat{r}$$

Total power radiated through a sphere of radius r is

$$\begin{aligned}
 P &= \oint da \hat{r} \cdot \vec{S} = 2\pi \int_0^\pi d\theta \sin\theta r^2 \hat{r} \cdot \vec{S} \\
 &= \frac{[\ddot{p}(t_0)]^2}{2c^3} \underbrace{\int_0^\pi d\theta \sin^3\theta}_{4/3}
 \end{aligned}$$

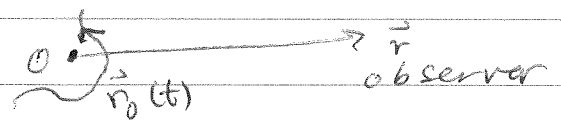
$$P = \frac{2}{3c^3} [\ddot{p}(t_0)]^2$$

For a point charge moving along a trajectory $\vec{r}_0(t)$

$$\vec{F}(t) = q \vec{r}_0(t)$$

$$\ddot{\vec{p}}(t) = q \ddot{\vec{r}}_0(t) = q \vec{a}(t)$$

↑ acceleration



$$P = \frac{2}{3} \frac{q^2 a^2(t_0)}{c^3}$$

Larmor's formula

← total power passing through a sphere of radius r at time t is due to acceleration at retarded time $t_0 = t - r/c$

power radiated \propto (acceleration)²

Larmor's formula above only holds in the non-relativistic limit since it is based on the electric dipole approx.