

Relation of Q_N to Helmholtz free energy.

Now $A = E - TS \Rightarrow E = A + TS = A - T \left(\frac{\partial A}{\partial T} \right)_{V,N}$

$$\begin{aligned} E &= A - T \left(\frac{\partial A}{\partial T} \right)_{V,N} = -T^2 \left[\frac{\partial}{\partial T} \left(\frac{A}{T} \right) \right]_{N,V} \\ &= \left[\frac{\partial (A/T)}{\partial (1/T)} \right]_{N,V} = \frac{\partial (\beta A)}{\partial \beta}_{N,V} \end{aligned}$$

Compare above $E = -\frac{\partial}{\partial \beta} (-\beta A)$

with $\langle E \rangle = -\frac{\partial}{\partial \beta} (\ln Q_N)$

$$\Rightarrow \boxed{A(T, V) = -k_B T \ln Q_N(T, V)}$$

compare to

$$S(E, V, N) = k_B \ln \Omega(E, V, N)$$

$$Q_N(T, V) = \int \frac{dE}{\Delta} \Omega(E) e^{-\beta E}$$

$\Rightarrow Q_N(T)$ is Laplace transform of $\Omega(E)$

A is Legendre transform of E

Note: Above relates A in canonical ensemble to average energy $\langle E \rangle$ in canonical ensemble $\langle S \rangle$

But original thermodynamic definition, in which E and $S(E)$ are fixed, should better be thought of as a relation in the microcanonical ensemble. Are the two formulations equivalent? We will soon see.

Energy fluctuations - In canonical ensemble, E is not fixed, but has a prob distr.
How wide is the distr in E ?

Consider

$$\begin{aligned}\frac{\partial \langle E \rangle}{\partial \beta} &= \frac{\partial}{\partial \beta} \left[\frac{\int dE E \Omega(E) e^{-\beta E}}{Q_N} \right] \\&= \frac{\int dE E \Omega(E) e^{-\beta E} (-E)}{Q_N} - \frac{\int dE E \Omega(E) e^{-\beta E}}{Q_N^2} \frac{\partial Q_N}{\partial \beta} \\&= -\langle E^2 \rangle - \langle E \rangle \frac{\partial \ln Q_N}{\partial \beta} \\&= -\langle E^2 \rangle + \langle E \rangle^2\end{aligned}$$

So

$$\langle E^2 \rangle - \langle E \rangle^2 = -\frac{\partial \langle E \rangle}{\partial \beta} = \frac{\partial^2}{\partial \beta^2} (\ln Q_N) = -\frac{\partial^2}{\partial \beta^2} (\beta A)$$

Note: $\langle (E - \langle E \rangle)^2 \rangle = \langle E^2 - 2E\langle E \rangle + \langle E \rangle^2 \rangle$

$$= \langle E^2 \rangle - 2\langle E \rangle \langle E \rangle + \langle E \rangle^2 = \langle E^2 \rangle - \langle E \rangle^2$$

So fluctuation in energy is

$$\begin{aligned}\langle (E - \langle E \rangle)^2 \rangle &= \langle E^2 \rangle - \langle E \rangle^2 = -\frac{\partial \langle E \rangle}{\partial \beta} = -k_B \frac{\partial \langle E \rangle}{\partial (1/T)} = k_B T^2 \frac{\partial \langle E \rangle}{\partial T} \\&= k_B T^2 C_V\end{aligned}$$

↑ specific heat at constant volume

Note: $\langle E \rangle \sim N$, $C_V \sim N$

$$\Rightarrow \frac{\sqrt{\langle E^2 \rangle - \langle E \rangle^2}}{\langle E \rangle} = \frac{\sqrt{k_B T^2 C_V}}{\langle E \rangle} \sim \frac{\sqrt{N}}{N} = \frac{1}{\sqrt{N}} \rightarrow 0 \text{ as } N \rightarrow \infty$$

Relation between Canonical A and Microcanonical A

$$A = E - TS$$

A is Legendre transform of $E(S, V, N)$ with respect to S .

Can also write

$$-\frac{A}{T} = S - \left(\frac{1}{T}\right)E \quad \text{where } \left(\frac{\partial S}{\partial E}\right)_{N,V} = \frac{1}{T}$$

so $-\frac{A}{T}$ is the Legendre transform of S with respect to E .

$$\Rightarrow -\frac{A}{T} = \max_E \left[S(E, V, N) - \frac{E}{T} \right]$$

$$\text{or } A = \min_E \left[E - TS(E, V, N) \right]$$

Having computed $S(E, V, N) = k_B \ln \Omega(E, V, N)$ in microcanonical ensemble, the above gives the corresponding microcanonical Helmholtz free energy A .

In microcanonical ensemble we can write

$$A(T, V, N) = \min_E [E - TS(E, V, N)] = \bar{E} - TS(\bar{E}, V, N)$$

Compare now to canonical ensemble calculation of A

$$Q_N = e^{-A/k_B T} = \int \frac{dE}{\Delta} \Omega(E) e^{-E/k_B T} = \int \frac{dE}{\Delta} e^{-(E - TS(E))/k_B T}$$

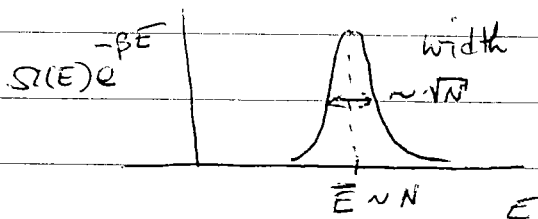
Consider $(E - TS(E))/k_B T$ expand about its ^{minimum} ~~maximum~~ at $\bar{E}$
 $E = \bar{E} + \delta E$

$$E - TS(E) \approx \bar{E} - TS(\bar{E}) + \delta E - T \left(\frac{\partial S}{\partial E} \right)_V \delta E - \frac{1}{2} T \left(\frac{\partial^2 S}{\partial E^2} \right)_V \delta E^2$$

$$= \bar{E} - T\bar{S} + \delta E - T \left(\frac{1}{T} \right) \delta E - \frac{1}{2} T \frac{\partial}{\partial E} \left(\frac{1}{T} \right)_V \delta E^2$$

$$\bar{S} \equiv S(\bar{E}) \quad = \bar{E} - T\bar{S} + \frac{1}{2} \frac{1}{T} \left(\frac{\partial T}{\partial E} \right)_V \delta E^2 = \bar{E} - T\bar{S} + \frac{\delta E^2}{2TC_V}$$

$$\Rightarrow e^{-A/k_B T} = e^{-(\bar{E} - T\bar{S})/k_B T} \underbrace{\int \frac{d\delta E}{\Delta} e^{-\delta E^2 / 2k_B T^2 C_V}}$$



integrand strongly peaked about $\delta E = 0$

$$\sqrt{\langle \delta E^2 \rangle} = \sqrt{k_B T^2 C_V} \sim \sqrt{N}$$

$$\text{so } \frac{\sqrt{\langle \delta E^2 \rangle}}{\bar{E}} \sim \frac{\sqrt{N}}{N} \sim \frac{1}{\sqrt{N}}$$

We can do the Gaussian integral $= \int dx e^{-x^2/2\sigma^2} = \sqrt{2\pi\sigma^2}$

$$\Rightarrow e^{-A/k_B T} = e^{-(\bar{E} - T\bar{S})/k_B T} \sqrt{2\pi k_B T^2 C_V}$$

$$\Rightarrow A = \bar{E} - T\bar{S} - \frac{1}{2} k_B T \ln(2\pi k_B T^2 C_V)$$

in the canonical, as compared to the microcanonical ensemble

the relative correction due to fluctuations $\sim \frac{\log N}{N} \rightarrow 0$ as $N \rightarrow \infty$

$\Rightarrow$ canonical ensemble gives same results as microcanonical ensemble in the thermodynamic limit. This is because in the limit $N \rightarrow \infty$, the most probable energy $\bar{E}$ is the same as the average energy $\langle E \rangle$; all other energies have negligible prob

Note:

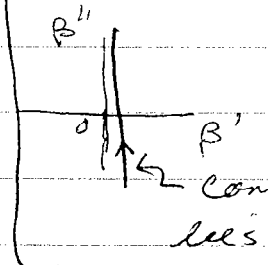
$$Q_N(\beta) = \int \frac{dE}{\Delta} \Omega(E) e^{-\beta E}$$

$Q_N(\beta)$ is Laplace transform of $\frac{\Omega(E)}{\Delta}$

$\Rightarrow \frac{\Omega}{\Delta}$ is inverse Laplace transform of Q_N

$$\begin{aligned} \frac{\Omega(E)}{\Delta} &= \frac{1}{2\pi i} \int_{\beta' - i\infty}^{\beta' + i\infty} e^{\beta E} Q_N(\beta) d\beta \quad (\beta' > 0) \\ &= \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{i(\beta' + i\beta'') E} Q_N(\beta' + i\beta'') d\beta'' \end{aligned}$$

where $\beta' = \text{Re}(\beta) = 0^+$



entropy $S = k_B \ln \Omega$

Helmholtz $-\frac{A}{T} = k_B \ln Q_N$

$$-\frac{A}{T} = S - \frac{E}{T}$$

{ Helmholtz free energy
is Legendre transform of S with
respect to E

Thermodynamic potentials which are Legendre transforms of each other, have ensemble partition functions that are Laplace transforms of each other.

Factorization of canonical partition function - the ideal gas

Consider a system of N noninteracting particles

$$\rightarrow \mathcal{H}[\vec{q}_i, \vec{p}_i] = \sum_{i=1}^N \mathcal{H}^{(i)}(\vec{q}_i, \vec{p}_i)$$

where $\mathcal{H}^{(i)}$ is the single particle Hamiltonian that depends only on the three coordinates $\vec{q}_i$ and three momenta $\vec{p}_i$ of particle i .

$$Q_N = \frac{1}{N! h^{3N}} \left(\prod_{i=1}^N \int d\vec{q}_i d\vec{p}_i \right) e^{-\beta \mathcal{H}}$$

$$= \frac{1}{N!} \left(\prod_{i=1}^N \int \frac{d\vec{q}_i d\vec{p}_i}{h^3} \right) e^{-\beta \sum_j \mathcal{H}^{(j)}(\vec{q}_j, \vec{p}_j)}$$

factor the exponential

$$= \frac{1}{N!} \prod_{i=1}^N \left(\int \frac{d\vec{q}_i d\vec{p}_i}{h^3} e^{-\beta \mathcal{H}^{(i)}(\vec{q}_i, \vec{p}_i)} \right)$$

↑
factor for particle i is
identical to factor for particle j

$$\rightarrow \boxed{Q_N = \frac{1}{N!} (Q_1)^N} \quad \text{for noninteracting particles}$$

where Q_1 is the one particle partition function

$$Q_1 = \int \frac{d\vec{q} d\vec{p}}{h^3} e^{-\beta H^{(1)}(\vec{q}, \vec{p})}$$

Apply to the ideal gas.

$$H^{(1)}(\vec{q}, \vec{p}) = \frac{p^2}{2m}$$

$$Q_1 = \int \frac{d\vec{q}}{h^3} \int d\vec{p} e^{-\beta \frac{p^2}{2m}}$$

$$\int d\vec{q} = V \quad \text{volume of system}$$

$$\int d\vec{p} e^{-\beta \frac{p^2}{2m}} = \left(\frac{2\pi m}{\beta} \right)^{3/2} \quad \text{3D Gaussian integral}$$

$$Q_1 = \frac{V}{h^3} (2\pi m k_B T)^{3/2}$$

$$\Rightarrow Q_N = \frac{1}{N!} \left(\frac{V}{h^3} \right)^N (2\pi m k_B T)^{3N/2}$$

$$A(T, V, N) = -k_B T \ln Q_N$$

$$\ln N! = N \ln N - N$$

using Stirling's formula

$$= -k_B T \left\{ N \ln \left[\frac{V}{h^3} (2\pi m k_B T)^{3/2} \right] - \overbrace{N \ln N + N} \right\}$$

$$A(T, V, N) = -k_B T N - k_B T N \ln \left[\frac{V}{h^3 N} (2\pi m k_B T)^{3/2} \right]$$

Compute average energy

$$\langle E \rangle = -\frac{\partial}{\partial \beta} (\ln Q_N) = -\frac{\partial}{\partial \beta} (-\beta A)$$

$$= -\frac{\partial}{\partial \beta} \left(N + N \ln \left[\frac{V}{h^{3N}} \left(\frac{2\pi m}{\beta} \right)^{3/2} \right] \right)$$

$$= -N \frac{\partial}{\partial \beta} (\ln \beta^{-3/2}) = \frac{3}{2} N \frac{\partial}{\partial \beta} \ln \beta = \frac{3}{2} N \frac{1}{\beta}$$

$$\langle E \rangle = \frac{3}{2} N k_B T \quad \text{as expected}$$

entropy

$$S = -\left(\frac{\partial A}{\partial T} \right)_{V,N} = k_B N + k_B N \ln \left[\frac{V}{h^{3N}} (2\pi m k_B T)^{3/2} \right]$$

$$+ k_B T N \frac{3}{2} \left(\frac{1}{T} \right)$$

↖ from derivative of log

$$S = \frac{5}{2} N k_B + N k_B \ln \left[\frac{V}{h^{3N}} (2\pi m k_B T)^{3/2} \right]$$

substitute in $k_B T = \frac{2}{3} \frac{E}{N}$ to get

$$\Rightarrow S(E, V, N) = \frac{5}{2} N k_B + N k_B \ln \left[\frac{V}{h^{3N}} \left(\frac{4\pi m E}{3N} \right)^{3/2} \right]$$

We have recovered the Sackur-Tetrode equation which we earlier derived from the microcanonical ensemble! Canonical and microcanonical approaches are equivalent.

Because in computing Q_N we sum over all states with any energy, as compared to computing Ω where we restrict the sum to states in a particular energy shell E , it is usually easier to compute Q_N , rather than Ω .

We introduced the canonical distribution as a means of describing a physical system in contact with a heat bath.

The canonical distrib gives the same result as the microcanonical because in the $N \rightarrow \infty$ (thermodynamic) limit, the canonical probability distribution

$$p(E) = \frac{\Omega(E) e^{-E/k_B T}}{\Delta Q_N(V, T)}$$

approaches a delta-function at the most probable energy = average energy, as set by the temperature T .

We could alternatively introduced the canonical ensemble just as a mathematical trick for computing $\Omega(E)$, removing the constraint of constant energy E by means of a Lagrange multiplier.