

Grand Canonical partition function for non-interacting systems

$$\mathcal{Z} = \sum_{N=0}^{\infty} z^N Q_N(T, V)$$

for non-interacting particles we saw

$$Q_N(T, V) = \frac{1}{N!} [Q_1(T, V)]^N \quad \text{indistinguishable particles} \\ \text{(as in ideal gas)}$$

$$= [Q_1(T, V)]^N \quad \text{distinguishable particles} \\ \text{(as in paramagnetic spins)}$$

⇒ Indistinguishable

$$\mathcal{Z} = \sum_{N=0}^{\infty} \frac{(z Q_1)^N}{N!} = e^{z Q_1}$$

Distinguishable

$$\mathcal{Z} = \sum_{N=0}^{\infty} (z Q_1)^N = \frac{1}{1 - z Q_1}$$

← must have $z Q_1 < 1$ for series to converge

Ideal gas

For a single gas of particles

$$Q_1 = \frac{\int d^3p \int d^3r e^{-\beta p^2/2m}}{h^3} = (2\pi m k_B T)^{3/2} \frac{V}{h^3}$$
$$= V f(T)$$

will have this form even for a more complicated gas in which the particles may have internal degrees of freedom.

$$\mathcal{Z} = e^{zQ_1} = e^{zVf(T)} \quad \ln \mathcal{Z} = zVf(T)$$

grand potential $\Sigma = -k_B T \ln \mathcal{Z} = -k_B T z V f(T) = -PV$

$$p = k_B T z f(T)$$

$$z = e^{\beta \mu}$$

$$N = -\frac{\partial \Sigma}{\partial \mu} = -\frac{\partial \Sigma}{\partial z} \frac{\partial z}{\partial \mu} = k_B T V f(T) \beta e^{\beta \mu}$$
$$= z V f(T)$$

Combine the above

$$\frac{p}{k_B T} = z f(T)$$

$$\frac{N}{V} = z f(T)$$

$$\Rightarrow pV = N k_B T$$

Ideal gas law -
independent of what
f is!

$$E = - \left(\frac{\partial \ln \mathcal{Z}}{\partial \beta} \right)_{z, V} = k_B T^2 \left(\frac{\partial \ln \mathcal{Z}}{\partial T} \right)_{z, V}$$

$$= k_B T^2 z V \frac{df}{dT} = k_B T^2 N \frac{(df/dT)}{f} = k_B T^2 N \left(\frac{\partial \ln f}{\partial T} \right)$$

use $N = z V f$

$$C_V = \left(\frac{\partial E}{\partial T} \right)_{N, V} = 2 k_B T N \frac{\partial \ln f}{\partial T} + k_B T^2 N \frac{\partial^2 \ln f}{\partial T^2}$$

Note, for harmonic degrees of freedom (ie \vec{p} , or harmonic internal degrees of freedom, such as internal vibrations of molecule) $f \propto T^n$ for some power n (for single particle, $n = 3/2$)

$$\Rightarrow \frac{\partial \ln f}{\partial T} = \frac{\partial (n \ln T)}{\partial T} = \frac{n}{T}$$

$$\frac{\partial^2 \ln f}{\partial T^2} = -\frac{n}{T^2}$$

and so

$$E = n k_B T N = n p V \Rightarrow \frac{E}{V} = n p$$

$$C_V = 2 n k_B N - k_B T^2 N \left(-\frac{n}{T^2} \right) = n k_B N$$

Helmholtz free energy

$$A = \sum + \mu N = -k_B T z V f(T) + k_B T (\ln z) z V f$$

$$= z V k_B T f(T) [\ln z - 1]$$

$$= N k_B T [\ln z - 1]$$

$$A(T, V, N) = N k_B T \left[\ln \left(\frac{N}{V f(T)} \right) - 1 \right]$$

above agrees with direct result from canonical ensemble:

$$Q_N = \frac{V^N f^N}{N!} \Rightarrow A = -k_B T \ln Q_N = -k_B T \ln \left(\frac{V^N f^N}{N!} \right)$$

$$= -k_B T N \ln V f + k_B T (N \ln N - N)$$

$$= -N k_B T + N k_B T \ln (N/V f)$$

entropy

$$S = - \left(\frac{\partial A}{\partial T} \right)_{V, N} = N k_B \left[\ln \left(\frac{N}{V f(T)} \right) - 1 \right]$$

$$- N k_B T \frac{d(\ln f)}{dT}$$

For distinguishable particles

- corresponds to situation where particles are localized - so we can distinguish them by their spatial location.

Now expect $Q_1 = \phi(T)$ is not proportional to V as the particles are localized.

$$\mathcal{Z} = \frac{1}{1 - zQ_1} = \frac{1}{1 - z\phi(T)} \quad (\text{if } Q_1 \propto V, \text{ then series would not converge!})$$

$$\Sigma = -k_B T \ln \mathcal{Z}$$

$$N = -\frac{\partial z}{\partial \mu} \frac{\partial \Sigma}{\partial z} = -\beta e^{\beta \mu} (-k_B T) \frac{1}{\mathcal{Z}} \frac{\partial \mathcal{Z}}{\partial z}$$

$$= z \frac{(1 - z\phi)(+\phi)}{(1 - z\phi)^2} = \frac{z\phi}{1 - z\phi}$$

$$N = \frac{z\phi}{1 - z\phi} \Rightarrow (1 - z\phi)N = z\phi$$

$$N = z\phi(1 + N)$$

$$z\phi = \frac{N}{1 + N} \approx 1 - \frac{1}{N} \quad \text{for } N \gg 1$$

$$E = -\left(\frac{\partial}{\partial \beta} \ln \mathcal{Z}\right)_{z, V} = k_B T^2 \left(\frac{\partial}{\partial T} \ln \mathcal{Z}\right)_{z, V}$$

$$= k_B T^2 (1 - z\phi) \frac{+z d\phi/dT}{(1 - z\phi)^2}$$

$$E = \frac{k_B T^2 z(d\phi/dT)}{1 - z\phi} \approx k_B T^2 N \frac{(d\phi/dT)}{\phi} = k_B T^2 N \left(\frac{d \ln \phi}{dT}\right)$$

$$A = \Sigma + \mu N = -k_B T \ln \left(\frac{1}{1-z\phi} \right) + k_B T (\ln z) N$$

$$= k_B T \left[\ln(1-z\phi) + \cancel{N} N \ln z \right]$$

use $1-z\phi \approx 1/N$ and $z \approx 1/\phi$
to get

$$A = -k_B T N \ln \phi(T) + o(\ln N)$$

Quantum Ensembles

Classical ensemble was a probability distribution in phase space $\rho(q_i, p_i)$ such that averages were

$$\langle X \rangle = \int \int dq_i dp_i X(q_i, p_i) \rho(q_i, p_i)$$

In quantum mechanics, the density function ρ becomes a density operator or density matrix.

In QM, the states of the ~~state~~ system are given by wave functions $|\psi\rangle$. Suppose we ~~are~~ have a system which we know has probability p_k to be in state $|\psi^k\rangle$. Then the average of some observable would be

$$\langle \hat{X} \rangle = \sum_k p_k \langle \psi^k | \hat{X} | \psi^k \rangle \quad \leftarrow \begin{array}{l} \text{Note this is an} \\ \text{incoherent sum.} \\ \text{Not a coherent superposition} \\ \text{of different states } |\psi^k\rangle \end{array}$$

we define the density ~~matrix~~ as operator as

$$\hat{\rho} = \sum_k |\psi^k\rangle p_k \langle \psi^k|$$

If $\{|n\rangle\}$ are a complete set of basis states (for example the energy eigenstates) then the density matrix is

$$\rho_{nm} = \langle n | \hat{\rho} | m \rangle = \sum_k \langle n | \psi^k \rangle p_k \langle \psi^k | m \rangle$$

Note:

$$\begin{aligned} \rho_{nm}^* &= \sum_k \langle \psi^k | n \rangle p_k \langle m | \psi^k \rangle & p_k \text{ is } \underline{\text{real}} \\ &= \sum_k \langle m | \psi^k \rangle p_k \langle \psi^k | n \rangle = \rho_{mn} \end{aligned}$$

So $\rho_{nm}^* = \rho_{mn} \Rightarrow \hat{\rho}$ is Hermitian, $\hat{\rho} = \hat{\rho}^\dagger$
 $\rightarrow \hat{\rho}$ can be diagonalized

For the average of any observable

$$\begin{aligned} \langle \hat{X} \rangle &= \sum_k p_k \langle \psi^k | \hat{X} | \psi^k \rangle \\ &= \sum_{m,n} p_k \langle \psi^k | n \rangle \langle n | \hat{X} | m \rangle \langle m | \psi^k \rangle \\ &= \sum_{m,n} X_{nm} \rho_{mn} = \text{trace}(\hat{X} \hat{\rho}) \end{aligned}$$

If we take $\hat{X} = \hat{I}$ identity operator, then we get the normalization condition

$$1 = \text{trace} \hat{\rho} = \sum_n \rho_{nn}$$

As for any operator in the Heisenberg picture, its equation of motion is

$$i\hbar \frac{\partial \hat{\rho}}{\partial t} = [\hat{H}, \hat{\rho}]$$

quantum analogue
of Liouville's eqn

⇒ if \hat{p} is to describe a stationary equilibrium, it is necessary that \hat{p} commutes with \hat{H} , $[\hat{H}, \hat{p}] = 0$, so $\partial \hat{p} / \partial t = 0$.

⇒ \hat{p} is diagonal in the basis formed by the energy eigenstates. If these states are $|\alpha\rangle$ then

$$\begin{aligned}\langle \alpha | \hat{H} \hat{p} | \beta \rangle &= E_\alpha \langle \alpha | \hat{p} | \beta \rangle \\ &= \langle \alpha | \hat{p} \hat{H} | \beta \rangle = E_\beta \langle \alpha | \hat{p} | \beta \rangle\end{aligned}$$

$$E_\alpha \langle \alpha | \hat{p} | \beta \rangle = E_\beta \langle \alpha | \hat{p} | \beta \rangle$$

$$\Rightarrow \langle \alpha | \hat{p} | \beta \rangle = 0 \text{ unless } E_\alpha = E_\beta$$

So \hat{p} only couples eigenstates of equal energy (i.e. degenerate states) but since \hat{p} is Hermitian, it is diagonalizable ⇒ we can always take appropriate linear combinations of degenerate eigenstates to make eigenstates of \hat{p} . In this basis \hat{p} is diagonal.

$$\hat{H} |\alpha\rangle = E_\alpha |\alpha\rangle, \quad \hat{p} |\alpha\rangle = p_\alpha |\alpha\rangle$$

$$\langle \alpha | \hat{H} | \beta \rangle = E_\alpha \delta_{\alpha\beta}, \quad \langle \alpha | \hat{p} | \beta \rangle = p_\alpha \delta_{\alpha\beta}$$

$$\delta_{\alpha\beta} = \begin{cases} 1 & \alpha = \beta \\ 0 & \alpha \neq \beta \end{cases} \quad \text{Kronecker delta}$$

Even though a stationary $\hat{\rho}$ is diagonal in the basis of energy eigenstates, we can always express it in terms of any other complete basis states

$$\rho_{nm} = \langle n | \hat{\rho} | m \rangle = \sum_{\alpha\beta} \langle n | \alpha \rangle \langle \alpha | \hat{\rho} | \beta \rangle \langle \beta | m \rangle$$

$$= \sum_{\alpha} \langle n | \alpha \rangle \rho_{\alpha} \langle \alpha | m \rangle$$

in this basis, $\hat{\rho}$ need not be diagonal

This will be useful because we may not know the exact eigenstates for \hat{H} . If $\hat{H} = \hat{H}^0 + \hat{H}^1$ we might know the eigenstates of the simpler \hat{H}^0 , but not the full \hat{H} . In this case it may be convenient to express $\hat{\rho}$ in terms of the eigenstates of \hat{H}^0 and treat \hat{H}^1 in perturbation. In general it is useful to have the above representation for $\hat{\rho}$ and

$\langle \hat{X} \rangle = \text{tr}(\hat{X} \hat{\rho})$ in an operator form that is indep of its representation in any particular basis

Microcanonical ensemble:

$$\hat{\rho} = \sum_{\alpha} |\alpha\rangle \rho_{\alpha} \langle \alpha| \quad \text{with } \rho_{\alpha} = \begin{cases} \text{const} & E \leq E_{\alpha} \leq E + \Delta \\ 0 & \text{otherwise} \end{cases}$$

$$\text{and } \sum_{\alpha} \rho_{\alpha} = 1$$

Canonical ensemble:

$$\hat{\rho} = \sum_{\alpha} |\alpha\rangle \rho_{\alpha} \langle \alpha| \quad \text{with } \rho_{\alpha} = \frac{e^{-\beta E_{\alpha}}}{Q_N}$$

$$\text{where } Q_N = \sum_{\alpha} e^{-\beta E_{\alpha}}$$