

Compare these to what one has Classically

If single particle states are labeled by energy ϵ_i with

$$E = \sum_i n_i \epsilon_i \quad n_i = \# \text{ particles in state } i$$

$$N = \sum_i n_i$$

Then if the particles are distinguishable, then for N particles with n_1 in state 1, n_2 in state 2, etc, the number of microstates corresponding to a given set of occupation numbers $\{n_i\}$ would be

$$\frac{N!}{n_1! n_2! \dots} = \# \text{ ways to distribute } N \text{ particles so that } n_i \text{ are in state } i$$

So we would have

$$Q_N = \sum_{\{n_i\}} \delta(\sum_i n_i - N) \frac{N!}{n_1! n_2! \dots} e^{-\beta \sum_i \epsilon_i n_i}$$

But we now recall Gibbs's correction factor $1/N!$ for indistinguishable particles, to get in this case

$$Q_N = \sum_{\{n_i\}} \delta(\sum_i n_i - N) \frac{1}{n_1! n_2! \dots} e^{-\beta \sum_i \epsilon_i n_i}$$

$$= \sum_{\{n_i\}} \delta(\sum_i n_i - N) \prod_i \left(\frac{1}{n_i!} (e^{-\beta \epsilon_i})^{n_i} \right)$$

Grand canonical

no constraint on $\{n_i\}$

$$\mathcal{Z} = \sum_{N=0}^{\infty} z^N Q_N = \sum_{\{n_i\}} \prod_i \frac{1}{n_i!} (z e^{-\beta \epsilon_i})^{n_i}$$

($z^N = \prod_i z^{n_i}$)

$$= \prod_i \left(\sum_{n=0}^{\infty} \frac{1}{n!} (z e^{-\beta \epsilon_i})^n \right)$$

Classical Gibbs

$$\mathcal{Z} = \prod_i \exp [z e^{-\beta \epsilon_i}] = \prod_i \exp [e^{-\beta(\epsilon_i - \mu)}]$$

$$\frac{PV}{k_B T} = \ln \mathcal{Z} = \sum_i e^{-\beta(\epsilon_i - \mu)}$$

$$= z \sum_i e^{-\beta \epsilon_i} = z Q_1 \leftarrow \begin{array}{l} \text{1 body canonical} \\ \text{partition func} \\ Q_1 = \sum_i e^{-\beta \epsilon_i} \end{array}$$

Note: $\frac{PV}{k_B T} = z Q_1$

also

$$N = z \frac{\partial \ln \mathcal{Z}}{\partial z} = z Q_1$$

$$\Rightarrow \frac{PV}{k_B T} = N$$

$$PV = N k_B T$$

ideal gas law!

Average Occupation Numbers

$$\langle N \rangle = \frac{1}{\beta} \frac{\partial}{\partial \mu} (\ln \mathcal{Z})_{T,V} = z \left(\frac{\partial \ln \mathcal{Z}}{\partial z} \right)_{T,V}$$

$$\langle E \rangle = - \left(\frac{\partial}{\partial \beta} \ln \mathcal{Z} \right)_{z,V}$$

↑ const z , not const μ

$$\ln \mathcal{Z} = \pm \sum_i \ln(1 \pm z e^{-\beta \epsilon_i}) \quad \begin{array}{l} + \text{FD} \\ - \text{BE} \end{array}$$

$$\langle N \rangle = \pm z \sum_i \frac{\pm e^{-\beta \epsilon_i}}{1 \pm z e^{-\beta \epsilon_i}} = \sum_i \frac{z e^{-\beta \epsilon_i}}{1 \pm z e^{-\beta \epsilon_i}}$$

$$\langle N \rangle = \sum_i \left(\frac{1}{\frac{1}{z} e^{\beta \epsilon_i} \pm 1} \right) = \sum_i \left(\frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1} \right)$$

$$\langle E \rangle = \mp \sum_i \frac{\mp z \epsilon_i e^{-\beta \epsilon_i}}{1 \pm z e^{-\beta \epsilon_i}} = \sum_i \frac{z \epsilon_i e^{-\beta \epsilon_i}}{1 \pm z e^{-\beta \epsilon_i}}$$

$$\langle E \rangle = \sum_i \left(\frac{\epsilon_i}{\frac{1}{z} e^{\beta \epsilon_i} \pm 1} \right) = \sum_i \frac{\epsilon_i}{e^{\beta(\epsilon_i - \mu)} \pm 1}$$

Now $N = \sum_i n_i$ so $\langle N \rangle = \sum_i \langle n_i \rangle$

and $E = \sum_i n_i \epsilon_i$ so $\langle E \rangle = \sum_i \epsilon_i \langle n_i \rangle$

Comparing with the above we get

$$\langle n_i \rangle = \frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1} \quad \begin{array}{l} + \text{FD} \\ - \text{BE} \end{array}$$

Classically

$$\ln \mathcal{Z} = \sum_i z e^{-\beta \epsilon_i}$$

$$\langle N \rangle = z \frac{\partial}{\partial z} \left(\sum_i z e^{-\beta \epsilon_i} \right) = z \sum_i e^{-\beta \epsilon_i} = \sum_i z e^{-\beta \epsilon_i}$$

$$= \ln \mathcal{Z} = \frac{PV}{k_B T}$$

again we get the ideal
gas law! $PV = N k_B T$

$$\langle E \rangle = - \frac{\partial}{\partial \beta} \sum_i z e^{-\beta \epsilon_i} = \sum_i \epsilon_i z e^{-\beta \epsilon_i}$$

$$\Rightarrow \langle n_i \rangle = z e^{-\beta \epsilon_i} = e^{-\beta(\epsilon_i - \mu)}$$

Quantum: $\ln Z = \pm \sum_i \ln(1 \pm e^{-\beta(\epsilon_i - \mu)})$

$$= \pm \sum_i \ln(1 \pm z e^{-\beta \epsilon_i})$$

+ FD
- BE

Classical $\ln Z = \sum_i z e^{-\beta \epsilon_i}$

we see that quantum \rightarrow classical in the limit $\boxed{z \ll 1}$
(then $\ln(1 \pm z e^{-\beta \epsilon_i}) \approx \pm z e^{-\beta \epsilon_i}$)

$$z = e^{\beta \mu} \ll 1 \Rightarrow \beta \mu \ll 0$$

~~need~~
 \Rightarrow need negative μ

Occupation numbers

quantum $\langle n_i \rangle = \frac{1}{e^{\beta(\epsilon_i - \mu)} \pm 1}$

+ FD
- BE

classical

$$\langle n_i \rangle = e^{-\beta(\epsilon_i - \mu)}$$

we see that quantum \rightarrow classical for states i such that $e^{\beta(\epsilon_i - \mu)} \gg 1$

$$\epsilon_i - \mu \gg k_B T$$

$$\Rightarrow \beta(\epsilon_i - \mu) \gg 0 \quad \text{or} \quad \epsilon_i - \mu \gg k_B T$$

Note: for bosons we need $(\epsilon_i - \mu) > 0$

so that $\langle n_i \rangle$ always is positive. For free particles where $\epsilon_p = \frac{\hbar^2 k^2}{2m}$ and $\epsilon = 0$ is the smallest energy

$$\text{this} \Rightarrow \mu < 0$$

Classical non interacting particles
phase space approach

$$\mathcal{Z} = \sum_{N=0}^{\infty} z^N Q_N = \sum_{N=0}^{\infty} \frac{(z Q_1)^N}{N!} \quad Q_N = \frac{Q_1^N}{N!}$$

$$= e^{z Q_1} \Rightarrow \boxed{\ln \mathcal{Z} = z Q_1} \quad Q_1 \text{ is single particle partition function}$$

where $Q_1 = \int \frac{d^3 q}{h^3} \int \frac{d^3 p}{h^3} e^{-\beta p^2/2m} = \frac{V}{h^3} (2\pi m k_B T)^{3/2}$

define $\lambda = \left(\frac{h^2}{2\pi m k_B T} \right)^{1/2}$ thermal wavelength

$$\Rightarrow \boxed{Q_1 = \frac{V}{\lambda^3}}$$

occupation number approach

$$\mathcal{Z} = \sum_{\{n_i\}} z^N \prod_i \left[\frac{1}{n_i!} (e^{-\beta \epsilon_i})^{n_i} \right] = \prod_i \left(\sum_{n_i} \frac{(z e^{-\beta \epsilon_i})^{n_i}}{n_i!} \right)$$

$$= \prod_i e^{z e^{-\beta \epsilon_i}}$$

$$\ln \mathcal{Z} = \sum_i z e^{-\beta \epsilon_i} = z \sum_i e^{-\beta \epsilon_i} = \boxed{z Q_1 = \ln \mathcal{Z}}$$

same as in phase space approach

$$Q_1 = \sum_i e^{-\beta \epsilon_i} \quad \text{single particle partition function.}$$

For quantized particles in a box, $\vec{p} = \hbar \vec{k}$, where
 $k_\alpha = \frac{2\pi}{L} n_\alpha \quad \alpha = x, y, z, \quad n_\alpha \text{ integer.}$

$\Rightarrow \Delta k = \frac{2\pi}{L}$ spacing between k vectors

$\Delta p = \frac{2\pi \hbar}{L} = \frac{h}{L}$ where $h = 2\pi \hbar$ is Planck's constant

$$Q_1 = \sum_i e^{-\beta \epsilon_i} = \sum_{\vec{p}} e^{-\beta p^2/2m} = \frac{1}{(\Delta p)^3} \int d^3p e^{-\beta p^2/2m}$$

$$= \left(\frac{L}{h}\right)^3 (2\pi m k_B T)^{3/2} \quad L^3 = V$$

$$= \frac{V}{h^3} (2\pi m k_B T)^{3/2}$$

$$\boxed{Q_1 = \frac{V}{\lambda^3}} \quad \text{where } \lambda = \left(\frac{h^2}{2\pi m k_B T}\right)^{1/2}$$

exact same result as in phase space method, but here we see that the phase space division h , which classically is arbitrary in the phase space method, should be taken as Planck's constant once we quantize the single particle states,

Validity of classical limit

We saw that the quantum partition function Z agreed with classical result in the limit $z \ll 1$.

$$\text{Classically } N = z \left(\frac{\partial \ln Z}{\partial z} \right) = z \frac{\partial (z Q_1)}{\partial z} = z Q_1$$

$$\text{so } z = \frac{N}{Q_1} = \frac{N}{V} \lambda^3 = n \lambda^3$$

where $n = \frac{N}{V}$ is the particle density.

We can define $n = 1/l^3$ where l is the average spacing between particles. Then

$$z = \left(\frac{\lambda}{l}\right)^3$$

and the condition $z \ll 1$ becomes

$$\left(\frac{\lambda}{l}\right)^3 \ll 1 \quad \text{or} \quad l \gg \lambda$$

Classical limit applies when interparticle spacing is very much larger than thermal wavelength.

Agrees with our earlier calculation of $\langle \vec{r}_1 \vec{r}_2 | \hat{\rho} | \vec{r}_1 \vec{r}_2 \rangle$ where we saw that the effect of quantum statistics on spatial correlations was only significant for distances $r_{12} < \lambda$.

Since $\lambda \sim 1/\sqrt{T}$, as T decreases, λ increases, and quantum effects become more important - for a system with fixed density n .

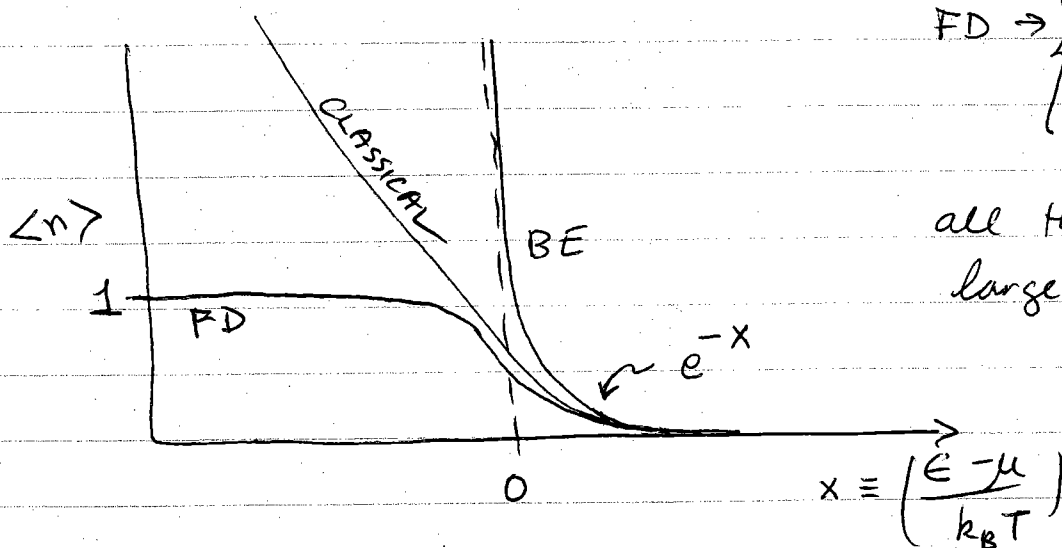
Equivalently, classical results should be OK provided

$$l \gg \lambda \Rightarrow k_B T \gg \frac{h^2}{2\pi m l^2}$$

Classical limit is a ~~large~~ T , or equivalently low density (large l) limit.

BE diverges as $\left(\frac{E-\mu}{k_B T}\right) \rightarrow 0$

$$FD \rightarrow \begin{cases} 1 & \text{for } \left(\frac{E-\mu}{k_B T}\right) \ll 0 \\ 0 & \text{for } \left(\frac{E-\mu}{k_B T}\right) \gg 0 \end{cases}$$



all three equal at large $(E-\mu)/k_B T$

For FD, $\langle n \rangle$ goes from 1 to 0 in an energy width of $O(k_B T)$

Harmonic Oscillator vs boson

Recall for harmonic oscillator $E_n = \hbar\omega(n + 1/2)$

We found

average level excitation \rightarrow

$$\begin{aligned} \langle n \rangle &= \frac{\sum_n e^{-\beta\hbar\omega(n+1/2)} n}{\sum_n e^{-\beta\hbar\omega(n+1/2)}} = \frac{\sum_n e^{-\beta\hbar\omega n} n}{\sum_n e^{-\beta\hbar\omega n}} \\ &= \frac{-1}{\hbar\omega} \frac{\partial}{\partial \beta} \left(\frac{\sum_n e^{-\beta\hbar\omega n}}{\sum_n e^{-\beta\hbar\omega n}} \right) = \frac{-1}{\hbar\omega} \frac{\partial}{\partial \beta} \ln \left[\frac{1}{1 - e^{-\beta\hbar\omega}} \right] \\ &= \frac{1}{\hbar\omega} \frac{\partial}{\partial \beta} \ln (1 - e^{-\beta\hbar\omega}) \\ &= \frac{1}{\hbar\omega} \frac{\hbar\omega e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}} = \frac{1}{e^{\beta\hbar\omega} - 1} \end{aligned}$$

Looks just like boson occupation number with $E = \hbar\omega$ and chemical potential $\mu = 0$.

⇒ quantized harmonic oscillators obey same statistics as bosons, with $\mu = 0$

we say that excitation level n of the oscillator is the same as n quanta or n "particles" of excitation.

Applies to: elastic oscillations of a solid \leftrightarrow "phonons"
oscillation of electromagnetic waves \leftrightarrow "photons"

Sound modes in solid

$$\omega = c_s |\vec{k}| \quad c_s = \text{speed of sound}, \vec{k} = \text{wave vector}$$

$$\Rightarrow \text{phonon modes } \langle n_{\vec{k}} \rangle = \frac{1}{e^{\beta \hbar c_s k} - 1}$$

electromagnetic waves

$$\omega = c |\vec{k}|, \quad c = \text{speed of light}, \vec{k} = \text{wave vector}$$

$$\text{photon modes } \langle n_{\vec{k}} \rangle = \frac{1}{e^{\beta \hbar c k} - 1}$$

Another way to see $\mu = 0$. Phonons and photons are not conserved particles - they can be created and destroyed