

$$\ln \mathcal{Z} = \frac{L_x L_y L_z}{2\pi} g \sum_{n=0}^{\infty} \int dk_z \ln \left[1 + z e^{-\beta \left(\frac{\hbar^2 k_z^2}{2m} + \hbar \omega_c (n + 1/2) \right)} \right]$$

Once we find $\ln \mathcal{Z}$, we can compute M , the total dipole moment, as follows:

Total energy in magnetic field is $E(B) = E(B=0) - M B$

$$\Rightarrow M = - \frac{\partial E}{\partial B} = - \left\langle \frac{\partial H}{\partial B} \right\rangle \quad H \text{ is Hamiltonian}$$

$$\text{Now } - \left\langle \frac{\partial H}{\partial B} \right\rangle = \frac{- \sum_{\alpha} e^{-\beta(H(\alpha) - \mu N_{\alpha})} \frac{\partial H}{\partial B}}{\sum_{\alpha} e^{-\beta(H(\alpha) - \mu N_{\alpha})}}$$

α \uparrow all many particle states α

$H(\alpha)$ is ^{total} energy in state α

$$= \frac{1}{\beta} \frac{\partial}{\partial B} \frac{\sum_{\alpha} e^{-\beta(H(\alpha) - \mu N_{\alpha})}}{\sum_{\alpha} e^{-\beta(H(\alpha) - \mu N_{\alpha})}}$$

$$= \frac{1}{\beta} \frac{\partial}{\partial B} \ln \sum_{\alpha} e^{-\beta(H(\alpha) - \mu N_{\alpha})}$$

$$M = \frac{1}{\beta} \frac{\partial}{\partial B} \ln \mathcal{Z}$$

or using grand potential

$$\Sigma = -k_B T \ln \mathcal{Z}$$

$$\Rightarrow M = - \frac{\partial \Sigma}{\partial B}$$

(Landau + Lifshitz Part I § 59)

$$V = L_x L_y L_z$$

$$\ln Z = \frac{V}{2\pi} g \sum_{n=0}^{\infty} \int_{-\infty}^{\infty} dk_z \ln \left[1 + z e^{-\beta \left(\frac{\hbar^2 k_z^2}{2m} + \hbar \omega_c (n+1/2) \right)} \right]$$

Define function $h(x) = \int_{-\infty}^{\infty} dk_z \ln \left[1 + e^{-\beta \left(\frac{\hbar^2 k_z^2}{2m} - x \right)} \right]$

using $z = e^{-\beta \mu}$

Then

$$\ln Z = \frac{V}{2\pi} g \sum_{n=0}^{\infty} h(\mu - \hbar \omega_c (n+1/2))$$

Consider the limit of very weak magnetic field $\hbar \omega_c \ll k_B T$. In this case many Landau levels occupied. We might think to replace \sum_n by $\int d\mu$, but it turns out that this would remove all dependence on B. To do better we need to use Euler summation formula (Pathria 8-2 Eq (44))

$$\sum_{n=0}^{\infty} f(n+1/2) \approx \int_0^{\infty} f(x) dx + \frac{1}{24} f'(0)$$

Apply to the above

$$\ln Z = \frac{V}{2\pi} g \int_0^{\infty} dx h(\mu - \hbar \omega_c x) + \frac{Vg}{2\pi} \frac{1}{24} (-\hbar \omega_c) \frac{dh(\mu)}{d\mu}$$

$$= \frac{V}{2\pi} \frac{2B}{\Phi_0} \left[\int_{-\infty}^{\mu} dy h(y) \left(\frac{1}{\hbar \omega_c} \right) - \frac{\hbar \omega_c}{24} \frac{dh(\mu)}{d\mu} \right]$$

use $\Phi_0 = \frac{hc}{e}$ $\omega_c = \frac{eB}{mc}$

$$\ln \mathcal{Z} = \frac{V}{2\pi} \frac{z_B}{\Phi_0} \frac{1}{\hbar \omega_c} \left[\int_{-\infty}^{\mu} dy h(y) - \frac{(\hbar \omega_c)^2}{24} \frac{dh(\mu)}{d\mu} \right]$$

$$= \frac{V}{2\pi} \frac{z_B e}{\hbar c} \frac{mc}{\hbar e B} \left[\int_{-\infty}^{\mu} dy h(y) - \frac{(\hbar \omega_c)^2}{24} \frac{dh(\mu)}{d\mu} \right]$$

$$= \frac{Vm}{\hbar^2} \left[\int_{-\infty}^{\mu} dy h(y) - \frac{(\hbar \omega_c)^2}{24} \frac{dh(\mu)}{d\mu} \right]$$

grand potential

$$\Sigma(T, V, M, B) = -k_B T \ln \mathcal{Z} = -\frac{k_B T V m}{\hbar^2} \left[\int_{-\infty}^{\mu} dy h(y) - \frac{(\hbar \omega_c)^2}{24} \frac{dh(\mu)}{d\mu} \right]$$

↑
index of B

1st term gives

$$\Sigma(T, V, M, 0) = -\frac{k_B T V m}{\hbar^2} \int_{-\infty}^{\mu} dy h(y)$$

Now note

$$-N = \left(\frac{\partial \Sigma}{\partial \mu} \right)_{T, V, B=0} = -\frac{k_B T V m}{\hbar^2} h(\mu)$$

$$-\left(\frac{\partial N}{\partial \mu} \right)_{T, V} = \left(\frac{\partial^2 \Sigma}{\partial \mu^2} \right)_{T, V} = -\frac{k_B T V m}{\hbar^2} \frac{dh(\mu)}{d\mu}$$

Combine to get

$$\Sigma(T, V, M, B) = \Sigma(T, V, M, 0) + \frac{(\hbar \omega_c)^2}{24} \left(\frac{\partial N}{\partial \mu} \right)_{T, V}$$

$$\Sigma(T, V, M, B) = \Sigma(T, V, M, 0) + \left(\frac{\hbar e B}{mc}\right)^2 \frac{1}{24} \left(\frac{\partial N}{\partial \mu}\right)_{T, V}$$

$$\mu_B = \frac{e \hbar}{2mc}$$

$$\Sigma(T, V, M, B) = \Sigma(T, V, M, 0) + \frac{1}{6} \mu_B^2 B^2 \left(\frac{\partial N}{\partial \mu}\right)_{T, V}$$

$$\text{Now } \frac{\partial N}{\partial \mu} = \frac{\partial}{\partial \mu} \left\{ V \int d\epsilon g(\epsilon) f(\epsilon, \mu) \right\} \quad f(\epsilon, \mu) = \frac{1}{e^{\beta(\epsilon - \mu)} + 1}$$

$$= V \int d\epsilon g(\epsilon) \frac{\partial f}{\partial \mu}$$

$$= V \int d\epsilon g(\epsilon) \left(-\frac{\partial f}{\partial \epsilon}\right)$$

$$\approx V g(\epsilon_F) \quad \text{to lowest order in Sommerfeld expansion} \\ \text{ie to } o\left(\frac{k_B T}{\epsilon_F}\right)$$

$$\Sigma(T, V, M, B) = \Sigma(T, V, M, 0) + \frac{V}{6} \mu_B^2 g(\epsilon_F) B^2$$

magnetization

$$M = -\frac{\partial \Sigma}{\partial B} = -\frac{V}{3} \mu_B^2 g(\epsilon_F) B$$

magnetic susceptibility

$$\chi_L = \frac{\partial (M/N)}{\partial B} = -\frac{1}{3} \mu_B^2 g(\epsilon_F) \quad < 0 \Rightarrow \text{diamagnetic}$$

$$\text{Compare } \chi_P = \mu_B^2 g(\epsilon_F)$$

$$\chi_L = -\frac{1}{3} \chi_P$$

Total magnetic susceptibility for a free electron gas is

$$\chi_{\text{tot}} = \chi_p + \chi_L = \frac{2}{3} \chi_p$$

For electrons in metal (as opposed to free electrons)

$\chi_p = \mu_B^2 g(E_F)$ comes from interaction with electron spin

$$\mu_B = \frac{\hbar e}{2mc} \quad m \text{ is rest mass of electron}$$

χ_L comes from orbital motion of electrons near fermi energy.

for such electrons the energy spectrum is

$$\epsilon(k) \approx \frac{\hbar^2 k^2}{2m^*} \quad \text{where } m^* \text{ is the effective mass of}$$

motion in the periodic potential of the ions (take P521!)

The μ_B in χ_L should therefore really be $\mu_B^* = \frac{\hbar e}{2m^*c}$

$$\text{then } \chi_L = -\frac{1}{3} \left(\frac{m}{m^*} \right) \chi_p$$

We derived χ_p and χ_L by separately considering effects of spin and orbital motion. One could get the same results by combining the derivations into a single one that includes both effects

$$\text{Note that } \chi_L = -\frac{1}{3} \mu_B^2 g(E_F) \quad g(E_F) = \frac{3}{2} \frac{m}{E_F}$$

$$= -\frac{1}{3} \left(\frac{\hbar e}{2mc} \right)^2 \frac{3}{2} \frac{m}{E_F}$$

Note: Landau diamagnetism is a purely quantum mechanical effect - does not exist classically

Classical N particle partition function:

$$Q_N = \frac{Q_1^N}{N!}$$

where

$$Q_1 = \int \frac{d^3r}{h^3} \int \frac{d^3p}{h^3} e^{-\beta H}$$
$$= \int \frac{d^3r}{h^3} \int \frac{d^3p}{h^3} e^{-\beta \left[\frac{1}{2m} \left(\vec{p} + \frac{e}{c} \vec{A}(\vec{r}) \right)^2 \right]}$$

just substitute $\vec{p}' = \vec{p} + \frac{e}{c} \vec{A}(\vec{r})$ to get

$$Q_1 = \int \frac{d^3r}{h^3} \int \frac{d^3p'}{h^3} e^{-\beta \frac{p'^2}{2m}}$$

same as partition function with $B=0$!

so Q_1 is independent of B

$$\Rightarrow \chi = -\frac{1}{V} \frac{\partial^2 \Sigma}{\partial B^2} = 0$$

$$M = -\frac{\partial \Sigma}{\partial B} = 0$$

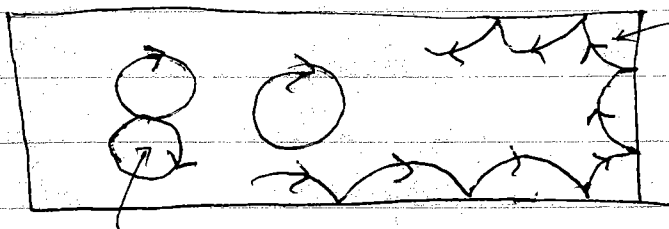
} orbital motion gives
no magnetization
classically

Bohr-Van Leeuwen theorem

Amusing aside:

The classical result $\chi=0$ may seem confusing if one considers that the classical electron in a uniform \vec{B} undergoes a circular motion \Rightarrow electron is effectively a current loop \Rightarrow should have an orbital magnetic moment from classical $\vec{r} \times \vec{j}$ (where \vec{j} is electric current). Each electron goes around in a circular orbit and so the contributions from all electrons should add and give $M \neq 0!$

Argument fails when one considers electrons traveling close to the finite boundaries of the system.



counter clockwise large orbits from electrons hitting the surface "skipping states"

clockwise closed orbits in interior

Moments from the interior orbits and moments from skipping states exactly cancel! Proof: For any fixed $|\vec{p}|$ at any point \vec{r} , we get contributions to current from electrons going in opposite directions. These always cancel,



True even near boundary when we average over all electron orbits the resulting average current at any point \vec{r} in the system vanishes! \Rightarrow no magnetic moment.

Sometimes it is important to consider in detail what happens at the boundaries!