

Density of states $g(\epsilon)$

Often we want to sum a quantity that depends on \vec{k} only through the energy ϵ_k , i.e.

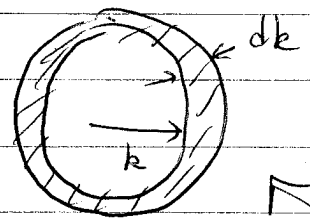
$$\sum_s \sum_{\vec{k}} F(\epsilon_k) = 2 \sum_{\vec{k}} F(\epsilon_k) = \frac{V}{4\pi^3} \int d^3k F(\epsilon_k)$$

\uparrow sum over spin states

Define $g(\epsilon)$ as the number of states per unit energy per unit volume, i.e.

$$\frac{1}{V} \sum_s \sum_{\vec{k}} = \frac{1}{4\pi^3} \int d^3k = g(\epsilon) d\epsilon$$

\vec{k} such that
 $\epsilon \leq \epsilon_k \leq \epsilon + d\epsilon$



Then $\frac{1}{V} \sum_s \sum_{\vec{k}} F(\epsilon_k) = \int d\epsilon g(\epsilon) F(\epsilon)$

When ϵ_k depends on \vec{k} only through $|\vec{k}|$, ~~we~~ can write in \vec{k} -space a spherical shell of width dk and we can write

$$\frac{1}{4\pi^3} \int d^3k = \frac{4\pi}{4\pi^3} \int dk k^2 = \frac{1}{\pi^2} \int dk k^2 = \int g(\epsilon) d\epsilon$$

$$\frac{1}{\pi^2} k^2 dk = g(\epsilon) d\epsilon$$

$$g(\epsilon) = \frac{1}{\pi^2} k^2 \frac{dk}{d\epsilon}$$

For free electrons $\epsilon = \frac{\hbar^2 k^2}{2m}$

$$g(\epsilon) = \frac{1}{\pi^2} \frac{2m\epsilon}{\hbar^2} \frac{1}{(d\epsilon/dk)}$$

$$= \frac{2m\epsilon}{\pi^2 \hbar^2} \frac{1}{\frac{2\hbar^2 k}{2m}} = \frac{2m\epsilon}{\pi^2 \hbar^2} \frac{k}{2\epsilon}$$

$$= \frac{m}{\pi^2 \hbar^2} \sqrt{\frac{2m\epsilon}{\hbar^2}}$$

$$g(\epsilon) = \frac{m}{\pi^2 \hbar^3} \sqrt{2m\epsilon} \propto \sqrt{\epsilon}$$

Total energy

$$E = \int_0^{\epsilon_F} \epsilon g(\epsilon) d\epsilon$$

write $g(\epsilon) = C \sqrt{\epsilon}$ where C is appropriate const

Density of electrons

$$n = \frac{1}{V} \sum_s \sum_{|\mathbf{k}| < k_F} 1 \quad \text{count states}$$

$$= \int_0^{\epsilon_F} d\epsilon g(\epsilon) = C \int_0^{\epsilon_F} d\epsilon \epsilon^{1/2} = \frac{2}{3} C \epsilon_F^{3/2}$$

$$\text{so } C = \frac{3}{2} \frac{n}{\epsilon_F^{3/2}}$$

$$g(\epsilon) = \frac{3}{2} \frac{n}{\epsilon_F} \sqrt{\frac{\epsilon}{\epsilon_F}}$$

Total energy

$$\frac{E}{V} = \int_0^{E_F} dE g(E) E = C \int_0^{E_F} dE E^{3/2} = \frac{2}{5} C E_F^{5/2}$$
$$= \frac{2}{5} \cdot \frac{3}{2} \frac{m}{E_F^{3/2}} E_F^{5/2}$$

$$\boxed{\frac{E}{V} = \frac{3}{5} m E_F} \quad \text{same as we found before}$$

We will see that $g(E)$, and in particular $g(E_F)$ is a very important quantity w/ determining the behavior of conduction electrons in a metal

Pressure of conduction electrons in a metal

From thermodynamics

$$dE = TdS - pdV$$

at $T=0$

$$dE = -pdV$$

$$P = - \left(\frac{dE}{dV} \right)$$

$$\text{Now } E = \frac{3}{5} V m E_F = \frac{3}{5} N E_F$$

$$E_F = \frac{\hbar^2 k_F^2}{2m} \quad \text{where } k_F = \left(3\pi^2 \frac{N}{V} \right)^{1/3}$$

$$\text{so } E_F \propto V^{-2/3} \quad \text{for constant } N$$

$$E = \text{const } V^{-2/3}$$

$$\text{const} = \frac{3}{5} N \frac{\hbar^2}{2m} (3\pi^2 N)^{2/3}$$

$$-\frac{dE}{dV} = \frac{2}{3} \text{const } V^{-5/3} = \frac{2}{3} \text{const } \frac{V^{-2/3}}{V}$$

$$\boxed{p = \frac{2}{3} \frac{E}{V}} = \frac{2}{3} \frac{3}{5} m E_F = \frac{2}{5} m E_F$$

Bulk modulus Compare to a classical ideal gas where

$$p = m k_B T \quad \text{classical}$$

we now have:

$$p = \frac{2}{5} m k_B T_F \quad \text{quantum where } E_F = k_B T_F$$

since $T_F \gg T$ for a metal

$$p_{\text{quantum}} \gg p_{\text{classical}}$$

Bulk modulus $B \equiv 1/\kappa$, $\kappa = \text{compressibility}$

$$B = -V \left(\frac{\partial p}{\partial V} \right)_N \leftarrow \text{derivative is at constant } N$$

$$p = \frac{2}{3} \frac{E}{V} \quad \text{and} \quad E = \text{const } V^{-2/3}$$

$$\Rightarrow p = \text{const } V^{-5/3}$$

$$B = -V \left(\frac{\partial p}{\partial V} \right)_N = +V \frac{5}{3} \text{const } V^{-7/3} = \frac{5}{3} \text{const } V^{-5/3}$$

$$= \frac{5}{3} p = \frac{5}{3} \cdot \frac{2}{3} \frac{E}{V} = \frac{10}{9} \frac{E}{V} = \frac{10}{9} \frac{3}{5} m E_F = \frac{2}{3} m E_F$$

$$\boxed{B = \frac{2}{3} n E_F} \quad \text{at } T=0$$

$$= \left[\frac{6.15}{(r_s/a_0)} \right]^5 \times 10^{10} \text{ dynes/cm}^2$$

<u>metal</u>	<u>B-theory</u>	<u>B-expt</u>	($\times 10^{10}$ dynes/cm ²)
Li	23.9	11.5	
Na	9.25	6.42	
Cu	63.8	134.3	
Ag	34.5	99.9	

Our free electron model gives correct order of magnitude. Note that our calculation gives only the contribution to B from the conduction electrons, whereas the measured B in experiment has also a contribution from the lattice of ions. We see that the electronic contribution is clearly just as important as the ionic contribution.

Free electron gas at finite temperature

Fermi occupation function - at finite temperature T

$$f_i = \frac{1}{e^{(\epsilon_i - \mu)/k_B T} + 1}$$

is the average number of electrons in single electron state "i", where ϵ_i is the energy of this state.

Note: for fermions, since there can only be 0 or 1 electron in any given state, f_i is also equal to the probability that state "i" is occupied.

Here "i" stands for the complete set of quantum numbers needed to specify the single electron state. For a free electron, "i" labels both the wavevector \vec{k} and spin s .

For a derivation of the fermi occupation function see Ashcroft + Mermin pages 40-42, or Kittel Appendix D, or any book on statistical mechanics.

μ is the chemical potential, determined by the condition that

$$\sum_i f_i = N$$

↑
sum over all single electron states

↑
total number of fermions in the system

or, dividing by volume V ,

$$\frac{N}{V} = n = \frac{1}{V} \sum_i f_i \quad \text{determines } \mu \text{ as function of density } n \text{ and temperature } T.$$

n is independent of V in the thermodynamic limit

In terms of the density of states $g(\epsilon)$

$$n = \frac{1}{V} \sum_i f_i = \frac{1}{V} \sum_{\vec{k}} \sum_s f_{\vec{k}s} = \int d\epsilon g(\epsilon) f(\epsilon)$$

since f depends on \vec{k} and s only through $\epsilon_{\vec{k},s}$

for metal, n is fixed constant at all temp T

$$n = \int d\epsilon g(\epsilon) \frac{1}{e^{(\epsilon - \mu)/k_B T} + 1} \quad \leftarrow \text{determines } \mu(T, n)$$

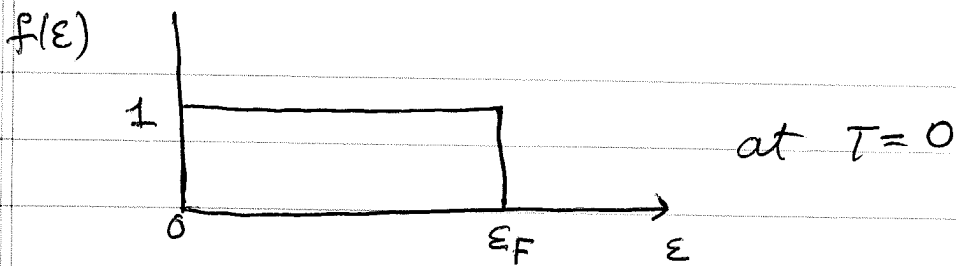
$$\text{As } T \rightarrow 0, \quad f(\epsilon) = \frac{1}{e^{(\epsilon - \mu)/k_B T} + 1} = \begin{cases} 0, & \epsilon > \mu \\ 1, & \epsilon < \mu \end{cases}$$

$$\text{So } n = \int_0^{\mu} d\epsilon g(\epsilon)$$

But earlier we had

$$n = \int_0^{\epsilon_F} d\epsilon g(\epsilon)$$

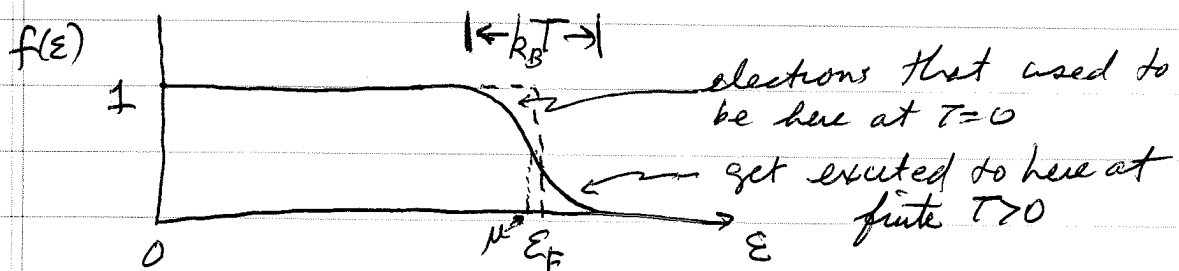
So as $T \rightarrow 0$, $\mu \rightarrow \epsilon_F$ the Fermi energy.



At finite temperature, electrons can absorb thermal energy $k_B T$ from the heat reservoir that keeps the system at temperature T (we are thinking of the canonical ensemble)

But for an electron with $\epsilon \ll \epsilon_F - k_B T$, absorbing $k_B T$ energy would just take it to a state that is already filled. Pauli exclusion principle prevents this from happening.

\Rightarrow The only electrons which can change their state when the system is at finite temperature are those within $\sim k_B T$ of ϵ_F . Only these states can go from a filled to an empty state by absorbing $k_B T$ energy from the heat bath



when $\epsilon = \mu$, $f(\epsilon) = 1/2$

Since $g(\epsilon) \sim \sqrt{\epsilon}$, there are more empty states

within $k_B T$ above E_F than there are within $k_B T$ below E_F .

Since $f(\mu) = 1/2$ is the dividing point ^{in energy} between states which are more likely filled and those that are more likely empty (since $f(\epsilon)$ is also the probability the state is occupied) we see that at finite temperature $\mu(T, n)$ must decrease somewhat from its value E_F at $T=0$.

A more precise calculation (see the "Sommerfeld expansion" in Ashcroft + Mermin pgs 45-47) gives the result (Ashcroft + Mermin Eq (2.78))

$$\mu(T, n) = E_F \left[1 - \frac{1}{3} \left(\frac{\pi k_B T}{2 E_F} \right)^2 \right]$$

↑ gives dependence on T
dependence of μ on n is via E_F

Since $T_F \sim 10^4$ K for most metals, the correction to the chemical potential at finite temperature are

$$|\mu(T) - E_F| \sim \left(\frac{T}{T_F} \right)^2 \sim 10^{-4} \text{ at room temperature}$$

So for most purposes it will be OK to approx $\mu(T) \approx E_F$ even at finite temperature.

Derivation of $\mu(T)$ for $T \ll T_F$

density of electrons is $n = \int_0^{\infty} d\varepsilon g(\varepsilon) f(\varepsilon)$

density of electrons remains fixed as temperature varies

$$\Rightarrow \frac{\partial n}{\partial T} = 0 = \int_0^{\infty} d\varepsilon g(\varepsilon) \frac{\partial f}{\partial T}$$

fermi function $f(\varepsilon) = \frac{1}{e^{x+1}}$ where $x = \frac{\varepsilon - \mu(T)}{k_B T}$

$$\text{so } \frac{\partial f}{\partial T} = \frac{\partial f}{\partial x} \frac{\partial x}{\partial T} = \left(-\frac{\partial f}{\partial x} \right) \left[\frac{x}{T} + \frac{1}{k_B T} \frac{d\mu}{dT} \right]$$

$$d\varepsilon = k_B T dx$$

$$\Rightarrow 0 = \int_{-\frac{\mu}{k_B T}}^{\infty} dx g(\varepsilon) \left(-\frac{\partial f}{\partial x} \right) \left[\frac{x}{T} + \frac{1}{k_B T} \frac{d\mu}{dT} \right]$$

since $\left(-\frac{\partial f}{\partial x} \right)$ is strongly peaked about $\varepsilon = \mu$ and vanishes for $\varepsilon \ll \mu$ or $\varepsilon \gg \mu$,

we can ① extend lower limit of integration to $-\infty$

② approximate $g(\varepsilon) \approx g(\mu) + g'(\mu)(\varepsilon - \mu)$

$$= g(\mu) + g'(\mu) k_B T x$$

then

$$0 = \int_{-\infty}^{\infty} dx \left[g(\mu) + g'(\mu) k_B T x \right] \left[\frac{x}{T} + \frac{1}{k_B T} \frac{d\mu}{dT} \right] \left(-\frac{\partial f}{\partial x} \right)$$

$$\text{Now } \left(-\frac{\partial f}{\partial x} \right) = \frac{e^x}{(e^x + 1)^2}$$

one can check that $\left(-\frac{\partial f}{\partial x} \right)$ is symmetric function of x

$$\text{Therefore } \int_{-\infty}^{\infty} dx \left(-\frac{\partial f}{\partial x} \right) x = 0$$

So we get

$$0 = \frac{g(\mu)}{k_B T} \frac{d\mu}{dT} \int_{-\infty}^{\infty} dx \left(-\frac{\partial f}{\partial x} \right) + g'(\mu) k_B \int_{-\infty}^{\infty} dx \left(-\frac{\partial f}{\partial x} \right) x^2$$

$$\text{where } \int_{-\infty}^{\infty} dx \left(-\frac{\partial f}{\partial x} \right) = f(-\infty) - f(\infty) = 1 - 0 = 1$$

$$\int_{-\infty}^{\infty} dx \left(-\frac{\partial f}{\partial x} \right) x^2 = \int_{-\infty}^{\infty} dx \frac{x^2 e^x}{(e^x + 1)^2} = \frac{\pi^2}{3}$$

$$0 = \frac{g(\mu)}{k_B T} \frac{d\mu}{dT} + \frac{\pi^2}{3} g'(\mu) k_B$$

$$\Rightarrow \frac{d\mu}{dT} = - \frac{\pi^2}{3} \frac{g'(\mu)}{g(\mu)} k_B T$$

$$\text{expect } \frac{g'(\mu)}{g(\mu)} \sim \frac{1}{\mu} \sim \frac{1}{E_F} \text{ at low } T$$

$$\text{so } \frac{d\mu}{dT} \sim \frac{k_B T}{E_F} \sim \left(\frac{T}{T_F} \right) \sim 10^{-2} \text{ is small for all metals}$$

Therefore can approximate

$$\frac{d\mu}{dT} \approx -\frac{\pi^2}{3} \frac{g'(\epsilon_F)}{g(\epsilon_F)} k_B^2 T$$

↑ replaced μ by ϵ_F in argument of g and g' . Corrections to this just give terms higher order in T/T_F

$$\mu(T) = \epsilon_F - \frac{\pi^2}{6} \frac{g'(\epsilon_F)}{g(\epsilon_F)} (k_B T)^2$$

$$\mu(T) = \epsilon_F \left[1 - \frac{\pi^2}{6} \frac{g'(\epsilon_F)}{g(\epsilon_F)} \frac{(k_B T)^2}{\epsilon_F} \right]$$

For free electron gas, $g(\epsilon) = \frac{3}{2} \frac{m}{\epsilon_F} \sqrt{\frac{\epsilon}{\epsilon_F}}$, $g(\epsilon_F) = \frac{3}{2} \frac{m}{\epsilon_F}$

$$g'(\epsilon) = \frac{3}{4} \frac{m}{\epsilon_F} \sqrt{\frac{\epsilon}{\epsilon_F}} \frac{1}{\epsilon}, \quad g'(\epsilon_F) = \frac{3}{4} \frac{m}{\epsilon_F^2}$$

$$\mu(T) = \epsilon_F \left[1 - \frac{\pi^2}{6} \frac{\left(\frac{3}{4} \frac{m}{\epsilon_F^2}\right)}{\left(\frac{3}{2} \frac{m}{\epsilon_F}\right)} \frac{(k_B T)^2}{\epsilon_F} \right]$$

$$\mu(T) = \epsilon_F \left[1 - \frac{\pi^2}{12} \left(\frac{k_B T}{\epsilon_F}\right)^2 \right]$$

so $\frac{\mu(T) - \epsilon_F}{\epsilon_F} \sim 0 \left(\frac{k_B T}{\epsilon_F}\right)^2 \sim 10^{-4}$

⇒ In most situations can ignore T dependence of μ and take $\mu(T) = \epsilon_F$