

Microcanonical Ensemble and Entropy

We saw that the microcanonical ensemble, at energy E , assigned equal weight to all systems on the surface in phase space of constant energy $H[q_i, p_i] = E$.

To count the number of such states on the energy surface we define the "density of states"

$$g(E) \equiv \frac{\int dq_i dp_i \delta(H[q_i, p_i] - E)}{h^{3N}}$$

(number of states per unit energy)

where h is a constant with units of $q_i p_i$. h^{3N} represents the volume of phase space occupied by one "state". Classically, h is totally arbitrary so our thermodynamic results should not depend on it. Quantum mechanically, we will see that h turns out to be Planck's constant.

At this stage, the factor $\frac{1}{h^{3N}}$ is introduced so that $g(E)$ has the units of $1/\text{energy}$.

We can now define the number of states in a shell of thickness Δ about the energy surface E .

$$\Omega(E, V, N) = \int_{E-\frac{\Delta}{2}}^{E+\frac{\Delta}{2}} dE' g(E')$$

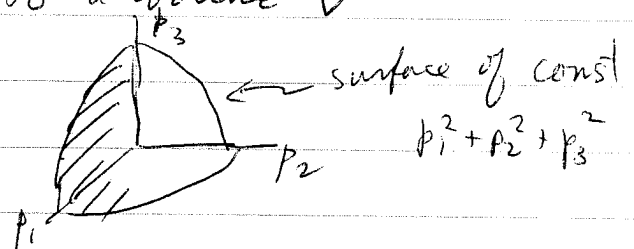
Ω is a pure number. Again, Δ is arbitrary, but

$$\frac{E}{N} \ll \Delta \ll E$$

assumed to be ~~small~~. It can be thought of as representing the finite accuracy with which one knows E . Our thermodynamic results should not depend on Δ . [both $\hbar$ and Δ are introduced so that ~~the~~ dimension of Ω is a dimensionless pure number that we can think of as being the ~~real~~ number of microscopic states occupied in the microcanonical ensemble at E]

Example: Compute Ω for the ideal gas of non interacting particles, confined to a volume V

$$H = \sum_{i=1}^N \frac{|\vec{p}_i|^2}{2m}$$



$$\begin{aligned} g(E) &= \frac{1}{h^{3N}} \prod_{i=1}^N \int d^3 \vec{p}_i \delta \left(\sum_i \frac{|\vec{p}_i|^2}{2m} - E \right) \\ &= \frac{V^N}{h^{3N}} \prod_{i=1}^N \int d^3 \vec{p}_i \delta \left(\sum_i \frac{|\vec{p}_i|^2}{2m} - E \right) \end{aligned}$$

The surface of constant energy is just the surface of a sphere in $3N$ dimensional momentum space given by the coords $p_{1x}, p_{1y}, p_{1z}, \dots, p_{Nx}, p_{Ny}, p_{Nz}$. The radius of the sphere is $\sqrt{2mE}$.

Let $P \equiv \sqrt{\sum_{i=1}^N |\vec{p}_i|^2}$ be the length of the momentum vector in the $3N$ dimensional momentum space.

Then $\prod_{i=1}^N d^3 p_i = dP P^{3N-1} d\Omega_{3N}$

↑ differential solid angle
in $3N$ dimensional space

$$g(E) = \frac{V^N}{h^{3N}} \int d\Omega_{3N} \int_0^\infty dP P^{3N-1} \delta\left(\frac{P^2}{2m} - E\right)$$

$$= \frac{V^N}{h^{3N}} S_{3N} \int_0^\infty dP P^{3N-1} \frac{\delta(P - \sqrt{2mE})}{(P/m)}$$

↑
area of unit
sphere in $3N$ -dim space

← from
converting the
 δ -function

$$\delta(f(x)) = \frac{\delta(x)}{f'(x)}$$

$$= \frac{V^N}{h^{3N}} S_{3N} m (2mE)^{\frac{3N-2}{2}}$$

From Appendix C of Pathria (eqn C.76) or elsewhere,
one has the area of unit sphere in d -dim space

$$S_d = \frac{2\pi^{d/2}}{\Gamma(d/2)}$$

where $\Gamma(n) = (n-1)!$

for integer n

Γ is the Gamma function

So $S_{3N} = \frac{2\pi^{3N/2}}{(\frac{3N}{2}-1)!}$

$$g(E) = \frac{V^N}{h^{3N}} \frac{2\pi^{3N/2}}{(\frac{3N}{2}-1)!} m \frac{(2mE)^{\frac{3N}{2}}}{2mE}$$

$$g(E) = \frac{V^N}{h^{3N}} \frac{(2\pi m E)^{3N/2}}{(\frac{3N}{2} - 1)!} \frac{1}{E}$$

$$\Omega(E) = \int_{E-\frac{\Delta}{2}}^{E+\frac{\Delta}{2}} dE' g(E') \approx g(E) \Delta$$

$$\Omega(E) = \frac{V^N}{h^{3N}} \frac{(2\pi m E)^{3N/2}}{(\frac{3N}{2} - 1)!} \frac{\Delta}{E}$$

For large N , $\Omega(E)$ is a very rapidly increasing function of E ! $\sim E^{\frac{3N}{2}-1}$

We will now argue that $\Omega(E)$ is related to the entropy of the system.

Consider two sub-systems separated by a wall

E_1	E_2
V_1	V_2
N_1	N_2

$$E_T = E_1 + E_2 \quad \text{energy conserved}$$

let $g_1(E_1)$ is density of states of system 1 with energy E_1
 $g_2(E_2)$ is density of states of system 2 with energy E_2

Now suppose the wall is thermally conducting so that energy can be transferred between the two systems. $\Rightarrow E_1$ can vary but $E_T = E_1 + E_2$ is fixed. What will be the value of E_1 when the system comes to equilibrium?

The density of states of the combined system will be

$$g_T(E_T) = \int_0^{E_T} dE_1 g_1(E_1) g_2(E_T - E_1)$$

$$\text{Now if } \Omega_T(E_T) \approx g_T(E_T) \Delta$$

$$\Omega_1(E_1) = g_1(E_1) \Delta$$

$$\Omega_2(E_2) = g_2(E_2) \Delta$$

Then the above can be written as

$$\Omega_T(E_T) = \int \frac{dE_1}{\Delta} \Omega_1(E_1) \Omega_2(E_T - E_1)$$

The integrand is the number of states with total energy E_T that also have system 1 with energy E_1 .

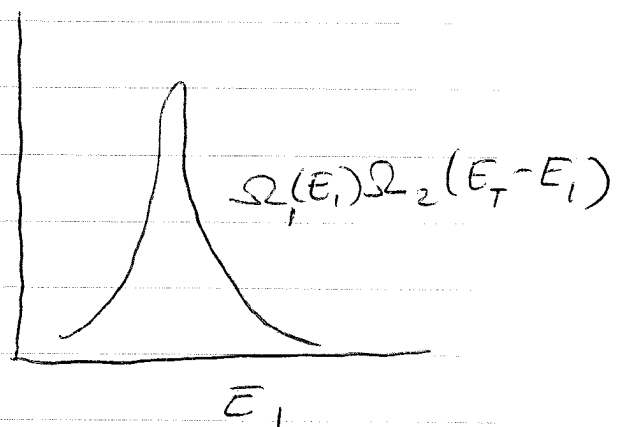
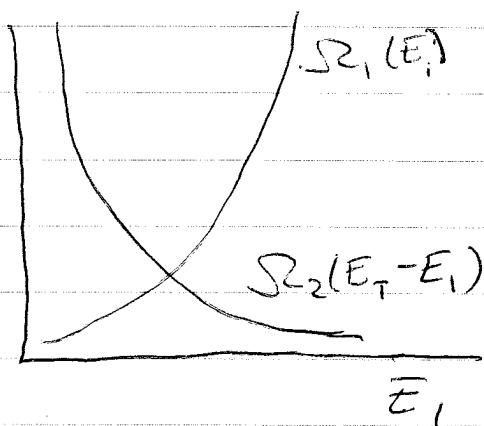
~~Then~~ Consider the behavior of the integrand

$\Omega_1(E_1)$ is a rapidly increasing function of E_1

$\Omega_2(E_2)$ is a rapidly increasing function of E_2

$\Rightarrow \Omega_2(E_T - E_1)$ is a rapidly decreasing function of E_1

$\Rightarrow$ the product $\Omega_1(E_1) \Omega_2(E_T - E_1)$ has a sharp maximum at some particular value of E_1



In the microcanonical ensemble, all states with total energy E_T are equally likely. But the value of E_1 that one is most likely to find as the energy of system 1 is the particular value that maximizes $\Omega_1(E_1)\Omega_2(E_T-E_1)$. That is, there are more states with this particular value of E_1 than with any other value of E_1 , and since all microscopic states are equally likely, this value of E_1 is the most likely. Moreover, since $\Omega_1(E_1)$ is rapidly increasing in E_1 and $\Omega_2(E_T-E_1)$ is rapidly decreasing in E_1 , the maximum is extremely sharp. So one is almost certain to find the maximizing value of E_1 (the probability to find any other value of E_1 will vanish as the size of the systems get infinitely large).

What condition determines this maximizing value of E_1 ?

$$\frac{\partial}{\partial E_1} [\Omega_1(E_1)\Omega_2(E_T-E_1)] = 0$$

$$\Rightarrow \left(\frac{\partial \Omega_1(E_1)}{\partial E_1} \right) \Omega_2(E_T-E_1) + \Omega_1(E_1) \left(\frac{\partial \Omega_2(E_T-E_1)}{\partial E_1} \right) = 0$$

$$\left(\frac{\partial \Omega_1(E_1)}{\partial E_1} \right) \Omega_2(E_T-E_1) - \Omega_1(E_1) \left(\frac{\partial \Omega_2(E_T-E_1)}{\partial E_2} \right) = 0$$

$$\Rightarrow \frac{1}{\Omega_1} \frac{\partial \Omega_1(E_1)}{\partial E_1} = \frac{1}{\Omega_2} \frac{\partial \Omega_2(E_T-E_1)}{\partial E_2}$$

$$\Rightarrow \frac{\partial}{\partial E_1} (\ln \Omega_1) = \frac{\partial}{\partial E_2} (\ln \Omega_2)$$

But from thermodynamics we know that ^{the} equilibrium value of E_1 will be determined by the condition

$$\frac{1}{T_1} = \frac{\partial S_1}{\partial E_1} = \frac{\partial S_2}{\partial E_2} = \frac{1}{T_2}$$

Therefore, following Boltzmann, we identify

$$S(E) \propto \ln \Omega(E)$$

as the entropy.

Since the relation between thermodynamics + mechanics should be fundamental, Boltzmann postulated that the proportionality constant in the above should be a universal number, and not depend on the particular system. This constant is Boltzmann's constant k_B .

$$\boxed{S(E) = k_B \ln \Omega(E)}$$

$\uparrow$ entropy $\uparrow$ # states with energy E

$S(E)$ is a monotonic increasing function of E as it should be.

Ideal gas

$$\text{we have } \Omega(E, V, N) = \frac{V^N (2\pi m E)^{3N/2}}{h^{3N} (\frac{3N}{2} - 1)!} \frac{\Delta}{E}$$

for large N we use Stirling's formula $\ln N! \approx N \ln N - N$

$$S(E, V, N) = k_B \ln \Omega(E, V, N)$$

$$= k_B \left\{ N \ln \left[\frac{V (2\pi m E)^{3/2}}{h^3} \right] - \left(\frac{3N}{2} - 1 \right) \ln \left(\frac{3N}{2} - 1 \right) + \left(\frac{3N}{2} - 1 \right) + \ln \frac{\Delta}{E} \right\}$$

$$\begin{aligned} \text{use } \ln \left(\frac{3N}{2} - 1 \right) &\approx \ln \frac{3N}{2} \left(1 - \frac{2}{3N} \right) \\ &= \ln \frac{3N}{2} + \ln \left(1 - \frac{2}{3N} \right) \\ &\approx \ln \frac{3N}{2} - \frac{2}{3N} \quad \text{expanding the log} \end{aligned}$$

$$S \approx k_B \left\{ N \ln \left[\frac{V (2\pi m E)^{3/2}}{h^3} \right] - \frac{3N}{2} \ln \frac{3N}{2} + \frac{3N}{2} \left(\frac{2}{3N} \right) + \ln \frac{3N}{2} - \frac{2}{3N} + \frac{3N}{2} - 1 + \ln \frac{\Delta}{E} \right\}$$

$$= k_B \left\{ N \ln \left[\frac{V (2\pi m E)^{3/2}}{h^3 \left(\frac{3N}{2} \right)} \right] + \frac{3N}{2} + \ln \frac{3N}{2} + O\left(\frac{1}{N}\right) + \ln \frac{\Delta}{E} \right\}$$

as $N \rightarrow \infty$ leading terms are

$$S(E, V, N) \approx N k_B \left\{ \frac{3}{2} + \ln \left[\frac{V}{h^3} \left(\frac{4\pi m E}{3N} \right)^{3/2} \right] \right\}$$

(since the terms $\ln \frac{3N}{2} + \ln \frac{\Delta}{E}$ are negligible)

Note: since we took Δ so that

$$\frac{E}{N} < \Delta \ll E$$

Then
$$-\ln N < \ln \frac{\Delta}{E} \ll 0$$

$$\approx |\ln \frac{\Delta}{E}| < \ln N$$

is of order $\ln N$
and so can be ignored
compared to terms of order N

$$S(E, V, N) = N k_B \left\{ \frac{3}{2} + \ln \left[\frac{V}{h^3} \left(\frac{4\pi m E}{3N} \right)^{3/2} \right] \right\}$$

note, our result does not depend on Δ ,
as we desired.

with the above, we recover the expected

$$\frac{1}{T} = \left(\frac{\partial S}{\partial E} \right)_{V,N} = \frac{\partial}{\partial E} \left(N k_B \frac{3}{2} \ln E \right) = \frac{3}{2} N k_B \frac{1}{E}$$

$$\Rightarrow E = \frac{3}{2} N k_B T$$

$$\frac{p}{T} = \left(\frac{\partial S}{\partial V} \right)_{E,N} = \frac{\partial}{\partial V} (N k_B \ln V) = N k_B \frac{1}{V}$$

$$\Rightarrow pV = N k_B T$$

so far so good!

But there is a problem - S above is not extensive.
If we take $E \rightarrow 2E$, $V \rightarrow 2V$, $N \rightarrow 2N$, we do not get $S \rightarrow 2S$.

$$(1) \quad S(E, V, N) = \frac{3}{2} k_B N + k_B N \ln \left[\frac{V}{h^3} \left(\frac{4}{3} \pi m \frac{E}{N} \right)^{3/2} \right]$$

the $\ln V$ term in above spoils the desired extensivity
compare the above to our earlier result for the
ideal gas, obtained from combining $pV = N k_B T$ and $E = \frac{3}{2} N k_B T$
with the Gibbs-Duhem relation

$$(2) \quad S(E, V, N) = \frac{N}{N_0} S_0 + k_B N \ln \left[\left(\frac{V}{V_0} \right) \left(\frac{E}{E_0} \right)^{3/2} \left(\frac{N}{N_0} \right)^{-5/2} \right]$$

↖ This version is extensive - it scales proportionate to N . here
 V_0, E_0, N_0 constants. > we have an extra factor N^{-1} in the log

Note: The Gibbs-Duhem relation was derived assuming S was extensive. Hence it should not be surprising that our expression (2) for S is extensive.

What is the physical reason why the expression (1) fails to be extensive?