

# Request Summary

## Frank's A Study of Colliding Flows and Feedback in Star Formation

Date submitted: Apr 15, 2014

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### Title / FOS

<b>TITLE</b>	A Study of Colliding Flows and Feedback in Star Formation
<b>REQUEST NUMBER</b>	AST130036
<b>REQUEST TYPE</b>	Renewal
<b>PRIMARY FIELD OF SCIENCE</b>	123 - Stellar Astronomy and Astrophysics
<b>SECONDARY FIELD(S) OF SCIENCE</b>	122 - Planetary Astronomy
<b>SCIENCE</b>	<b>KEYWORDS</b>
Star Formation	Colliding Flows
	Molecular Clouds
	Feedback

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## Supporting Grant(s) Information

<b>PI NAME</b>	Adam Frank
<b>FUNDING AGENCY</b>	National Science Foundation (NSF)
<b>FUNDING AGENCY DIVISION</b>	
<b>PROGRAM OFFICER NAME</b>	Katharina Lodders
<b>PROGRAM OFFICER EMAIL</b>	klodders@nsf.gov
<b>FUNDING TITLE</b>	From Central Engine to Bipolar Outflow: Binaries, MHD and the Evolution of Planetary Nebulae
<b>AWARD NUMBER</b>	406,000.00
<b>AWARDED AMOUNT</b>	135333
<b>PERCENTAGE OF AWARD</b>	50
<b>SUPPORTING THIS REQUEST</b>	
<b>START DATE</b>	09/01/2011
<b>EXPIRATION DATE</b>	08/31/2014
<b>FIELD OF SCIENCE</b>	Planetary Astronomy
<b>COMMENT</b>	

<b>PI NAME</b>	Adam Frank
<b>FUNDING AGENCY</b>	Dept of Energy (DoE and labs)
<b>FUNDING AGENCY DIVISION</b>	
<b>PROGRAM OFFICER NAME</b>	Sean Finnegan
<b>PROGRAM OFFICER EMAIL</b>	sean.finnegan@science.doe.gov
<b>FUNDING TITLE</b>	Resolving the Issue: The Dynamics of Magnetized Astrophysical Jets through Pulsed Powered HEDP Laboratory Studies
<b>AWARD NUMBER</b>	1,725,000.00
<b>AWARDED AMOUNT</b>	575000
<b>PERCENTAGE OF AWARD</b>	50
<b>SUPPORTING THIS REQUEST</b>	
<b>START DATE</b>	08/15/2012
<b>EXPIRATION DATE</b>	08/14/2015

FIELD OF SCIENCE	Stellar Astronomy and Astrophysics
COMMENT	
PI NAME	Adam Frank
FUNDING AGENCY	Other
FUNDING AGENCY DIVISION	Space Telescope Sci Institute
PROGRAM OFFICER NAME	Paula Sessa
PROGRAM OFFICER EMAIL	sessa@stsci.edu
FUNDING TITLE	STSci - Hubble Telescope - The Reel Deal: Interpreting HST Multi-Epoch Movies of YSO JetsSpace
AWARD NUMBER	HST -AR-12128.01-A
AWARDED AMOUNT	95000
PERCENTAGE OF AWARD	60
SUPPORTING THIS REQUEST	
START DATE	10/01/2010
EXPIRATION DATE	09/30/2014
FIELD OF SCIENCE	Stellar Astronomy and Astrophysics
COMMENT	
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## Resources Requested

Please estimate what percentage of the work you expect to do in this allocation will be the following types (the 3 numbers should sum to 100):

- Production (actually doing research): 99
- Exploration/porting (preparing to do research): 1

Please estimate what percentage of the jobs you expect to run in this allocation will be the following types (the 3 numbers should sum to 100):

- Submitted through command line/script: 100

Please estimate what percentage of the science runs you expect to perform in this allocation will be the following types (the 4 numbers should sum to 100):

- Independent but related (such as jobs that make up an ensemble or parameter sweeps): 100

RESOURCE NAME	SDSC Appro with Intel Sandy Bridge Cluster (Gordon Compute Cluster)
RESOURCE REQUESTED AMOUNT	2500000
RESOURCE AWARDED AMOUNT	

RESOURCE NAME	SDSC Medium-term disk storage (Data Oasis)
RESOURCE REQUESTED AMOUNT	3000
RESOURCE AWARDED AMOUNT	

RESOURCE NAME	TACC Dell PowerEdge C8220 Cluster with Intel Xeon Phi coprocessors (Stampede)
RESOURCE REQUESTED AMOUNT	5000000
RESOURCE AWARDED AMOUNT	

RESOURCE NAME	TACC Long-term tape Archival Storage (Ranch)
RESOURCE REQUESTED	12000
AMOUNT	RESOURCE AWARDED AMOUNT

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## Abstract

Observational evidence from local star-forming regions mandates that star formation occurs shortly after, or even during, molecular cloud formation. Models of the formation of molecular clouds in large-scale colliding flows have identified the physical mechanisms driving the necessary rapid fragmentation. They also point to global gravitational collapse driving supersonic turbulence in molecular clouds. In this work we propose exploring the effect of {\it magnetic fields} and {\it shear} in the colliding flow on the resulting clouds and the ensuing gravitational collapse. We also explore the role of triggering in star formation, namely the ability of supersonic flows to drive stable pre-existing clouds into collapse. We are particularly interested in the formation of (planet forming) accretion disks in the aftermath of triggering and the role of fields in inhibiting or altering collapse. Three new publications came from results obtained with our previous XSEDE allocation (AST130036) and we now seek resources to continue and expand our work. To achieve this goal we request the support of 5.0 million SU's on Stampede at TACC, 2.5 million SU's on Gordon at SDSC.

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