Today in Physics 218: forces in relativity

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forces

X-ray image of the pulsar-driven center of the Crab Nebula in Taurus, the remnant of the supernova of 1054 AD (Chandra Xray observatory image, NASA and Center for Astrophysics).



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Newton's laws in relativity

We will be interested in learning how to solve force problems in relativity, because force is ultimately how we relate to the fields. Can we still use Newton's laws? (Are those among the laws of physics that are valid in inertial reference frames?)

- □ **First law**: inertia. ("A body in uniform motion will remain…") Clearly this is still true in relativity; otherwise we would have had more trouble last time with momentum and energy conservation.
- □ Second law: *F* = *ma*. This is still true if one takes *m* to be the rest mass, and expresses it as

$$F = m \frac{dv}{dt} = \frac{dp}{dt}$$
 ,

Newton's laws in relativity (continued)

and uses the relativistic momentum thenceforth:

$$p = \frac{mu}{\sqrt{1 - u^2/c^2}}$$

The easiest way to demonstrate this is to note that mechanical work still increases mechanical energy in relativity, just as it always has:

$$W = \int \mathbf{F} \cdot d\ell = \int \frac{d\mathbf{p}}{dt} \cdot d\ell = \int \frac{d\mathbf{p}}{dt} \cdot \frac{d\ell}{dt} dt = \int \frac{d\mathbf{p}}{dt} \cdot \mathbf{u} dt$$
$$= \int \frac{d}{dt} \left(\frac{m\mathbf{u}}{\sqrt{1 - u^2/c^2}} \right) \cdot \mathbf{u} dt$$

Newton's laws in relativity (continued)

$$W = \int \left(\frac{mu}{\left(1 - u^{2}/c^{2}\right)^{3/2}} \frac{u}{c^{2}} \cdot \frac{du}{dt} + \frac{m}{\sqrt{1 - u^{2}/c^{2}}} \frac{du}{dt} \right) \cdot udt$$
$$= \int \left(\frac{mu^{2}/c^{2}}{\left(1 - u^{2}/c^{2}\right)^{3/2}} u + \frac{m\left(1 - u^{2}/c^{2}\right)}{\left(1 - u^{2}/c^{2}\right)^{3/2}} u \right) \cdot \frac{du}{dt} dt$$
$$= \int \left(\frac{m}{\left(1 - u^{2}/c^{2}\right)^{3/2}} u \right) \cdot \frac{du}{dt} dt = \int \frac{d}{dt} \left(\frac{mc^{2}}{\sqrt{1 - u^{2}/c^{2}}} \right) dt$$
$$= \int \frac{dE}{dt} dt = E_{\text{final}} - E_{\text{initial}} \quad \text{, q.e.d.}$$

Newton's laws in relativity (continued)

□ **Third law**: "for every action and equal and opposite reaction." This clearly doesn't apply in relativity:

- Suppose two extended objects exert forces *F*(*t*) and -*F*(*t*) on each other in some reference frame, so that the third law is satisfied at all times *t*: *F*(*t*) and -*F*(*t*) are simultaneously applied
- A observer in a different reference frame would see those forces applied at different times! Since the objects are not at the same spatial point, events simultaneous in the first frame will not appear so in the second frame, so unless the forces are *constant* they will not appear equal and opposite.

This won't surprise those who remember radiation reaction.

The Minkowski (four-)force

Since F = dp/dt, the values of a force seen from different inertial reference frames are not related simply by a Lorentz transformation, but instead by a transformation similar to velocity addition.

However, the vector

$$\boldsymbol{K} = \frac{d\boldsymbol{p}}{d\tau} = \frac{dt}{d\tau} \frac{d\boldsymbol{p}}{dt} = \frac{1}{\sqrt{1 - u^2/c^2}} \boldsymbol{F}$$

can clearly be part of a four-vector:

$$K^{0} = \frac{dp^{0}}{d\tau} = \frac{d}{d\tau}\frac{E}{c} \implies K^{\mu} = \frac{dp^{\mu}}{d\tau} \quad Minkowski$$
 force

The scalar product of *K* with itself is therefore Lorentzinvariant (Griffiths problem 12.39):

$$\begin{split} K_{\mu}K^{\mu} &= -\left(K^{0}\right)^{2} + K \cdot K = -\left(\frac{d}{d\tau}\frac{E}{c}\right)^{2} + \frac{F^{2}}{1 - u^{2}/c^{2}} \quad .\\ \frac{d}{d\tau}\frac{E}{c} &= \frac{1}{c}\frac{dt}{d\tau}\frac{dE}{dt} = \frac{1}{c}\frac{1}{\sqrt{1 - u^{2}/c^{2}}}\frac{d}{dt}\left(\frac{mc^{2}}{\sqrt{1 - u^{2}/c^{2}}}\right)\\ &= \frac{1}{c}\frac{1}{\sqrt{1 - u^{2}/c^{2}}}\frac{mc^{2}}{\left(1 - u^{2}/c^{2}\right)^{3/2}}\frac{u}{c^{2}}\frac{du}{dt} = \frac{m}{c}\frac{1}{\left(1 - u^{2}/c^{2}\right)}u \cdot a \end{split}$$

Compare this to

$$\mathbf{u} \cdot \mathbf{F} = \mathbf{u} \cdot \frac{d\mathbf{p}}{dt} = \mathbf{u} \cdot \frac{d}{dt} \left(\frac{m\mathbf{u}}{\sqrt{1 - u^2/c^2}} \right)$$

We worked this out in the middle of an integral a few pages ago:

$$u \cdot F = \frac{m}{\left(1 - u^2/c^2\right)^{3/2}} u \cdot \frac{du}{dt} = \frac{m(u \cdot a)}{\left(1 - u^2/c^2\right)^{3/2}} = c\sqrt{1 - u^2/c^2} K^0$$

$$uF \cos \theta = ;$$

$$K^0 = \frac{uF \cos \theta}{c\sqrt{1 - u^2/c^2}} .$$

Thus

$$\begin{split} K_{\mu}K^{\mu} &= -\left(\frac{uF\cos\theta}{c\sqrt{1-u^2/c^2}}\right)^2 + \frac{F^2}{1-u^2/c^2} \\ &= \frac{1\!-\!\left(u^2/c^2\right)\!\cos^2\theta}{1\!-\!u^2/c^2} F^2 \quad . \end{split}$$

Utility: if *F* is measured at rest, an observer in a moving frame will measure

$$\frac{1 - \left(u^2/c^2\right)\cos^2\overline{\theta}}{1 - u^2/c^2}\overline{F}^2 = F^2$$

Can we use the Minkowski force to cast Newton's second law? Yes, as it turns out. (Griffiths problem 12.38)

□ First define the four-acceleration in terms of the four-velocity:

$$\alpha^{\mu} = d\eta^{\mu} / d\tau = d^2 x^{\mu} / d\tau^2$$

□ In these terms, the second law is

$$K^{\mu} = \frac{dp^{\mu}}{d\tau} = m\frac{d\eta^{\mu}}{d\tau} = m\alpha^{\mu}$$

This bears an odd relationship to the four-velocity itself, as we can see from the various Lorentz invariants we can construct:

The inner product of the four-velocity with itself turns out to be constant:

$$\eta_{\mu}\eta^{\mu} = -\left(\eta^{0}\right)^{2} + \eta \cdot \eta = -\frac{c^{2}}{1 - u^{2}/c^{2}} + \frac{u^{2}}{1 - u^{2}/c^{2}} = -c^{2}$$

And this turns out to mean that the four-velocity and four-acceleration are "orthogonal:"

$$\frac{d}{d\tau} \left(\eta_{\mu} \eta^{\mu} \right) = \alpha_{\mu} \eta^{\mu} + \eta_{\mu} \alpha^{\mu} = 2 \eta_{\mu} \alpha^{\mu} = \frac{d}{d\tau} \left(-c^2 \right) = 0 \quad ;$$

$$\eta_{\mu} \alpha^{\mu} = 0 \quad .$$

 \Box Similarly, $K^{\mu}\eta_{\mu} = 0$.

Alas, real forces are not like the Minkowski force; we still need to derive their transformations. To wit:

□ For finite intervals of momentum and time, seen from two inertial reference frames in relative motion along the x axis,

$$\Delta \overline{p}_{x} = \gamma \left(\Delta p_{x} - \beta \frac{E}{c} \right)$$
$$\Delta \overline{p}_{y} = \Delta p_{y}$$
$$\Delta \overline{p}_{z} = \Delta p_{z}$$
$$\Delta \overline{t} = \gamma \left(\Delta t - \beta \frac{\Delta x}{c} \right)$$

□ Thus we can, in the limit, get a component of the force:

$$\overline{F}_{x} = \lim_{\Delta p, \Delta t \to 0} \frac{\Delta \overline{p}_{x}}{\Delta \overline{t}} = \lim_{\Delta p, \Delta t \to 0} \frac{\gamma \left(\Delta p_{x} - \beta \frac{E}{c}\right)}{\gamma \left(\Delta t - \beta \frac{\Delta x}{c}\right)}$$

Note that

$$\Delta x = \frac{1}{2}a\Delta t^{2} = \frac{1}{2m}\frac{dp_{x}}{dt}\Delta t^{2} \quad \text{and} \quad \Delta E = \frac{1}{2m}\left(\frac{dp_{x}}{dt}\Delta t\right)^{2}$$

are both second order in Δt , so in the limit the second term in both numerator and denominator are small compared to the first.

 $\overline{F}_{x} \cong \lim_{\Delta p, \Delta t \to 0} \frac{\gamma \Delta p_{x}}{\gamma \Delta t} = \frac{dp_{x}}{dt} = F_{x} \quad .$

□ By the same token

□ Thus,

$$\overline{F}_{y} = \lim_{\Delta p, \Delta t \to 0} \frac{\Delta \overline{p}_{y}}{\Delta \overline{t}} = \lim_{\Delta p, \Delta t \to 0} \frac{\Delta p_{y}}{\gamma \left(\Delta t - \beta \frac{\Delta x}{c} \right)}$$
$$= \frac{1}{\gamma} \frac{dp_{y}}{dt} = \frac{F_{y}}{\gamma} ,$$
$$\overline{F}_{z} = \frac{F_{z}}{\gamma} \quad \text{similarly.}$$

□ More compactly,

$$\overline{F}_{\parallel} = F_{\parallel}$$
 , $\overline{F}_{\perp} = \frac{1}{\gamma}F_{\perp}$,

where \parallel and \perp means parallel and perpendicular to the direction of relative motion between the two different inertial frames.