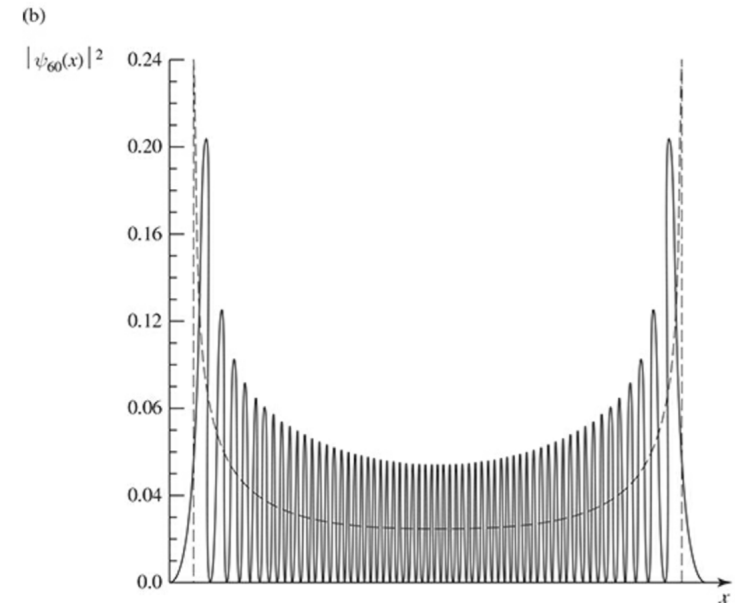
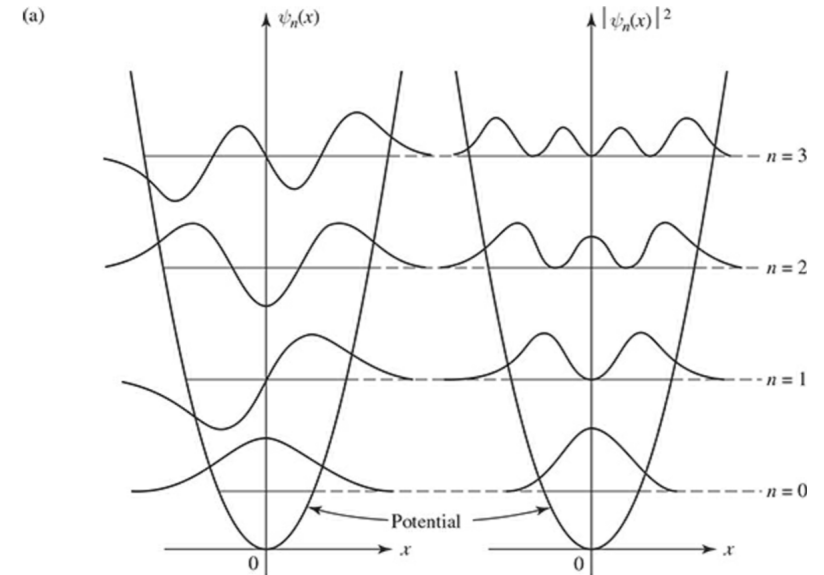


Today in Physics 237: normalization, non-normalizability

- Normalizing the wavefunctions for the quantum simple harmonic oscillator
- Outline of the analytical solution to the Schrödinger equation for the quantum simple harmonic oscillator
- The free particle: travelling-wave Schrödinger-equation solutions



Finishing touches to 1-D QSHO, operator-algebra style:

- Last time we found the ground state of the quantum simple harmonic oscillator, and the ladder operators which can generate the spectrum of this system's states:

$$\psi_0(x) = A_0 e^{-m\omega x^2/\hbar} \quad , \quad \hat{a}_{\pm} = \frac{1}{\sqrt{2m\hbar\omega}} (\mp i\hat{p} + m\omega x) \quad .$$

- Implicit in this is the spectrum of states: $\psi_\nu(x) = A_\nu (\hat{a}_+)^{\nu} e^{-m\omega x^2/\hbar}$ and $E_\nu = (\nu + 1/2)\hbar\omega$; implicit also is the other part of the separation solution, $\varphi(t) = e^{-iE_\nu t/\hbar}$.

- It remains to find the normalizing factor A_ν . This goes quickest with a lemma about the ladder operators:

For any square-integrable functions $f(x)$ and $g(x)$, $\int_{-\infty}^{\infty} f^*(\hat{a}_\pm g) dx = \int_{-\infty}^{\infty} (\hat{a}_\mp f)^* g dx$. Proof:

$$\int_{-\infty}^{\infty} f^*(\hat{a}_\pm g) dx = \frac{1}{\sqrt{2m\hbar\omega}} \int_{-\infty}^{\infty} f^* \left(\mp \hbar \frac{d}{dx} + m\omega x \right) g dx = \frac{1}{\sqrt{2m\hbar\omega}} \left[\int_{-\infty}^{\infty} f^* \left(\mp \hbar \frac{d}{dx} \right) g dx + m\omega \int_{-\infty}^{\infty} x f^* g dx \right]$$

Finishing touches (continued)

- With $u = f^*$, $du = \frac{df^*}{dx} dx$, $dv = \frac{dg}{dx} dx$, $v = g$, do the first of these integrals by parts, and note that $\mp i\hbar = (\pm i\hbar)^*$:

$$\mp \hbar \int_{-\infty}^{\infty} f^* \frac{dg}{dx} g dx = uv \Big|_{-\infty}^{\infty} - \int_{-\infty}^{\infty} v dx = \cancel{\mp \hbar f^* g} \Big|_{-\infty}^{\infty} - (\mp \hbar) \int_{-\infty}^{\infty} \frac{df^*}{dx} g dx = - \int_{-\infty}^{\infty} \left(\mp \hbar \frac{df}{dx} \right)^* g dx \quad .$$

- Put this result into the last expression on the previous page:

$$\begin{aligned} \int_{-\infty}^{\infty} f^* (\hat{a}_{\pm} g) dx &= \frac{1}{\sqrt{2m\hbar\omega}} \left[- \int_{-\infty}^{\infty} \left(\mp \hbar \frac{df}{dx} \right)^* g dx + m\omega \int_{-\infty}^{\infty} x f^* g dx \right] = \frac{1}{\sqrt{2m\hbar\omega}} \left[- \int_{-\infty}^{\infty} (\mp i\hat{p}f)^* g dx + m\omega \int_{-\infty}^{\infty} x f^* g dx \right] \\ &= \frac{1}{\sqrt{2m\hbar\omega}} \left[\int_{-\infty}^{\infty} \left(+[\pm i\hat{p}f]^* g + [m\omega x f^*] g \right) dx \right] = \int_{-\infty}^{\infty} (\hat{a}_{\mp} f)^* g dx \quad \boxed{\text{q.e.d.}} \end{aligned}$$

Finishing touches (continued)

- It is also handy to note that, since $\hat{H}\psi_\nu = \hbar\omega\left(\hat{a}_\pm a_\mp \pm \frac{1}{2}\right)\psi_\nu = \hbar\omega\left(\nu + \frac{1}{2}\right)\psi_\nu$, $\hat{a}_+\hat{a}_-\psi_\nu = \nu\psi_\nu$ and $\hat{a}_-\hat{a}_+\psi_\nu = (\nu+1)\psi_\nu$.
- Applied to an excited state, $\hat{a}_+\hat{a}_-$ returns the same state, times the number of $\hbar\omega$ s in its energy above the ground state.
- Onward to normalize $\hat{a}_\pm\psi_\nu$. Suppose that $\hat{a}_+\psi_\nu = c_\nu\psi_{\nu+1}$ and $\hat{a}_-\psi_\nu = d_\nu\psi_{\nu-1}$, where c_ν and d_ν are constants. Now,

$$\int_{-\infty}^{\infty} (\hat{a}_\pm\psi_\nu)^* (\hat{a}_\pm\psi_\nu) dx = \int_{-\infty}^{\infty} (\hat{a}_\mp\hat{a}_\pm\psi_\nu)^* \psi_\nu dx$$

according to our new lemma on pages 2-3

$$\left\{ \begin{array}{l} |c_\nu|^2 \\ |d_\nu|^2 \end{array} \right\} \int_{-\infty}^{\infty} \psi_{\nu\pm 1}^* \psi_{\nu\pm 1} dx = \left\{ \begin{array}{l} \nu+1 \\ \nu \end{array} \right\} \int_{-\infty}^{\infty} \psi_\nu^* \psi_\nu dx$$

by definition (LHS), and by the Handy expressions above (RHS)

$$\Rightarrow |c_\nu|^2 = \nu+1 \quad , \quad |d_\nu|^2 = \nu$$

since we suppose that A_ν , on page 2, is the correct normalization factor for ψ_ν , etc.

Finishing touches (continued)

- Thus $\hat{a}_+ \psi_v = \sqrt{v+1} \psi_{v+1}$ and $\hat{a}_- \psi_v = \sqrt{v} \psi_{v-1}$.
- These expressions give us the ratios of the normalization factors A_v , so we can start at the ground state and work our way up from ψ_0 :

$$\psi_1 = \frac{1}{\sqrt{1}} \hat{a}_+ \psi_0 \quad , \quad \psi_2 = \frac{1}{\sqrt{2}} \hat{a}_+ \psi_1 = \frac{1}{\sqrt{2 \cdot 1}} (\hat{a}_+)^2 \psi_0 \quad , \quad \psi_3 = \frac{1}{\sqrt{3}} \hat{a}_+ \psi_2 = \frac{1}{\sqrt{3 \cdot 2 \cdot 1}} (\hat{a}_+)^3 \psi_0 \quad ,$$

$$\psi_4 = \frac{1}{\sqrt{4}} \hat{a}_+ \psi_3 = \frac{1}{\sqrt{4 \cdot 3 \cdot 2 \cdot 1}} (\hat{a}_+)^4 \psi_0 \quad \dots \quad \psi_v = \frac{1}{\sqrt{v!}} (\hat{a}_+)^v \psi_0 \quad ,$$

where, as we found [last lecture](#), $\psi_0(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} e^{-m\omega x^2/2\hbar}$.

- The states of different v are orthogonal too:

Finishing touches (continued)

$$\int_{-\infty}^{\infty} \psi_u^* (\hat{a}_+ \hat{a}_-) \psi_v dx = v \int_{-\infty}^{\infty} \psi_u^* \psi_v dx$$

by the Handy expressions at the top of page 4

$$\int_{-\infty}^{\infty} (\hat{a}_- \psi_u)^* (\hat{a}_- \psi_v) dx =$$

by the lemma on pages 2-3

$$\int_{-\infty}^{\infty} (\hat{a}_+ \hat{a}_- \psi_u)^* \psi_v dx =$$

by the lemma again

$$u \int_{-\infty}^{\infty} \psi_u^* \psi_v dx =$$

by those Handy expressions again

- If $u \neq v$, this requires $\int_{-\infty}^{\infty} \psi_u^* \psi_v dx = 0$. And since the ψ_v are normalized, $\int_{-\infty}^{\infty} \psi_u^* \psi_v dx = \delta_{uv}$.

- So the ψ_v for the 1-D quantum simple harmonic oscillator are a complete orthonormal set of functions, permitting Fourier series formulation of the time-dependent wavefunctions and Fourier's Trick, just like the infinite square well.

Now for the hard way.

We won't do this in detail though.

- Back to the time-independent Schrödinger equation: $\left(-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + \frac{1}{2} m \omega^2 x^2 \right) \psi = E \psi$.
- Make some abbreviations: $\xi = \sqrt{\frac{m\omega}{\hbar}} x$, $K = \frac{2E}{\hbar\omega} \Rightarrow \left(\frac{d^2}{d\xi^2} - \xi^2 + K \right) \psi = 0$. This looks hard to solve, and is.
- To make a long and complicated story short: this equation has the solution

$$\psi_\nu(\xi) = \left(\frac{m\omega}{\pi\hbar} \right)^{1/4} \frac{1}{\sqrt{2^\nu \nu!}} H_\nu(\xi) e^{-\xi^2/2} , \quad \text{with eigenvalues } K_\nu = 2\nu + 1 , \quad \nu = 1, 2, \dots$$

and where the functions $H_\nu(\xi)$ are the **Hermite polynomials**.

- Note that, as expected from the solution we got via operator algebra, it contains a Gaussian factor, and familiar normalization factors.

The hard way (continued)

- Compare apples to apples:

$$\psi_\nu(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} \frac{1}{\sqrt{\nu!}} \frac{1}{\sqrt{2^\nu}} H_\nu\left(\sqrt{\frac{m\omega}{\hbar}}x\right) e^{-m\omega x^2/2\hbar}, \quad E_\nu = \hbar\omega\left(\nu + \frac{1}{2}\right), \quad \nu = 0, 1, 2, \dots$$

$$\psi_\nu(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} \frac{1}{\sqrt{\nu!}} (\hat{a}_+)^{\nu} e^{-m\omega x^2/2\hbar}, \quad E_\nu = \left(\nu + \frac{1}{2}\right)\hbar\omega, \quad \nu = 0, 1, 2, \dots$$

- The extra factor of $1/\sqrt{2^\nu}$ is a convention in the definition of the Hermite polynomials.
- So if you need a large- ν wavefunction and don't have time to calculate $(\hat{a}_+)^{\nu} e^{-m\omega x^2/2\hbar}$, you can look up the Hermite polynomials in tables instead.

The hard way (continued)

- For those experienced in MATH 281 and/or PHYS 217: the Hermite polynomials have some of the properties familiar from **Legendre polynomials**, including
 - they're just polynomials: easy to integrate.
 - they're orthonormal and complete: easy to integrate products of them, and represent functions in Fourier-like series with coefficients solvable *via* Fourier's Trick.
 - they can be computed *via* a **Rodrigues formula**, $H_\nu(\xi) = (-1)^\nu e^{\xi^2} \left(\frac{d}{d\xi} \right)^\nu e^{-\xi^2}$, which looks a lot like the product-of-creation-operators version (see G&S Problem 2.16a).
 - they have a generating function, $e^{-z^2+2z\xi}$, which preserves the form of the ground state but yields to a simple power-series expression (see G&S problem 2.16d).
 - The generating function of Legendre polynomials is $1/\nu = 1/|\mathbf{r} - \mathbf{r}'|$, for which the power-series representation leads to the definition and use of multipole moments in electrostatics and magnetostatics.
 - ... and is why we will see them when we deal with the hydrogen atom.

The hard way (continued)

- The first six Hermite polynomials:

$$H_0 = 1,$$

$$H_1 = 2\xi,$$

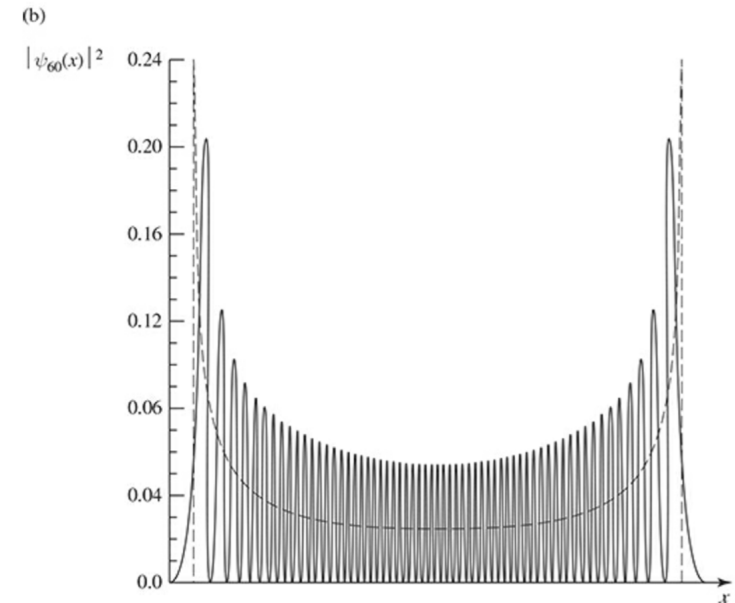
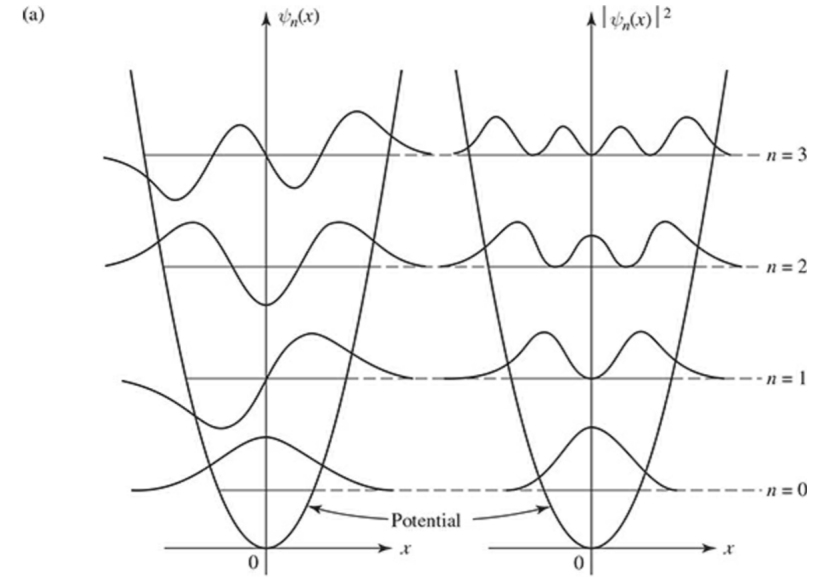
$$H_2 = 4\xi^2 - 2,$$

$$H_3 = 8\xi^3 - 12\xi,$$

$$H_4 = 16\xi^4 - 48\xi^2 + 12,$$

$$H_5 = 32\xi^5 - 160\xi^3 + 120\xi.$$

Note that they alternate between even and odd with respect to their argument, as Legendre polynomials also do.



The free quantum

- The simplest looking case, but a subtle one.
- The time-independent Schrödinger equation for the free quantum is merely $-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} \psi = E\psi = -k^2\psi$,
same as inside the infinite square well, but this time with no boundary conditions.
- In the infinite square well, the boundary conditions “chose” standing waves, for which sine and cosine particular solutions are most convenient.
- Perforce, the free quantum travels, and we will find the Euler-formula equivalent, complex exponential solutions handier:

$$\psi(x) = Ae^{ikx} + Be^{-ikx}$$

$$\Psi(x,t) = Ae^{ikx - iEt/\hbar} + Be^{-ikx - iEt/\hbar}$$

$$= Ae^{ik(x - \hbar kt/2m)} + Be^{-ik(x + \hbar kt/2m)} .$$

The free quantum (continued)

- The two terms represent travelling waves, in which a peak or a wavepacket travels to $+x$ (first term) or $-x$ (second).
- We can bundle the two terms together by letting k be a signed variable:

$$k = \pm\sqrt{2mE}/\hbar: \quad + \text{ travels to } +x, \quad - \text{ to } -x;$$

$$\Psi_k(x,t) = A_k e^{i(kx - \hbar k^2 t/2m)} .$$

- Features detectable at a glance:
 - Wavelength, momentum, and speed: $\lambda = 2\pi/k$, $p = \hbar k$, $v = \hbar k/2m$.
 - Quantum speed not the same as classical: $v = \hbar k/2m = \sqrt{E/2m}$, not $E = mv^2/2 \Rightarrow v_{\text{classical}} = \sqrt{2E/m} = 2v_{\text{quantum}}$.
 - **Wavefunction not normalizable:** $\int_{-\infty}^{\infty} |\Psi_k(x,t)|^2 dx = |A|^2 \int_{-\infty}^{\infty} dx \rightarrow \pm\infty$.