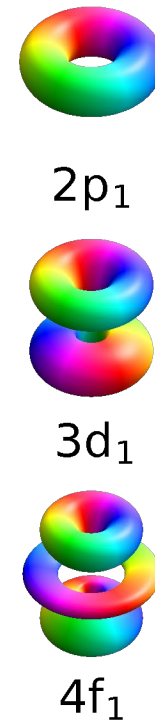
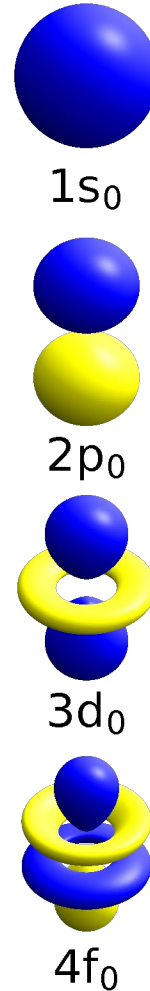
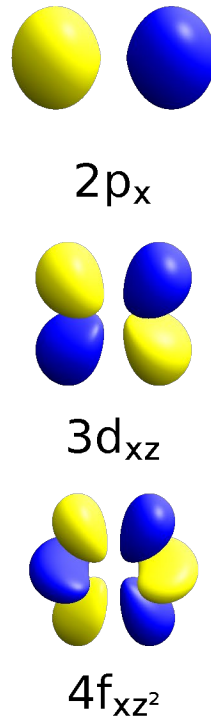
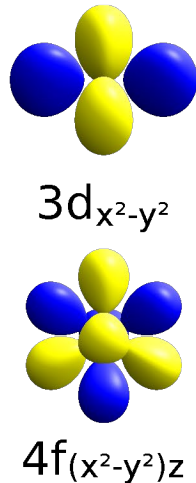
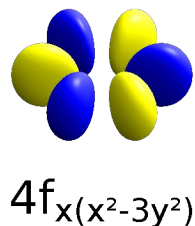


Today in Physics 237: the hydrogen atom

- The radial equation, the solution to which is long and tedious
- The Bohr series
- The complete H wavefunction

As chemists like to picture H-atom orbitals



As physicists like to picture H-atom orbitals

Solution of the radial part of the time-independent Schrödinger equation

- Protons and neutrons are about 1836 times more massive than electrons. To excellent approximation, the H atom's center of mass is the proton's location; r is distance from there.

- The potential energy of the electron-proton system is, of course,

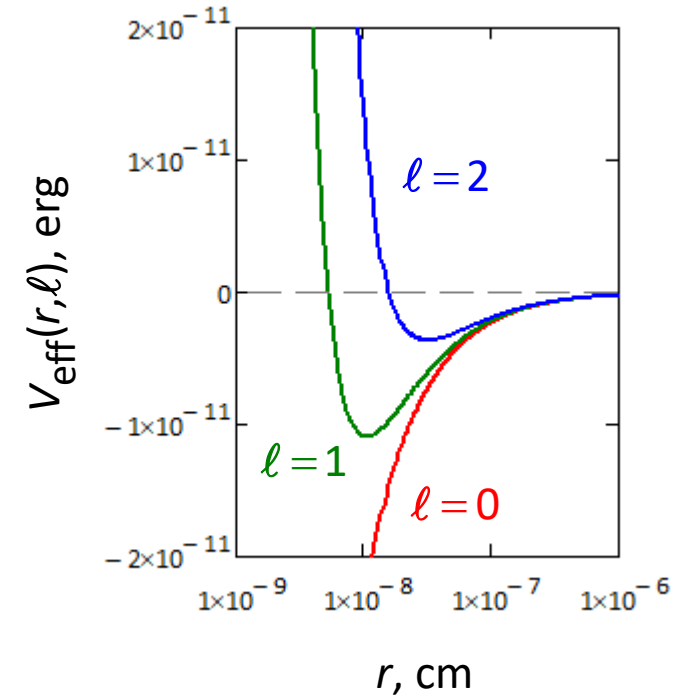
$$V(r) = -\frac{e^2}{r} \quad \text{(in cgs units; replace } e^2 \text{ with } e^2/4\pi\epsilon_0 \text{ for SI)} .$$

- Including the centrifugal term ([Lecture 14](#), page 13), the effective potential energy is

$$V_{\text{eff}}(r, \ell) = -\frac{e^2}{r} + \frac{\hbar^2}{2m_e} \frac{\ell(\ell+1)}{r^2}$$

- Using $u(r) = rR(r)$ as in [Lecture 14](#), the radial equation becomes

$$-\frac{\hbar^2}{2m_e} \frac{d^2 u}{dr^2} + V_{\text{eff}}(r, \ell) u = E u \quad . \quad (V_{\text{eff}})_{\text{min}} \leq E \leq 0 \text{ for bound states}$$



The radial part (continued)

- Now one usually makes some substitutions to render the equation dimensionless:

$$\kappa = \frac{\sqrt{-2m_e E}}{\hbar} \quad , \quad \rho = \kappa r \quad , \quad \rho_0 = \frac{2m_e e^2}{\hbar^2 \kappa} \quad :$$

so that

$$\frac{d^2 u}{d\rho^2} = \left[1 - \frac{\rho_0}{\rho} + \frac{\ell(\ell+1)}{\rho^2} \right] u \quad .$$

- Not a differential equation for which we know the solution. This time it will help to look at **two** solutions for **asymptotic** forms of the equation, which will give multiplicative factors of the solution. We would do so in hope that the rest of $u(\rho)$ will be solved more easily than *via* the differential equation above.

- If $\rho \rightarrow \infty$, the first term in [...] is much larger than the other two, which gives a familiar differential equation:

0, to keep the wavefunction normalizable as $\rho \rightarrow \infty$

$$\rho \rightarrow \infty: \quad \frac{d^2 u}{d\rho^2} \cong u \quad \Rightarrow \quad u = Ae^{-\rho} + \cancel{Be^{\rho}} = Ae^{-\rho} \quad .$$

The radial part (continued)

- In the other, $\rho \rightarrow 0$, limit, the $\ell(\ell+1)/\rho^2$ dominates:

$$\rho \rightarrow 0: \frac{d^2 u}{d\rho^2} \cong \frac{\ell(\ell+1)}{\rho^2} u .$$

- As those who took PHYS 217 (their [Lecture 10](#), pp. 3-4) are aware, one should try $u = C\rho^\alpha$ for such equations. I will suggest that $u = C\rho^{\alpha+1}$ may save a step:

$$\frac{du}{d\rho} = (\alpha+1)C\rho^\alpha \quad , \quad \frac{d^2 u}{d\rho^2} = \alpha(\alpha+1)C\rho^{\alpha-2} = \ell(\ell+1)C\rho^{\alpha-2} \quad ;$$

$$\alpha^2 + \alpha - \ell(\ell+1) = 0 \quad \Rightarrow \quad \alpha = \frac{1}{2} \left(1 \pm \sqrt{1 + 4\ell(\ell+1)} \right) \quad \Rightarrow$$

$$1 + 4\ell(\ell+1) = 4 \left(\ell^2 + \ell + \frac{1}{4} \right) = 4 \left(\ell + \frac{1}{2} \right)^2 \quad \Rightarrow \quad \alpha = \frac{1}{2} \pm \left(\ell + \frac{1}{2} \right) = \begin{cases} -\ell & \text{or} \\ \ell + 1 \end{cases} .$$

The radial part (continued)

- So in this limit, the solution of the radial equation must converge on

0, to keep the wavefunction normalizable as $\rho \rightarrow 0$

$$\rho \rightarrow 0: u = C\rho^{\ell+1} + \cancel{D\rho^{-\ell}} = C\rho^{\ell+1} .$$

- Now we return to the full radial equation, and this time seek a solution of the form

$$u(\rho) = \rho^{\ell+1} e^{-\rho} v(\rho) ,$$

which has the correct behavior as $\rho \rightarrow 0$ and $\rho \rightarrow \infty$, to seek the new function $v(\rho)$.

- Convert the derivatives of u to those of v , doing the algebra slowly. First derivative:

$$\frac{du}{d\rho} = (\ell+1)\rho^{\ell} e^{-\rho} v - \rho^{\ell+1} e^{-\rho} v + \rho^{\ell+1} e^{-\rho} \frac{dv}{d\rho} = \rho^{\ell} e^{-\rho} \left[(\ell+1-\rho)v + \rho \frac{dv}{d\rho} \right] .$$

The radial part (continued)

- Second derivative:

$$\begin{aligned}\frac{d^2u}{d\rho^2} &= (\ell\rho^{\ell-1}e^{-\rho} - \rho^\ell e^{-\rho}) \left[(\ell+1-\rho)v + \rho \frac{dv}{d\rho} \right] + \rho^\ell e^{-\rho} \left[-v + (\ell+1-\rho) \frac{dv}{d\rho} + \frac{dv}{d\rho} + \rho \frac{d^2v}{d\rho^2} \right] \\ &= \rho^\ell e^{-\rho} \left\{ \left(\frac{\ell}{\rho} - 1 \right) \left[(\ell+1-\rho)v + \rho \frac{dv}{d\rho} \right] - v + (\ell+2-\rho) \frac{dv}{d\rho} + \rho \frac{d^2v}{d\rho^2} \right\} \\ &= \rho^\ell e^{-\rho} \left\{ \left[\frac{\ell(\ell+1)}{\rho} - \ell - (\ell+1) + \rho - 1 \right] v + 2(\ell+1-\rho) \frac{dv}{d\rho} + \rho \frac{d^2v}{d\rho^2} \right\} \\ &= \rho^\ell e^{-\rho} \left\{ \left[\frac{\ell(\ell+1)}{\rho} - 2(\ell+1) + \rho \right] v + 2(\ell+1-\rho) \frac{dv}{d\rho} + \rho \frac{d^2v}{d\rho^2} \right\} .\end{aligned}$$

The radial part (continued)

- Combine in the radial equation to produce a differential equation for $v(\rho)$:

$$\frac{d^2 u}{d\rho^2} = \left[1 - \frac{\rho_0}{\rho} + \frac{\ell(\ell+1)}{\rho^2} \right] u \Rightarrow$$

$$\rho^\ell e^{-\rho} \left\{ \left[\frac{\ell(\ell+1)}{\rho} - 2(\ell+1) + \rho \right] v + 2(\ell+1-\rho) \frac{dv}{d\rho} + \rho \frac{d^2 v}{d\rho^2} \right\} = \rho^{\ell+1} e^{-\rho} \left[1 - \frac{\rho_0}{\rho} + \frac{\ell(\ell+1)}{\rho^2} \right] v$$

$$\rho \frac{d^2 v}{d\rho^2} + 2(\ell+1-\rho) \frac{dv}{d\rho} + \left[\frac{\ell(\ell+1)}{\rho} - 2(\ell+1) + \rho \right] v = \rho \left[1 - \frac{\rho_0}{\rho} + \frac{\ell(\ell+1)}{\rho^2} \right] v$$

$$\rho \frac{d^2 v}{d\rho^2} + 2(\ell+1-\rho) \frac{dv}{d\rho} + [\rho_0 - 2(\ell+1)] v = 0 \quad .$$

That helped more than it probably looks; all the terms in $1/\rho$ are gone.

The radial part (continued)

- And since all the terms in $1/\rho$ are gone, it works better to propose a power-series solution, $v(\rho) = \sum_{n=0}^{\infty} c_n \rho^n$.
 - ... at the risk of re-admitting the solution we rejected by setting $D = 0$ on page 3; [see below, page 11](#).
- Again calculate the derivatives:

$$\frac{dv}{d\rho} = \sum_{n=0}^{\infty} n c_n \rho^{n-1} = \sum_{j=-1}^{\infty} (j+1) c_{j+1} \rho^j = \cancel{(-1+1) \frac{c_0}{\rho}} + \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j, \quad \frac{d^2v}{d\rho^2} = \sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^{j-1}.$$

- Into the differential equation they go:

$$\rho \frac{d^2v}{d\rho^2} + 2(\ell+1-\rho) \frac{dv}{d\rho} + [\rho_0 - 2(\ell+1)]v = 0$$

$$\sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^j + 2(\ell+1) \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j - 2 \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^{j+1} + [\rho_0 - 2(\ell+1)] \sum_{j=0}^{\infty} c_j \rho^j = 0$$

The radial part (continued)

- The third term can be re-indexed for ease of comparing coefficients of ρ^j :

$$2 \sum_{j=0}^{\infty} (j+1)c_{j+1}\rho^{j+1} = 2 \sum_{n=1}^{\infty} nc_n\rho^n = 2 \sum_{j=0}^{\infty} jc_j\rho^j \quad . \quad \text{Can start sum at zero instead of 1, since the first term is zero regardless.}$$

- This leaves
$$\sum_{j=0}^{\infty} j(j+1)c_{j+1}\rho^j + 2(\ell+1) \sum_{j=0}^{\infty} (j+1)c_{j+1}\rho^j - 2 \sum_{j=0}^{\infty} jc_j\rho^j + [\rho_0 - 2(\ell+1)] \sum_{j=0}^{\infty} c_j\rho^j = 0 \quad , \quad \text{or}$$

$$\sum_{j=0}^{\infty} \{ [j(j+1) + 2(\ell+1)(j+1)]c_{j+1} + [\rho_0 - 2(\ell+1) - 2j]c_j \} \rho^j = 0 \quad .$$

- For this to be valid at all values of ρ , the coefficient of each ρ^j must be zero:

The radial part (continued)

$$c_{j+1} = \frac{2(\ell+1)+2j-\rho_0}{j(j+1)+2(\ell+1)(j+1)} c_j = \frac{2(\ell+j+1)-\rho_0}{(j+1)(j+2\ell+2)} c_j .$$

- If we know any of the c_j , this formula gives us all the rest.
 - We can always count on knowing one – say, c_0 – from normalization. Thus we **almost** have the solution.
- However, one nuance remains, as suspected on page 8. Consider the coefficients at large j , for which

$$j \gg 1: \quad c_{j+1} \cong \frac{2j}{j(j+1)} c_j = \frac{2}{j+1} c_j .$$

- In this limit, $c_1 \cong \frac{2}{1} c_0$, $c_2 \cong \frac{2}{2} c_1 = \frac{2 \cdot 2}{2 \cdot 1} c_0$, $c_3 \cong \frac{2}{3} c_2 = \frac{2 \cdot 2 \cdot 2}{3 \cdot 2 \cdot 1} c_0$, ... $c_j \cong \frac{2^j}{j!} c_0$.

Ionization and normalization

- If this were an exact result instead of an approximation, then the radial solution would be

$$v(\rho) = \sum_{j=0}^{\infty} c_j \rho^j = c_0 \sum_{j=0}^{\infty} \frac{1}{j!} (2\rho)^j = c_0 e^{2\rho} \Rightarrow u(\rho) = \rho^{\ell+1} e^{-\rho} v(\rho) = c_0 \rho^{\ell+1} e^{+\rho} \xrightarrow{\rho \rightarrow \infty} \infty .$$

That is, it would not be normalizable: with the power-series form for $v(\rho)$, **we have inadvertently re-admitted the $D \neq 0$ solutions we ruled out on page 3.**

- So, out they go again. The easiest way to exclude them is to **terminate the sum**, rather than taking it to infinity. That is, there is some index N such that

$$c_{N-1} \neq 0 \quad , \quad c_N = 0 \quad , \quad c_{N+1} = \frac{2}{(N+1)!} c_N = 0 \quad , \quad \dots \quad , \quad c_{M \geq N} = 0 \quad .$$

- Good news: there is a natural point at which to terminate the sum, as follows.

Ionization and normalization (continued)

- From page 10, for $j = N$, we have $c_N = \frac{2(\ell + N) - \rho_0}{N(N + 2\ell + 1)} c_{N-1} = 0 \Rightarrow \ell + N = \frac{\rho_0}{2}$.

- Define

$$n = N + \ell, \quad n = 1, 2, 3, \dots$$

Principal quantum number

- Then, from page 3,

$$\rho_0 = \frac{2m_e e^2}{\hbar^2 \kappa} = \frac{2m_e e^2}{\hbar \sqrt{-2m_e E}} = 2n \Rightarrow E_n = -\frac{m_e e^4}{2\hbar^2} \frac{1}{n^2} \left[-\frac{m_e e^4}{(4\pi\epsilon_0)^2 2\hbar^2} \frac{1}{n^2} \text{ in SI units} \right] \quad \text{Bohr series}$$

- The Bohr series limit, and in turn the truncation of the sum on page 11, corresponds to the fact that **the H atom can be ionized.**
- Before Bohr, the energy spectrum of H was measured accurately by Johannes Rydberg to be $E_n = -Ry/n^2$, $n = 1, 2, 3, \dots$, where $Ry = 2.18 \times 10^{-11}$ erg = 13.6 eV. Bohr's triumph was in showing – much differently than we just did – that Ry is a product of more fundamental constants ($m_e e^4 / \hbar^2$), rather than being an independent physical constant.

The wavefunction of hydrogen

- Furthermore, $\kappa = \frac{m_e e^2}{\hbar^2} \frac{1}{n} \equiv \frac{1}{na}$, $a = \frac{\hbar^2}{m_e e^2} \cdot \left[\frac{4\pi\epsilon_0 \hbar^2}{m_e e^2} \text{ in SI units} \right]$ **Bohr radius**

- Finally, we can write out the wavefunction, by undoing all our various substitutions and abbreviations:

$$\psi_{n\ell m}(r, \vartheta, \varphi) = R_{n\ell}(r) Y_{\ell}^m(\vartheta, \varphi) \quad ,$$

$$R_{n\ell}(r) = \frac{u(r)}{r} = \frac{1}{r} \rho^{\ell+1} e^{-\rho} v(\rho) = \frac{1}{na} \left(\frac{r}{na} \right)^{\ell} e^{-r/na} v_{n\ell} \left(\frac{r}{na} \right) \quad ,$$

where v is a polynomial in ρ given by the terminated sum on page 11:

$$v_{n\ell}(\rho) = \sum_{j=0}^{n-\ell-1} c_j \rho^j \quad , \quad \text{where} \quad c_{j+1} = \frac{2(\ell+j+1) - \rho_0}{(j+1)(j+2\ell+2)} c_j \quad .$$

Normalization of the ground-state H wavefunction

- The lowest energy (ground) state of H has $n=1, \ell=m=0$. From the past few pages above, and from the discussion of spherical harmonics in [Lecture 14](#),

$$\left. \begin{aligned}
 P_0(\cos\vartheta) &= \frac{1}{2^0 0!} \left(\frac{d}{d(\cos\vartheta)} \right)^0 (\cos^2\vartheta - 1)^0 = 1 \quad ; \\
 P_0^0(\cos\vartheta) &= (-1)^0 (1-x^2)^0 \left(\frac{d}{dx} \right)^0 P_0(x) = 1 \quad ; \\
 Y_0^0(\vartheta, \varphi) &= \sqrt{\frac{(0+1)(0)!}{4\pi(0)!}} e^{i0} P_0^0(\cos\vartheta) = \sqrt{\frac{1}{4\pi}} \quad ; \\
 R_{10}(r) &= \frac{1}{a} \left(\frac{r}{a} \right)^0 e^{-r/a} v_{10} \left(\frac{r}{a} \right) = \frac{c_0}{a} e^{-r/a} \quad .
 \end{aligned} \right\} \left\{ \begin{aligned}
 1 &= \int \psi_{100}^* \psi_{100} d\tau = \frac{1}{4\pi} \frac{|c_0|^2}{a^2} \int_0^\pi \int_0^{2\pi} \sin\vartheta d\vartheta d\varphi \int_0^\infty e^{-2r/a} r^2 dr \\
 &= \frac{1}{4\pi} \frac{|c_0|^2}{a^2} (4\pi) \left(\frac{a}{2} \right)^3 \int_0^\infty e^{-u} u^2 du = \frac{|c_0|^2}{a^2} \left(\frac{a}{2} \right)^3 (2!) = \frac{|c_0|^2 a}{4} \\
 \boxed{c_0 = \frac{2}{\sqrt{a}}} &\Rightarrow \boxed{\psi_{100}(r, \vartheta, \varphi) = \frac{1}{\sqrt{\pi a^3}} e^{-r/a}} \quad .
 \end{aligned} \right.$$

Associated Laguerre functions

- All the other coefficients in $v_{n\ell}(\rho) = \sum_{j=0}^{n-\ell-1} c_j \rho^j$ come from this value of c_0 , and $c_{j+1} = \frac{2(\ell+j+1) - \rho_0}{(j+1)(j+2\ell+2)} c_j$.
- An applied mathematician would recognize the expression $v_{n\ell}(\rho)$ and c_{j+1}/c_j . They would write it as

$$v_{n\ell}(\rho) = L_{n-\ell-1}^{2\ell+1}(2\rho) \quad , \quad \text{where} \quad L_q^p(x) = (-1)^p \left(\frac{d}{dx} \right)^p L_{p+q}(x)$$

is called the **associated Laguerre function**, and where

$$L_q(x) = \frac{e^x}{q!} \left(\frac{d}{dx} \right)^q (x^q e^{-x})$$

is the (q th) **Laguerre polynomial**. A few of each are listed in the next pages, which we show to prevent fear that they are dangerous beasts: they are, after all, polynomials. It doesn't – or shouldn't – scare you to perform any mathematical operation on a polynomial.

Laguerre polynomials

$$q \quad L_q(x)$$

$$0 \quad 1$$

$$1 \quad -x + 1$$

$$2 \quad \frac{1}{2}(x^2 - 4x + 2)$$

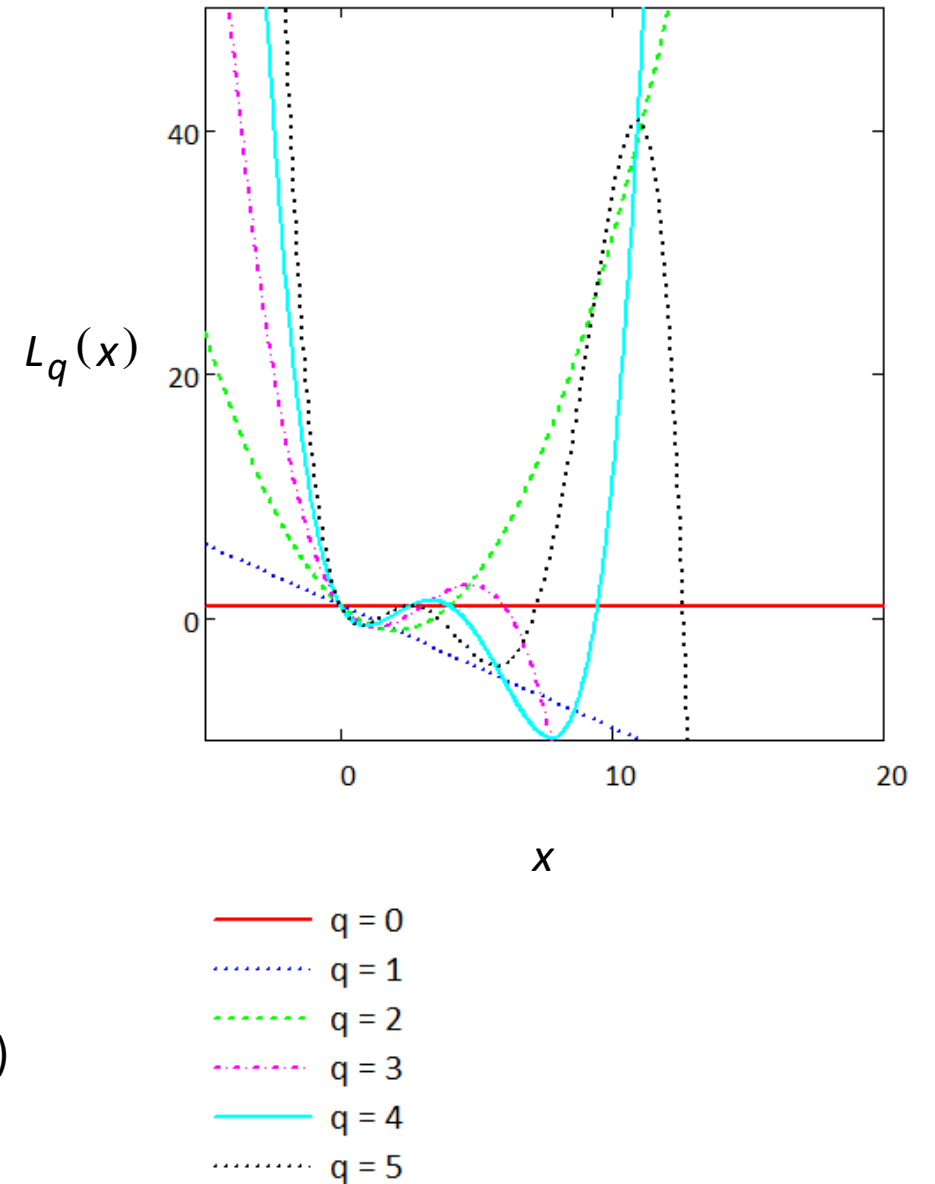
$$3 \quad \frac{1}{6}(-x^3 + 9x^2 - 18x + 6)$$

$$4 \quad \frac{1}{24}(x^4 - 16x^3 + 72x^2 - 96x + 24)$$

$$5 \quad \frac{1}{120}(-x^5 + 25x^4 - 200x^3 + 600x^2 - 600x + 120)$$

$$\vdots$$

$$q \quad \frac{1}{q!}((-x)^q + q^2(-x)^{q-1} + \dots + \left(\frac{q!}{p!(q-p)!}\right)^2 p!(-x)^{q-p} + \dots + q(q!)(-x) + q!)$$



Normalized H wavefunctions

- In all, the H wavefunctions are

$$\psi_{n\ell m}(r, \vartheta, \varphi) = \sqrt{\left(\frac{2}{na}\right)^3 \frac{(n-\ell-1)!}{2n(n+\ell)!}} e^{-r/na} \left(\frac{2r}{na}\right)^\ell L_{n-\ell-1}^{2\ell+1}\left(\frac{2r}{na}\right) Y_\ell^m(\vartheta, \varphi) .$$

- And, as we are getting used to, they are orthonormal:

$$\int_0^{2\pi} \int_0^\pi \int_0^\infty \psi_{n\ell m}^* \psi_{n'\ell'm'} r^2 dr \sin\vartheta d\vartheta d\varphi = \delta_{nn'} \delta_{\ell\ell'} \delta_{mm'} .$$

- They are quite a chore to visualize, though. See G&S page 153 for some monochrome attempts, but don't neglect to search the web for color-coded renditions which may make more sense to you. Page 1, for example.
- Next time: example use of the H wavefunctions