

Lecture #7

We learned last time that

(1)

Entropy (thermodynamics) $S = k_B \ln \Omega$; $\Omega = \#$ of states occupied at energy E .
(Microscopic)

should be helpful in H.W. problem #3.

Simple Example #2: Magnetic Systems

Take N spins localized in space so we are only concerned with energy:

$$H = -g \frac{\mu_B}{\hbar} \vec{S} \cdot \vec{B} ; \vec{S} = \text{spin operator}$$

Align $\vec{B} = B_0 \hat{z}$

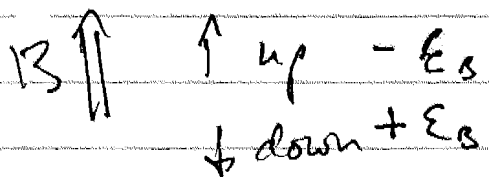
$g = g\text{-factor} (\approx 2 \text{ for free electron})$
 $\mu_B = \frac{e\hbar}{2m} = \text{Bohr Magnetron.}$

$$S_z |\chi_s\rangle = \pm \frac{\hbar}{2} |\chi_s\rangle$$

$$\therefore E_{\pm} = \pm \frac{g_s}{2} \mu_B B_0 \text{ are the energy eigenstates}$$

$$\equiv \pm E_B$$

$N_{\pm} = \#$ of spins that are \uparrow or \downarrow



We now will find Ω by counting the $\#$ of ways to get energy U for N spins

$$U = (N_-) E_B + N_+ (-E_B) = (N_- - N_+) E_B$$

$$N = N_- + N_+$$

$$\underline{\Omega} \quad N_- = \frac{N + U/E_B}{2} , \quad N_+ = \frac{N - U/E_B}{2}$$

$\downarrow \sim -$ first spin has 3 possibilities
 second spin has 2 possibilities
 third " " " 1 "

How many ways can we assign the N spins such that the energy U is constant? (equivalent to fixed N_+, N_-)

(2)

$B \uparrow$
 $N_- = 1$
 $N_+ = 2$

$\uparrow \uparrow \downarrow \quad \Theta = U = -\epsilon_B \Rightarrow \Omega = 3$

$\uparrow \downarrow \uparrow$
 $\downarrow \uparrow \uparrow$
 If we assign the spins without regard to \uparrow or \downarrow , there are $N!$

ways of assigning them. However, each \uparrow spin is indistinguishable from the other \uparrow spins, & same for the \downarrow . So we divide by the # of permutations of \uparrow and \downarrow , $N_+!$ and $N_-!$

$\Rightarrow \Omega = \frac{N!}{N_+! N_-!} = \frac{N!}{\left(\frac{N + U/\epsilon_B}{2}\right)! \left(\frac{N - U/\epsilon_B}{2}\right)!}$

To make further progress, we use Stirling's Approximation

$n! \approx \sqrt{2\pi n} \exp[n \ln n - n] \left[1 - \frac{1}{12n} - \frac{17!}{32n^2} \dots \right]$

We did this last time.

Why? Consider $n! = \int_0^\infty e^{-x} x^n dx \equiv \Gamma(n)$

Integrate by parts: $\Gamma(n) = n \int_0^\infty e^{-x} x^{n-1} dx = n \Gamma(n-1)$

Define Gamma function $\Gamma(z+1) = \int_0^\infty e^{-x} x^z dx \quad \forall z \in \mathbb{C}$

$= \int_0^\infty e^{-x + z \ln x} dx$ evaluate by steepest descent.

Applying this, we see

$$\Omega = \frac{N!}{\left(\frac{N - \frac{U}{\epsilon_B}}{2}\right)! \left(\frac{N + \frac{U}{\epsilon_B}}{2}\right)!}$$

$$\Rightarrow \ln \Omega = \ln N! - \ln \left[\frac{N - \frac{U}{\epsilon_B}}{2}\right]! - \ln \left[\frac{N + \frac{U}{\epsilon_B}}{2}\right]!$$

Replace $\ln x! \approx x \ln x - x$

$$\frac{S}{k_B} = \cancel{N \ln N - N} - \left(\frac{N - \frac{U}{\epsilon_B}}{2}\right) \ln \left[\frac{N - \frac{U}{\epsilon_B}}{2}\right] + \frac{N - \frac{U}{\epsilon_B}}{2} - \left(\frac{N + \frac{U}{\epsilon_B}}{2}\right) \ln \left[\frac{N + \frac{U}{\epsilon_B}}{2}\right] + \frac{N + \frac{U}{\epsilon_B}}{2}$$

$$\frac{S}{k_B} = N \ln N - \frac{N}{2} \ln \left[\left(N - \frac{U}{\epsilon_B}\right) \left(N + \frac{U}{\epsilon_B}\right) \right] + N \ln 2 + \frac{U}{2\epsilon_B} \ln \left[\frac{N - \frac{U}{\epsilon_B}}{N + \frac{U}{\epsilon_B}} \right]$$

$$\begin{aligned} \text{2nd term} &= -\frac{N}{2} \ln \left[N^2 \left(1 - \frac{U}{N\epsilon_B}\right) \left(1 + \frac{U}{N\epsilon_B}\right) \right] \\ &= (-N \ln N) - \frac{N}{2} \ln \left[1 - \frac{U^2}{N^2 \epsilon_B^2} \right] \end{aligned}$$

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$$\therefore \frac{S}{k_B} = N \ln 2 - \frac{N}{2} \ln \left[1 - \frac{U^2}{N^2 \epsilon_B^2} \right] + \frac{U}{2 \epsilon_B} \ln \left[\frac{1 - \frac{U}{N \epsilon_B}}{1 + \frac{U}{N \epsilon_B}} \right]$$

$$\frac{S^*}{k_B} = N \left\{ \ln 2 - \frac{1}{2} \left(1 + \frac{U}{N \epsilon_B} \right) \ln \left(1 + \frac{U}{N \epsilon_B} \right) - \frac{1}{2} \left(1 - \frac{U}{N \epsilon_B} \right) \ln \left(1 - \frac{U}{N \epsilon_B} \right) \right\}$$

We now introduce the temperature:

However, we have $S^*(U, B, N)$, NOT $S(U, M, T)$
↑ field (intensive)

$$dU = T dS + \underbrace{\mu_0 B dM}_{dw \text{ (work)}} + \mu dN$$

$$\text{So } dS = \frac{1}{T} dU - \frac{\mu_0 B dM}{T} - \frac{\mu}{T} dN$$

$$S^* = S + \frac{\mu_0 B}{T} M ; dS^* = \frac{1}{T} dU + \mu_0 M d\left(\frac{B}{T}\right) - \frac{\mu}{T} dN$$

$$\therefore \frac{1}{T} = \left. \frac{\partial S^*}{\partial U} \right|_{\frac{B}{T}, N} \quad \text{; We can't easily hold } \frac{B}{T} \text{ fixed, given the form of } S^* \quad (6)$$

Consider instead a new potential, $U^*(S^*, B, N)$

$$U^* = U - \mu_0 M B \Rightarrow dU^* = T dS^* - \mu_0 M dB + \mu dN$$

$$\Rightarrow T = \left. \frac{\partial U^*}{\partial S^*} \right|_{B, \mu}$$

$$\therefore \frac{S^*}{k_B} = N \left[\ln 2 - \frac{1}{2} \left(1 + \frac{U^*}{N \epsilon_B} \right) \ln \left(1 + \frac{U^*}{N \epsilon_B} \right) - \frac{1}{2} \left(1 - \frac{U^*}{N \epsilon_B} \right) \ln \left(1 - \frac{U^*}{N \epsilon_B} \right) \right]$$

~~Now take~~ Now take $\frac{\partial}{\partial S^*} [\dots]_{B, N}$

$$1 = N k_B \left\{ \left(-\frac{1}{2N \epsilon_B} \right) \frac{\partial U^*}{\partial S^*} \ln \left[1 + \frac{U^*}{N \epsilon_B} \right] + \left(\frac{1}{2N \epsilon_B} \right) \frac{\partial U^*}{\partial S^*} \ln \left[1 - \frac{U^*}{N \epsilon_B} \right] \right\}$$

$$- \frac{1}{2} \left(1 + \frac{U^*}{N \epsilon_B} \right) \left(1 + \frac{U^*}{N \epsilon_B} \right)^{-1} \frac{\partial U^*}{\partial S^*} - \frac{1}{2} \left(1 - \frac{U^*}{N \epsilon_B} \right) \left(1 - \frac{U^*}{N \epsilon_B} \right)^{-1} \left(-\frac{\partial U^*}{\partial S^*} \right)$$

$$1 = \frac{N k_B}{2N \epsilon_B} \left\{ \ln \left[\frac{1 - U^*/N \epsilon_B}{1 + U^*/N \epsilon_B} \right] \right\} T$$

$$\exp \left[\frac{2\epsilon_B}{k_B T} \right] = \frac{1 - U^*/N\epsilon_B}{1 + U^*/N\epsilon_B}$$

$$\Rightarrow U^* = -N\epsilon_B \left[\frac{e^{\epsilon_B/k_B T} - e^{-\epsilon_B/k_B T}}{e^{\epsilon_B/k_B T} + e^{-\epsilon_B/k_B T}} \right]$$

$$U^* = -N\epsilon_B \tanh \left(\frac{\epsilon_B}{k_B T} \right)$$

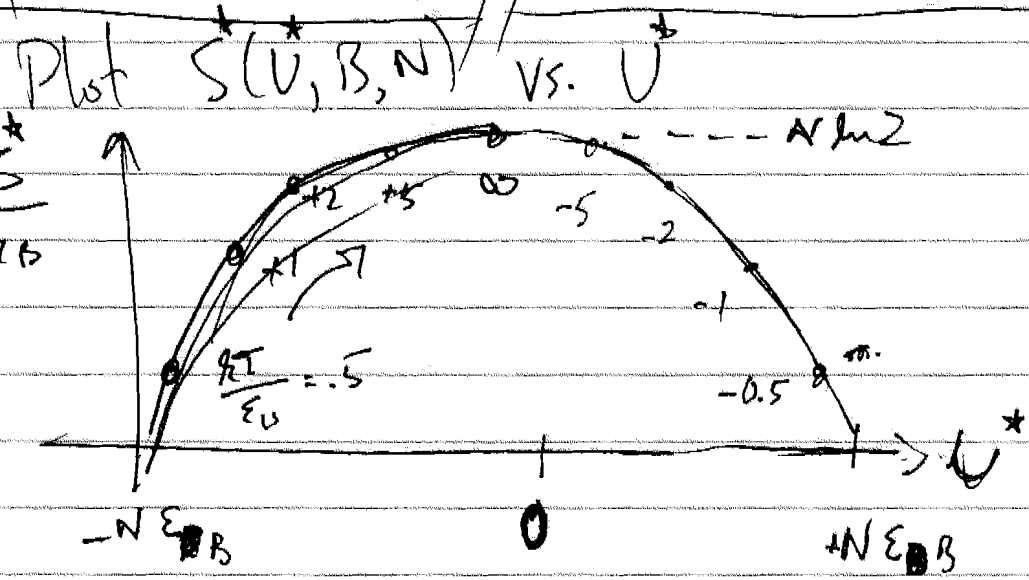
Consider Magnetization M , Each spin has magnetic moment $\mu_B = g_s \mu_B / 2$

(short-cut)
we could also use

$$\therefore M = N_+ \mu_B - \mu_B N_- = \mu_B (N_+ - N_-)$$

$$\mu_B M = \frac{\partial U^*}{\partial B} \Big|_{S, N}$$

$$\Rightarrow M = \frac{-\mu_B N}{\epsilon_B} U^* = N \mu_B \tanh \left(\frac{\epsilon_B}{k_B T} \right)$$



Note 1: for $T \rightarrow 0$ $\tanh(\infty) \rightarrow 1$

$$U^* \rightarrow -N\epsilon_0$$

\therefore there is only 1 microstate (all aligned)

$$S = k_B \ln(1) = 0$$

$$\Rightarrow \text{slope } \frac{\partial S}{\partial U} = \frac{1}{T} \rightarrow \infty \text{ as } T \rightarrow 0$$

Note 2: for $T \rightarrow \infty$, $\tanh(0) = 0$,

$$\text{so } U^* \rightarrow 0$$

here $\frac{1}{2}$ of spins are \uparrow , half are \downarrow .

$S = N \ln 2$ is the maximum entropy.

Recall
 $\Omega = \frac{N!}{N_+! N_-!}$

$$\sum_{N_+} \Omega = \sum_{N_+} \frac{N!}{N_+! (N-N_+)!} = 2^N$$

for $N=3$
 $2 \cdot 2 \cdot 2 = 2^3$

Here slope $T^{-1} = \frac{\partial S}{\partial U}$ is 0.

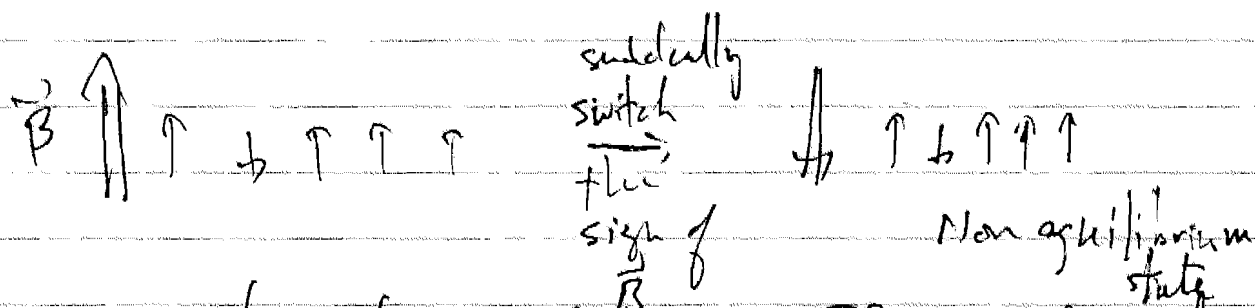
This corresponds to the maximum # of microstates, for the system, $2^N = \Omega_{\text{max}} = \sum_n \Omega_n$

Note 3: for $U > 0$, there can be states of negative temperature, that can be obtained by systems in isolation (NMR pulses to flip spins into anti-parallel state.)

Self-consistent thermodynamics can be developed,

Key For $U > 0$ ($T < 0$) the magnetization is opposite in direction to that of the applied field!

~~Consider~~ Consider the following experiment.
Purcell + Pound (1951)



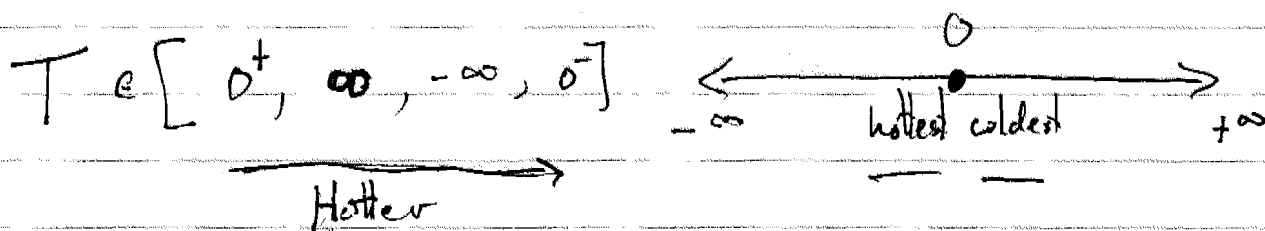
⇒ Negative temperature? If t_1 characterized time of mutual interaction among spins, and t_2 describes time to interact spins with the lattice, then for times $t_1 \leq t \leq t_2$

the ~~sys~~ spins are in equilibrium with each other but not with the lattice.

⇒ Negative temperatures!

for longer times, spin system comes to equilibrium with lattice ⇒ overall positive temp.

Note: $T < 0$ only possible if system has an upper limit to its energy.



Limiting Behavior:

(10)

$$U = -N \epsilon_B \tanh\left(\frac{\epsilon_B}{kT}\right)$$

for $T \rightarrow 0$, ϵ_B/kT is big $\Rightarrow e^{\epsilon_B/kT}$ big
 $e^{-\epsilon_B/kT}$ small

$$U = (-N \epsilon_B) \left[\frac{1 - e^{-2\epsilon_B/kT}}{1 + e^{-2\epsilon_B/kT}} \right] = -N \epsilon_B \left(\frac{1 - \delta}{1 + \delta} \right)$$

$$U \approx -N \epsilon_B [1 - 2\delta + \dots] = -N \epsilon_B \left(1 - 2e^{-\frac{2\epsilon_B}{kT}} + \dots \right)$$

$$\therefore C_B = \left. \frac{\partial U}{\partial T} \right|_{B,N} = 2N \epsilon_B \left(\frac{2\epsilon_B}{kT^2} e^{-\frac{2\epsilon_B}{kT}} + \dots \right)$$

$$C_B = \frac{4N \epsilon_B^2}{kT^2} e^{-\frac{2\epsilon_B}{kT}} \Rightarrow \text{heat capacity decreases exponentially.}$$

for $T \rightarrow \infty$ ϵ_B/kT is small

$$U = -N \epsilon_B \left(\frac{1 - e^{-\delta}}{1 + e^{-\delta}} \right) \approx -N \epsilon_B \left(\frac{1 - (1 - \delta + \dots)}{1 + (1 - \delta + \dots)} \right)$$

$(\delta = 2\epsilon_B/kT)$

$$U = -N \epsilon_B^2 \left(\frac{1}{kT^2} \left(1 + \delta \left(\frac{1}{T^2} \right) \right) \right) \Rightarrow \text{power law decay}$$

$$\therefore C_B = \frac{N \epsilon_B^2}{kT^2} \left(1 + \delta \left(\frac{1}{T^2} \right) \right)$$

