



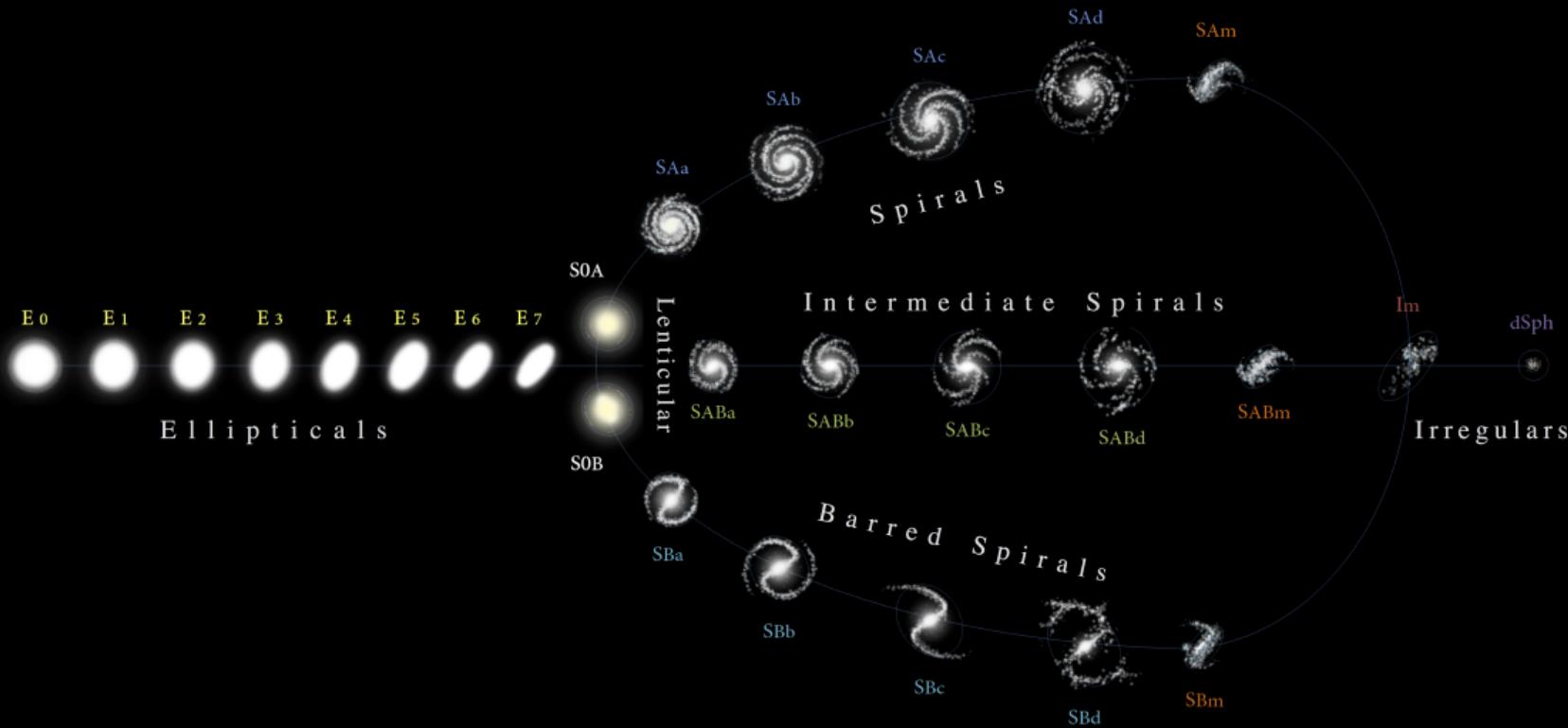
Stars & Galaxies

Galaxy classification
Galaxy population statistics
Clusters & groups of galaxies
Galaxies at high redshifts

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University of Rochester

Galaxy classification: the Hubble sequence



Elliptical galaxies

Messier 87 & NGC 1332



Sérsic profile

$$I(R) = I_0 e^{-\beta_n \left(\frac{R}{R_e} \right)^{1/n}} = I_e e^{-\beta_n \left(\left(\frac{R}{R_e} \right)^{1/n} - 1 \right)}$$

where

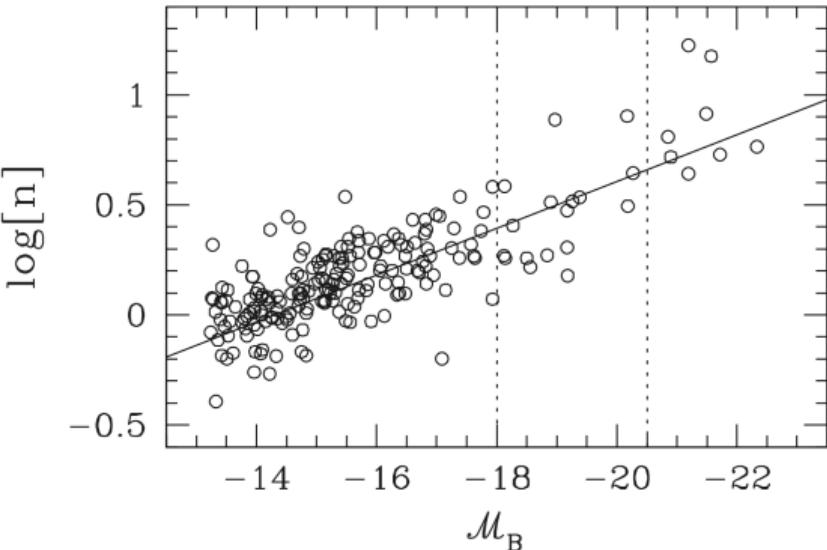
I_0 ≡ central surface brightness

n ≡ Sérsic index

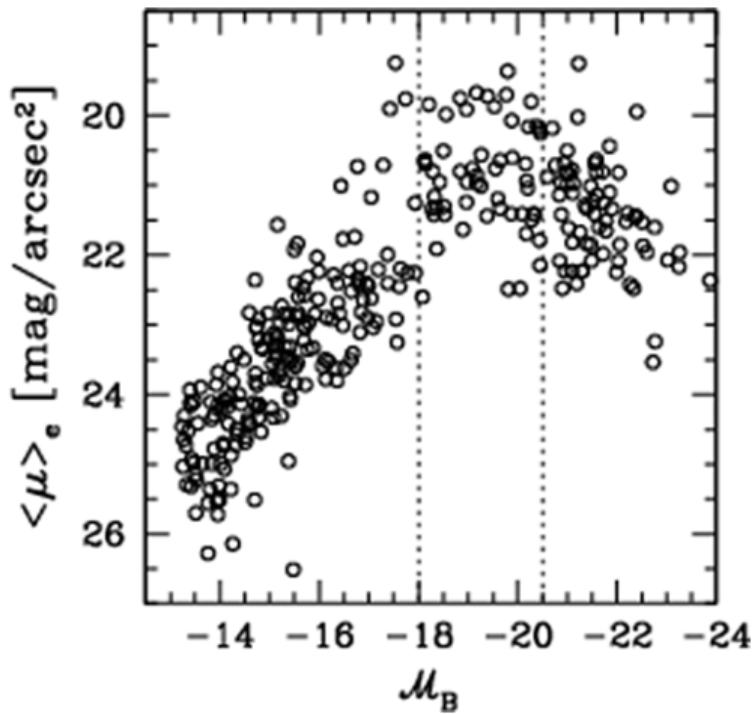
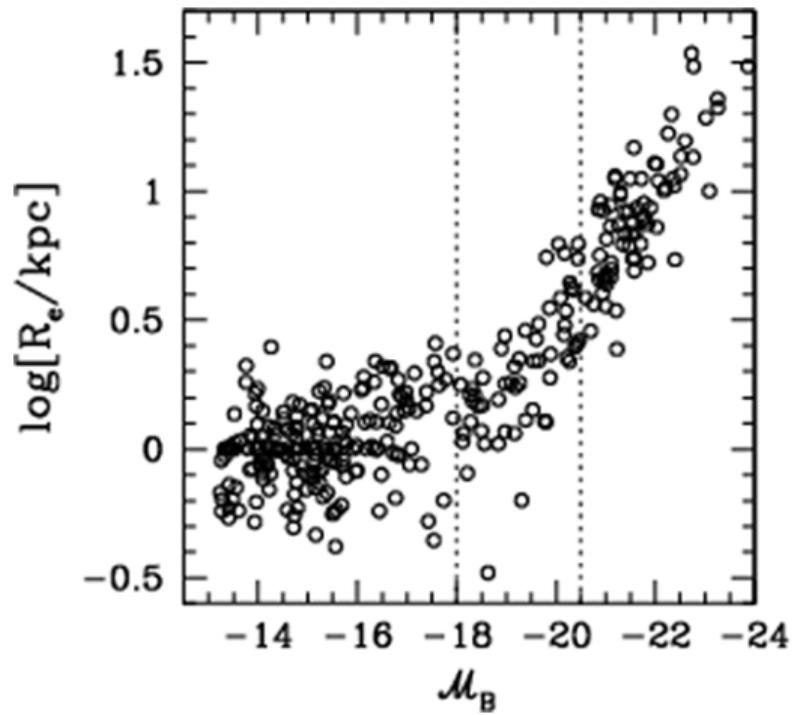
R_e ≡ effective radius enclosing 50%
of the total light

$I_e = I(R_e)$

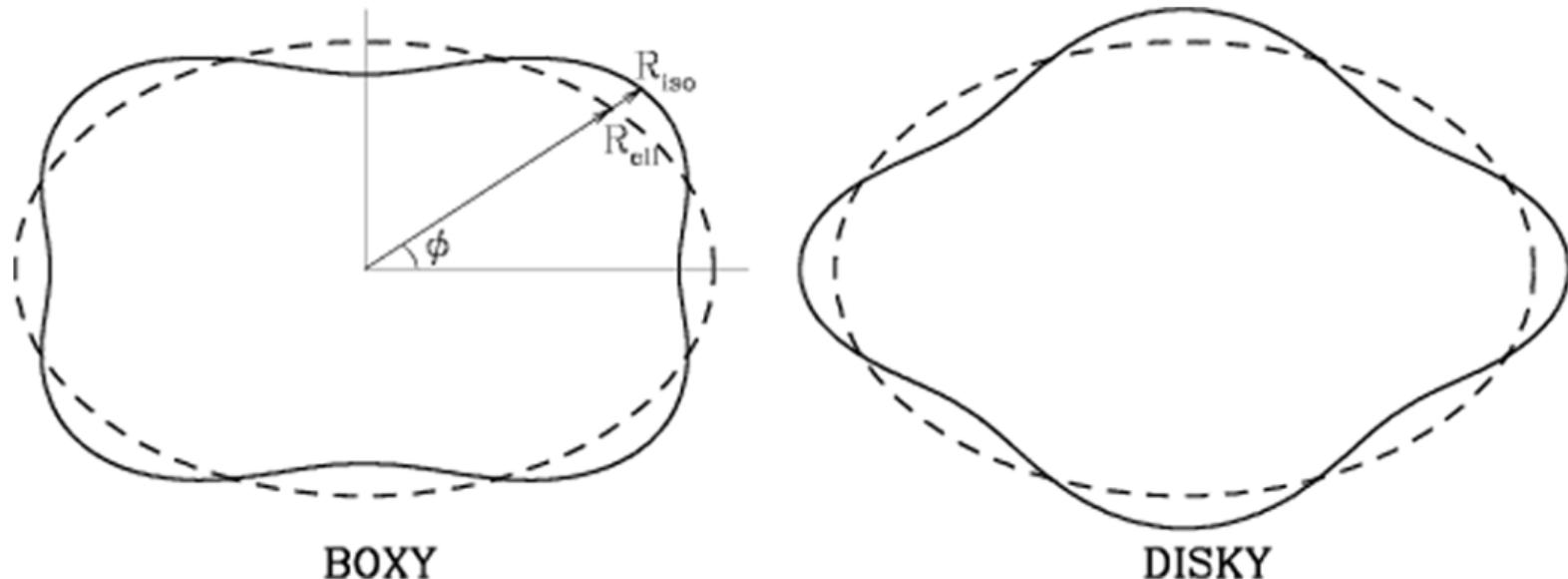
$\beta_n = 2n - 0.324$ for $n \gtrsim 1$



Elliptical galaxy trends



Ellipticity of elliptical galaxies



Kinematics of elliptical galaxies

For isotropic, oblate galaxies flattened by the centrifugal force due to rotation,

$$\frac{v_m}{\bar{\sigma}} \approx \sqrt{\frac{\epsilon}{1 - \epsilon}}$$

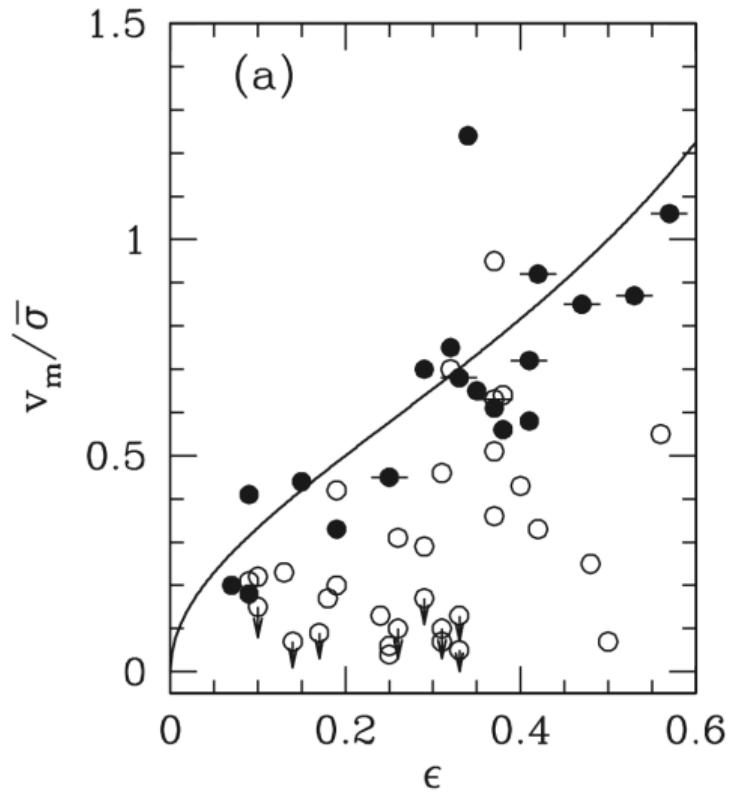
where

v_m \equiv maximum relative velocity along the line of sight

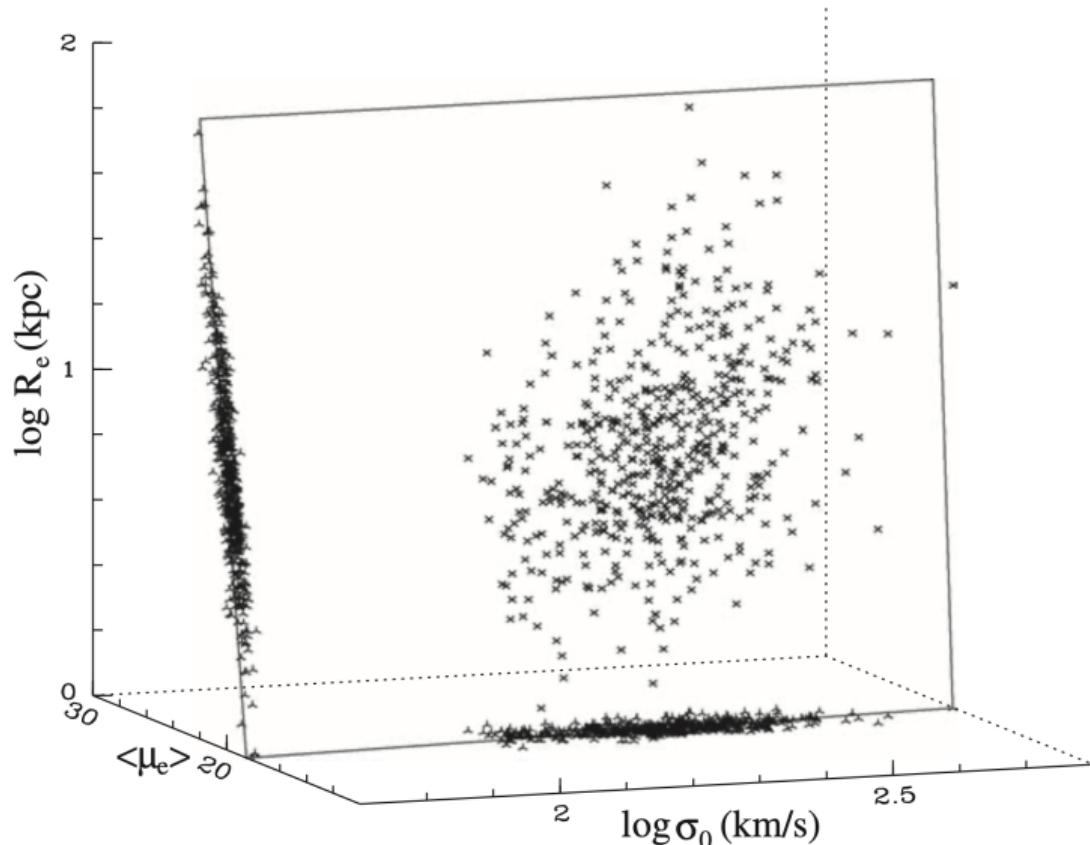
$\bar{\sigma}$ \equiv average line-of-sight velocity dispersion interior to $\sim 0.5R_e$

Open circles: bright galaxies

Filled circles: fainter galaxies



Fundamental plane



Spiral galaxies

Messier 74 & NGC 3898



Barred spiral galaxies

NGC 1398 & NGC 1365



Exponential profile

$$I(R) = I_0 e^{-R/R_d}$$

where

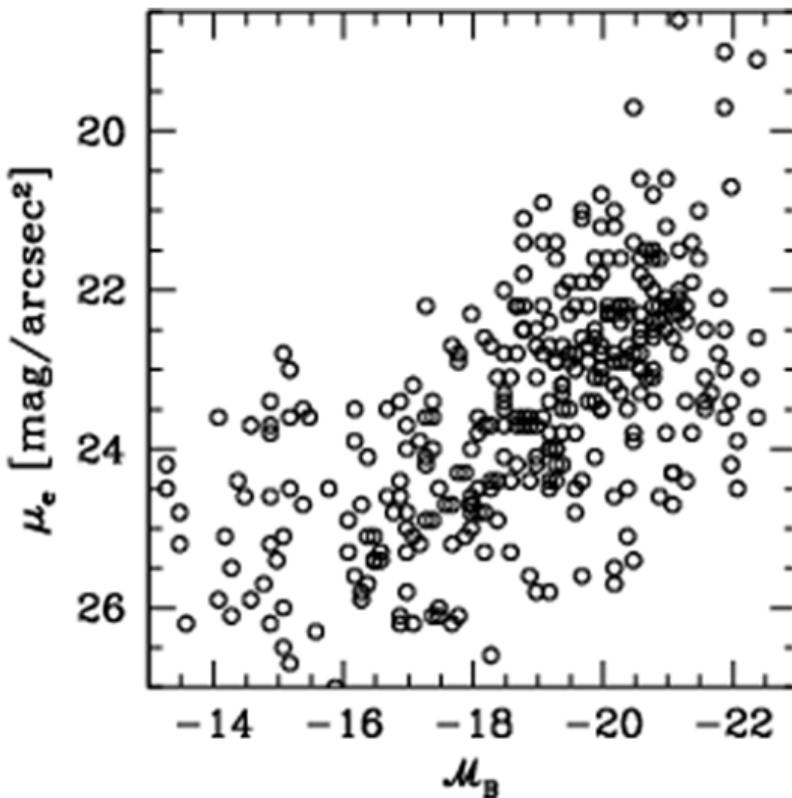
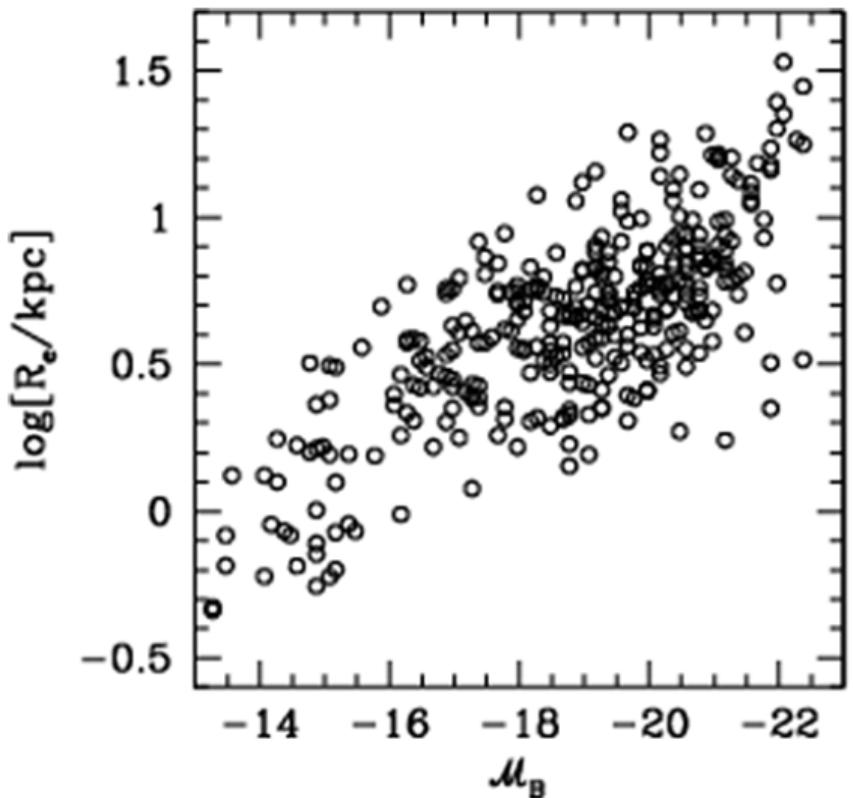
I_0 = $L/2\pi R_d^2$; central luminosity surface density

R ≡ cylindrical radius

R_d ≡ exponential scale length

L ≡ total luminosity

Spiral galaxy trends



Vertical luminosity profile

The surface brightness distribution in the vertical direction is independent of the distance from the center.

$$\nu(R, z) = \nu_0 e^{-R/R_d} f(z) \quad f_n(z) = \operatorname{sech}^{2/n} \left(\frac{n|z|}{2z_d} \right)$$

where

ν ≡ luminosity density in 3D

n controls the shape of the profile near $z = 0$ ($n = 1$ corresponds to a self-gravitating isothermal disk)

z_d ≡ vertical scale height of the disk

Spiral arm patterns

Grand design v. flocculent system

