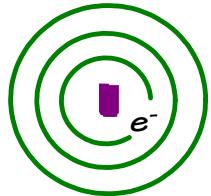


# Physics 123, April 3, 2013

- Exam 2 looms large      Apr. 11, Dewey 1101, 0800
- Up to start of Bohr's model of atom (Apr. 1 lecture)
- With all the crazy optics stuff . . .  
Can use Both Sides of an  $8.5 \times 11$  inch sheet  
for notes + Formulas  
I will provide a formula sheet as well
- More on Material coverage in email soon . . .  
But it will cover from end of Exam 1 material  
to start of Bohr Model

(2)



### Bohr model

quantized circular orbits

$$n=1, 2, 3 \dots$$

$$r_n = \frac{n^2 h^2}{k z e^2 m}$$

with quantized energies

$$E_{\text{TOTAL}} = -\frac{m k^2 Z^2 e^4}{2 n^2 h^2}$$

- $e^-$  in circular orbits
- $e^-$  held in atom via Coulomb attraction
- Single  $e^-$
- $e^-$  exists in discrete stable orbits
- photons emitted as  $e^-$  jumps between orbits
- Photon energy corresponds to difference in energy between the two energy levels

er ... last class I forgot to mention something important

(3)

## Bohr's Model WORKS!!

For single  $e^-$  Atoms Bohr's model reproduces  
Fairly well the experimentally determined  
discrete Atomic Spectra!

Not perfect ... revisions made

e.g., allow nucleus to move + have  $e^-$  orbit center-of-mass  
elliptical orbits

:

- Bohr model gets people to take Rutherford's nuclear atom seriously
- relates quanta of light and quantization of
- Atomic Energy levels

- Generally a decent starting place for intuition

But

- Fails to describe spectra of larger atoms
- Treats  $e^-$  classically (can think of de Broglie Motivating the quantization  
... But this Mixes Wave + Particle for  $e^-$ )
- Fails to describe detailed structure ("fine" structure, Zeeman Splitting  
... etc.)

Successful model of atom must come from a theory  
that treats  $e^-$  as wave from the ground up  $\rightarrow$  quantum Mechanics

## Particles $\hat{R}$ waves leads to a Strange new world

(5)

Recall from waves + Fourier analysis

A Can describe waves in position or spatial frequency  
 $X \quad k = 2\pi/\lambda$

If wave is narrow in  $X$ , require higher frequency components

B Can describe waves in time or temporal frequency  
 $t \quad \omega = 2\pi/T$

If wave is narrow in  $t$ , require higher frequency components

C So, for all waves  $\underbrace{\Delta X \Delta k}_{\text{width in position space}} \sim 1$  and  $\underbrace{\Delta t \Delta \omega}_{\text{width in frequency space}} \sim 1$

(6)

$$\text{De Broglie} \rightarrow p = \frac{h}{\lambda} = h \frac{k}{2\pi} = \hbar k \rightsquigarrow \Delta k = \frac{\Delta p}{\hbar}$$

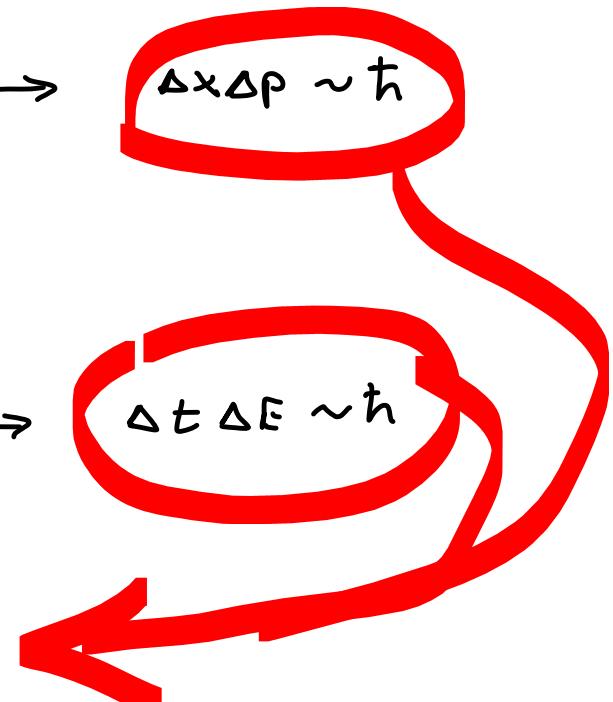
$\hbar\text{-bar} \equiv \frac{h}{2\pi}$

$$\Delta x \Delta k \sim 1 \rightarrow \Delta x \frac{\Delta p}{\hbar} \sim 1 \rightarrow \boxed{\Delta x \Delta p \sim \hbar}$$

$$E = h\nu = \hbar\omega \rightarrow \Delta\omega = \frac{\Delta E}{\hbar}$$

$$\Delta t \Delta \omega \sim 1 \rightarrow \Delta t \frac{\Delta E}{\hbar} \sim 1 \rightarrow \boxed{\Delta t \Delta E \sim \hbar}$$

Two forms of  
Heisenberg's Uncertainty Principle



(7)

In world of Newton and Particles 'R Particles:

Know  $\begin{pmatrix} \text{Position} \\ \text{Momentum} \\ \text{Forces} \end{pmatrix}$   Know the future

Universe is deterministic !

Not so in world of Particles 'R Waves

The better I determine the position  
The less well I know the momentum  
+ vice versa

Inherent Uncertainty !

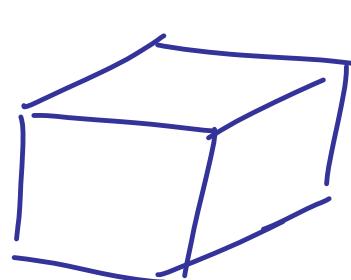


⑧

Mourn the loss of Deterministic universe !

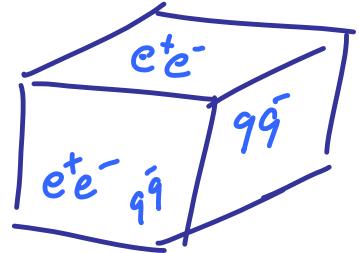
Rejoice in the New freedom !

The uncertainty Principle  
opens up Harry Potterish possibilities



a box of  
Empty space  
vacuum

①



a box of  
Empty Space  
Vacuum

Consider this bit of empty space over a very short time  
Energy over that time is very uncertain

Can have fluctuations in the energy  $\rightsquigarrow$  Particle-Antiparticle  
Pairs

Virtual particles, Fleeting existence

The vacuum is a seething mess of virtual particles



## Quantum Mechanics

(10)

recall

$$\vec{F} = -\vec{\nabla} V$$

Need a wave equation that allows us to calculate the motion of a particle under the influence of a force (Moving in a "potential")

Why not use  $\vec{\nabla}^2 \psi = \frac{1}{c^2} \frac{\partial^2 \psi}{\partial t^2}$  ?

This is for free traveling wave (no force)

want: 

- consistency w/  $E = h\nu$ ,  $P = h/\lambda$

Also need consistency w/  $E = h\nu$   
 $P = h/\lambda$

- Energy conservation including potential

$$E = \frac{P^2}{2M} + V$$

- Linear in  $\psi(x,t)$  in order to Satisfy Superposition

$$\begin{array}{l} \textcircled{A} \quad E = h\nu = \hbar\omega \\ \textcircled{B} \quad p = \frac{h}{\lambda} = k\hbar \end{array} \xrightarrow{\quad} \textcircled{C} \quad E = \frac{p^2}{2m} + V \quad \curvearrowright \quad \textcircled{D} \quad \hbar\omega = \frac{k^2\hbar^2}{2m} + \nu$$

Let's assume particle is a 1-d wave, Amplitude = 1

$$\textcircled{D} \quad \psi(x, t) = e^{i(kx - \omega t)}$$

→

$$-\frac{\hbar}{i} \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi$$

← Mult by  $i/i$

1-d  
nonrelativistic  
time-dependent  
Schrödinger  
equation

$$i\hbar \frac{\partial \psi(x, t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x, t)}{\partial x^2} + V\psi(x, t)$$

Two of the  
Big players  
in the early  
development of  
QM

## Werner Karl Heisenberg (1901 - 1976)

Nobel Prize in physics - 1932  
for "the creation of quantum  
Mechanics"

(Max Born, Pascual Jordan - co-workers)



(12)

Matrix  
formulation



## Erwin Rudolf Josef Alexander Schrödinger (1887 - 1961) Austria

1933 Nobel Prize in physics

1926 - Paper on Wave Mechanics of Matter  
*Annalen der Physik*

"for discovery of new and productive forms of  
atomic theory"

Wave equation  
formulation

$$-i\hbar \frac{\partial \psi_{(x,t)}}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi_{(x,t)}}{\partial x^2} + V(x,t) \psi_{(x,t)}$$

What do you do with this?  
 Put in the physical situation  $V(x,t)$   
 Solve for  $\psi_{(x,t)}$  and Energy

]

(13)

Will do Examples shortly  
but let me carry on for a bit with more background/theory

Can imagine it is useful to consider Static Situations

$$\text{let } V(x,t) \rightarrow V(x) \quad \text{Potential is constant in time}$$

$$\psi_{(x,t)} \rightarrow \psi_{(x)} \psi_{(t)} \quad \text{Assume Space + time dependence are Separable}$$

$$-i\hbar \frac{\partial \psi_{(x)} \psi_{(t)}}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi_{(x)} \psi_{(t)}}{\partial x^2} + V(x) \psi_{(x)} \psi_{(t)}$$

$$-i\hbar \psi_{(x)} \frac{\partial \psi_{(t)}}{\partial t} = -\frac{\hbar^2}{2m} \psi_{(t)} \frac{\partial^2 \psi_{(x)}}{\partial x^2} + V(x) \psi_{(x)} \psi_{(t)}$$

$$\div \text{ by } \psi(x)\psi(t) \quad -i\hbar \frac{1}{\psi(t)} \frac{\partial \psi(t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{1}{\psi(x)} \frac{\partial^2 \psi(x)}{\partial x^2} + V(x)$$

Function of  $t$  only      Function of  $x$  only

(14)

Equality only holds for all  $x$  and  $t$   
 if both sides are equal to same constant  
 $\rightarrow$  choose it to be  $E$

(A)

$$-i\hbar \frac{1}{\psi(t)} \frac{\partial \psi(t)}{\partial t} = E$$

one soln of this is  $\psi(t) = e^{-i\frac{E}{\hbar}t}$

And (B)  $-\frac{\hbar^2}{2m} \frac{1}{\psi(x)} \frac{\partial^2 \psi(x)}{\partial x^2} + V(x) = E$



for our  
1d wave  $\psi(x,t) = e^{i(kx - \omega t)}$  =  $\psi(x)\psi(t) = e^{ikx} e^{-i\omega t}$

recall  $E = \hbar\omega$

$$\psi(x,t) = \psi(x)e^{-i\frac{E}{\hbar}t}$$

$$-\frac{\hbar^2}{2m} \frac{1}{\psi(x)} \frac{\partial^2 \psi(x)}{\partial x^2} + V(x) = E$$

Time independent  
1d nonrelativistic  
Schrödinger eqn

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} + V(x) \psi(x) = E \psi(x)$$

(15)

Woohoo! We got wave equations . . .

(16)

But what's waving?

What is  $\Psi(x, t)$ ? it is complex (i.e., has  $i\sqrt{-1}$  in it  
as opposed to being complicated)  
So association of  $\Psi$  w/  
real world is NOT obvious

Max Born - 1926 - Born interpretation



Square of  $\Psi$  represents  
the probability density  
for the particle/wave

(17)

Born's postulate : at time  $t$ ,  $\psi^*(x,t) \psi(x,t) dx = |\psi(x,t)|^2 dx$   
 gives the probability of finding the particle  
 between  $x$  and  $x+dx$

Normalize  $\psi(x,t)$  so that the total probability is 1

$$\int_{-\infty}^{\infty} |\psi(x,t)|^2 dx = 1$$

The particle must be  
somewhere

in 3d ...  $\int |\psi(\vec{r},t)|^2 d\vec{v} = 1 \dots$

Are we having fun yet? Need to do a couple of examples

(18)

Example: 1d Free particle

Time ind.  
Schr. eqn

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} + V(x) \psi(x) = E \psi(x)$$

↙  
0

(A)

Try  $\psi(x, t) = e^{i(kx - \frac{E}{\hbar}t)}$  as a solution

$$\psi(x) = e^{ikx} \quad \frac{\hbar^2 k^2}{2m} e^{ikx} = E e^{ikx}$$

Works if  $k = \sqrt{\frac{2mE}{\hbar^2}}$

(19)

$\psi(x,t)$  here is a wave traveling in  $+x$  direction

$$\psi(x) = e^{-ikx} \text{ also a soln} \dots k = \sqrt{\frac{2ME}{\hbar^2}}$$

$\psi(x,t)$  here is a wave traveling in  $-x$  direction

so, general solution

$$\psi(x) = A e^{ikx} + B e^{-ikx}$$

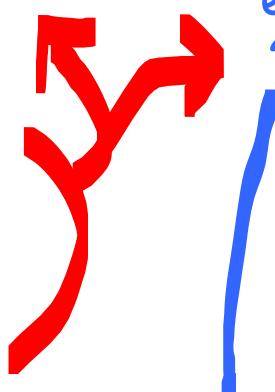
$$\text{where } k^2 = \frac{2ME}{\hbar^2}$$

or  
equivalently

$$\psi(x) = A' \sin(kx) + B' \cos(kx)$$

$$k^2 = \frac{2ME}{\hbar^2}$$

You will find that it is easiest to solve different types of problems using one form or the other



$$\frac{e^{i\theta} + e^{-i\theta}}{2}$$

$e^{i\theta} = \cos\theta + i\sin\theta$

recall Euler's formula

$$\frac{e^{i\theta} - e^{-i\theta}}{2}$$

20

$$A' \sin kx + B' \cos kx$$

$$\frac{A'}{2}(e^{ikx} - e^{-ikx}) + \frac{B'}{2}(e^{ikx} + e^{-ikx})$$

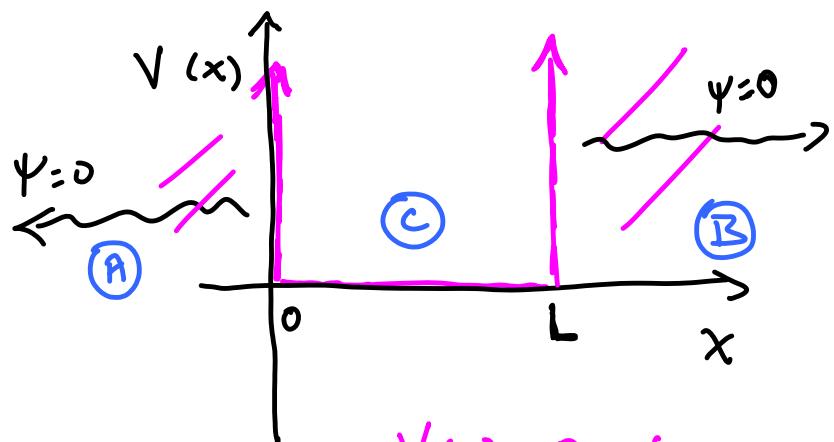
$$\left(\frac{A'}{2} + \frac{B'}{2}\right)e^{ikx} + \left(\frac{B'}{2} - \frac{A'}{2}\right)e^{-ikx}$$

$\equiv A$        $\equiv B$

# 1-d Particle in a box

(21)

infinite Square Well potential



$$V(x) = 0 \text{ for } 0 < x < L$$

$$V(x) = \infty \text{ for } x < 0, L < x$$

as  $V \rightarrow \infty$  Schr eqn makes no sense  
so  $\psi = 0$  there

for  $0 < x < L \quad V(x) = 0$   
particle is free  
inside the Box

$$\psi(x) = A \sin kx + B \cos kx$$

(22)

Use Boundary conditions to constrain Solution:

at  $x=0$ ,  $\psi(x)=0$

$$0 = A \sin kx + B \cos kx$$



$$\Rightarrow B = 0$$

$$\psi(x) = A \sin kx$$

at  $x=L$ ,  $\psi(x)=0$

$$0 = A \sin kL$$

$$kL = n\pi \quad n=1, 2, 3 \dots$$

$n=0$  not considered because for  $n=0$  either

$$\begin{aligned} k &= 0 & \psi(x) &= 0 \\ \text{or} \\ L &= 0 & \text{no box} \end{aligned}$$

$$\psi(x) = A \sin\left(\frac{n\pi}{L}x\right) \quad n=1, 2, 3 \dots$$

in region  $0 < x < L$

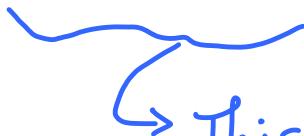
$$\textcircled{A} \quad k = \frac{n\pi}{L} \quad \text{so, } \textcircled{B} \quad E_n = \frac{n^2 \pi^2 \hbar^2}{L^2 2m} \quad \text{where } n=1, 2, 3 \dots \quad \textcircled{23}$$

energy is quantized! You'll find this is always true  
in attractive potentials

as in wanting to bind  
or capture

$$E_1 = \frac{\pi^2 \hbar^2}{L^2 2m}$$

$$E_2 = \frac{4\pi^2 \hbar^2}{L^2 2m} \quad \dots$$



This is the lowest energy allowed  
Note that it is NOT zero as we  
would have in classical physics

(24)

Determine A by using normalization  
to probability of 1

$$\int \psi(x)^* \psi(x) dx = 1 \quad \rightsquigarrow \quad \int_0^L A^2 \sin^2(kx) dx = 1 \quad (A)$$

Done  
in Giancoli:  
ex 38-6  
p. 1032

$$A^2 \int_0^L \sin^2(kx) dx = 1$$

$\underbrace{\qquad\qquad\qquad}_{= \frac{L}{2}}$

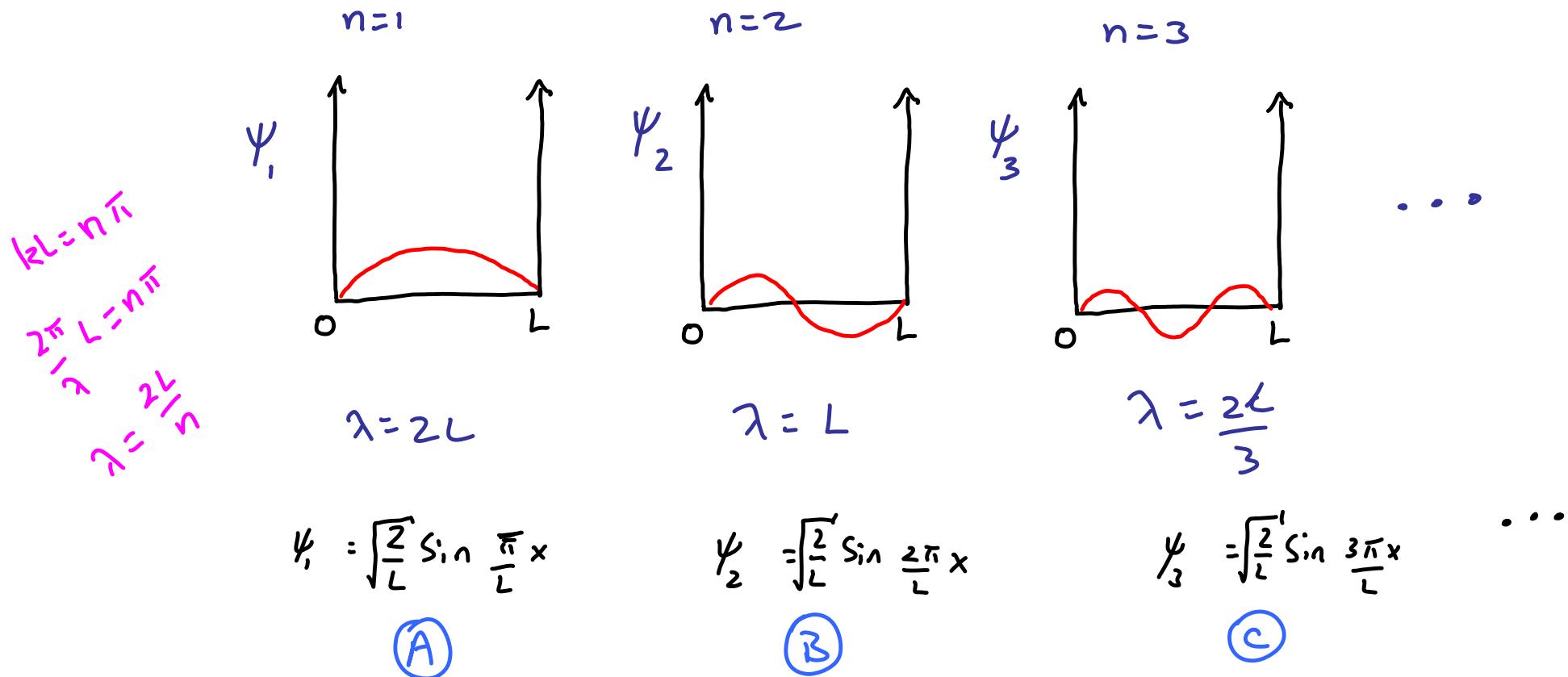
$$A = \sqrt{\frac{2}{L}}$$

1d  
square well  
wave function

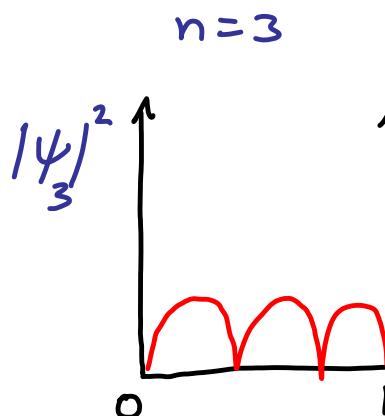
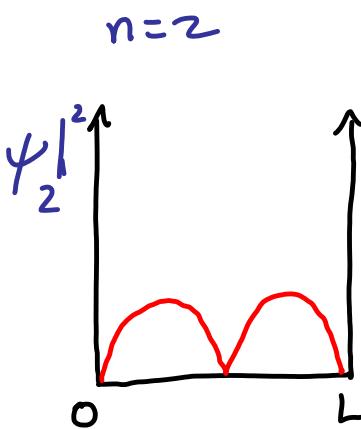
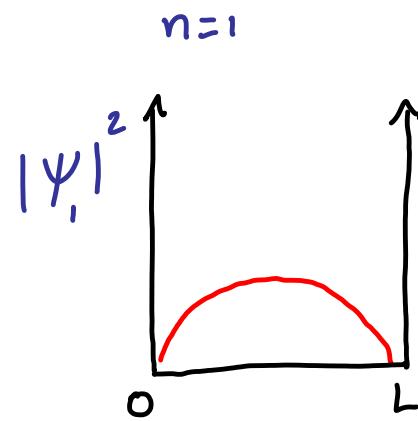
$$\psi_n(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi}{L}x\right) \quad \text{for } n=1, 2, 3, \dots \quad \text{in } 0 < x < L$$

What does the wavefunction look like?

(25)



what is the probability density for seeing particle at position  $x$  (26)



...

$$\lambda = 2L$$

$$|\psi_1|^2 = \frac{2}{L} \sin^2 \frac{\pi}{L} x$$

(A)

$$\lambda = L$$

$$|\psi_2|^2 = \frac{2}{L} \sin^2 \frac{2\pi}{L} x$$

(B)

$$\lambda = \frac{2L}{3}$$

$$|\psi_3|^2 = \frac{2}{L} \sin^2 \frac{3\pi}{L} x$$

(C)

...