

$\vec{S}_{EI} \sim \hat{r} \Rightarrow$  energy flows radially outwards

$\vec{S}_{EI} \sim \frac{1}{r^2} \Rightarrow$  energy conserved

$$\oint da \hat{n} \cdot \langle \vec{S}_{EI} \rangle = \text{constant for all } R$$

sphere  
radius  $R$

time averaged energy current

- Question - what about the  
 $\frac{1}{r^n}, n > 2$ , terms if we do not  
make radiation zone approx?

$$\langle \vec{S}_{EI} \rangle = \frac{1}{T} \int_0^T dt \vec{S}_{EI}(\vec{r}_I t)$$

$T$  is period  $T = \frac{2\pi}{\omega}$

$$= \frac{c}{8\pi} k^4 p_\omega^2 \frac{\sin^2 \theta}{r^2} \hat{r}$$

$$\langle \cos^2(\theta) \rangle = \frac{1}{2}$$

average energy flowing through an element  
of area at spherical angles  $\theta, \phi$  is

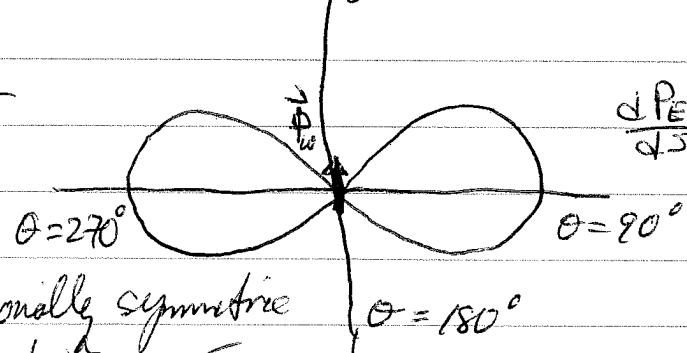
$$dP_{EI} = \hat{r} \cdot \underbrace{\langle \vec{S}_{EI} \rangle r^2 \sin \theta d\theta d\phi}_{\text{area of surface element}}$$

$$= r^2 d\Omega \quad \Omega \text{ is solid angle}$$

$$= \hat{r} \cdot \langle \vec{S}_{EI} \rangle r^2 d\Omega$$

$$\frac{dP_{EI}}{d\Omega} = \hat{r} \cdot \langle \vec{S}_{EI} \rangle r^2 = \frac{c}{8\pi} k^4 p_\omega^2 \sin^2 \theta \sim \omega^4 \sin^2 \theta$$

polar plot



rotationally symmetric  
about  $\hat{z}$  axis

$$\frac{dP_{EI}}{d\Omega} \sim \sin^2 \theta$$

most of power is  
directed outwards  
into plane  $\perp \vec{p}_\omega$ ,  
i.e. peaked about  $\theta = 90^\circ$

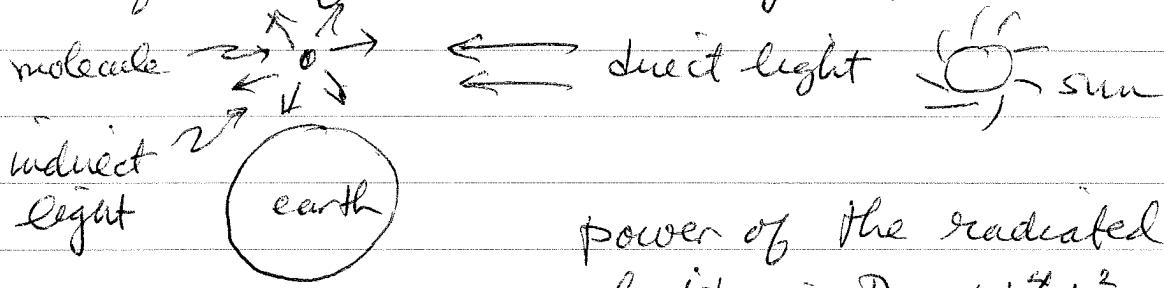
Total power radiated is

$$P_{EI} = \int \frac{dP_{EI}}{dS} dS = \frac{ck^4 p_0^2}{8\pi} 2\pi \underbrace{\int_0^\pi \sin \theta \sin^2 \theta}_{4/3} \pi$$

$$P_{EI} = \frac{ck^4 p_0^2}{3} = \frac{p_0^2 w^4}{3c^3} \sim w^4$$

why the sky is blue - Lord Rayleigh

When look up at sky, you are seeing the indirect light of the sun, which is the light emitted by the atoms and molecules of the atmosphere as they oscillate, and so radiate, due to the electric field of the direct light from the sun



power of the radiated indirect light is  $P \sim w^4 p_0^2$

$$\vec{P} = \alpha \vec{E} \quad \alpha \sim \frac{e^2}{m} \frac{1}{w_0^2 - w^2 - \epsilon \omega \gamma}$$

For molecules in atmosphere ( $N_2$  etc)  $w_0$  is typically at a freq higher than visible spectrum. Therefore, in visible spectrum  $\alpha \sim \frac{e^2}{m w_0^2}$  indep of  $w$ .

$\Rightarrow$  power radiated is  $P \sim w^4$

$P \sim w^4$  largest at high freq

Since light from sun is "white light"  
it has components of all freqs. Of these  
freqs, the higher ones are scattered the  
most & make up the indirect light we see.

Since blue is the largest  $\omega$  in visible spectrum,  
the sky is blue!

When we look at sunrise or sunset, we  
are looking at the direct rays of the sun.  
Since these rays are least scattered at  
low  $\omega \Rightarrow$  sunset and sunrise are red!

## Magnetic dipole approx - Radiation Zone for $r \gg 1$

$$\vec{A}_{M1} = \frac{e^{ikr}}{r} \left( \hat{r} - ik \right) \left( -\hat{r} \times \vec{m}_\omega \right)$$

$\approx ik \hat{r} \times \vec{m}_\omega \frac{e^{ikr}}{r}$  in RZ

$$\vec{B}_{M1} = \vec{\nabla} \times \vec{A}_{M1} = (\vec{\nabla} e^{ikr}) \times \left( ik \frac{\hat{r} \times \vec{m}_\omega}{r} \right)$$

$+ e^{ikr} \vec{\nabla} \times \left( \frac{i k \hat{r} \times \vec{m}_\omega}{r} \right)$

will give terms of  $\mathcal{O}(1/r^2)$   
so ignore in RZ approx

$$\boxed{\vec{B}_{M1} = -k^2 \frac{e^{ikr}}{r} \hat{r} \times (\hat{r} \times \vec{m}_\omega)}$$

From Ampere's Law

$$\vec{E}_{M1} = \frac{i}{k} \vec{\nabla} \times \vec{B}_{M1} = -ik (\vec{\nabla} e^{ikr}) \times \left( \hat{r} \times \left[ \hat{r} \times \vec{m}_\omega \right] \right)$$

$-ik e^{ikr} \vec{\nabla} \times \left( \hat{r} \times \left[ \hat{r} \times \vec{m}_\omega \right] \right)$

will give terms of  $\mathcal{O}(1/r^2)$   
so ignore in RZ approx

$$\vec{E}_{M1} = k^2 \frac{e^{ikr}}{r} \hat{r} \times (\hat{r} \times (\hat{r} \times \vec{m}_\omega))$$

triple product rule

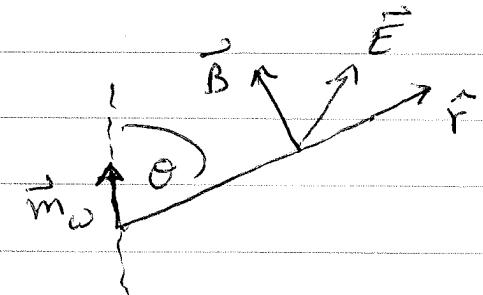
$$= k^2 \frac{e^{ikr}}{r} \left\{ \hat{r} [\hat{r} \cdot (\hat{r} \times \vec{m}_\omega)] - (\hat{r} \times \vec{m}_\omega) [\hat{r} \cdot \hat{r}] \right\}$$

$$\boxed{\vec{E}_{M1} = -k^2 \frac{e^{ikr}}{r} (\hat{r} \times \vec{m}_\omega)}$$

If align axes so that  $\vec{m}_\omega = m_\omega \hat{z}$  then

$$\vec{E}_{M1} = m_\omega \frac{k^2}{r} e^{ikr} \sin\theta \hat{\phi}$$

$$\vec{B}_{M1} = -m_\omega \frac{k^2}{r} e^{ikr} \sin\theta \hat{\theta}$$



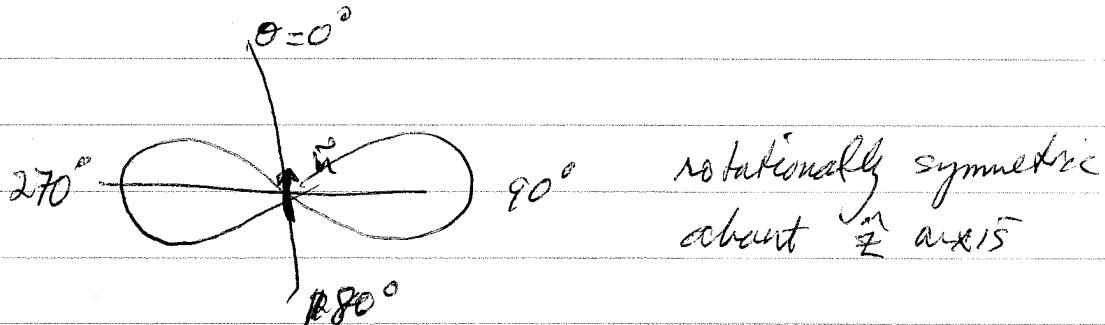
Poynting vector

$$\vec{S}_{M1} = \frac{c}{4\pi} \operatorname{Re}\{\vec{E}_{M1}\} \times \operatorname{Re}\{\vec{B}_{M1}\}$$

$$= \frac{c}{4\pi} \frac{k^4 m_\omega^2}{r^2} \cos^2(kr - \omega t) \sin^2\theta \hat{r}$$

$$\langle \vec{S}_{M1} \rangle = \frac{c}{8\pi} \frac{k^4 m_\omega^2}{r^2} \sin^2\theta \hat{r}$$

$$\frac{dP_{M1}}{dS2} = \hat{r} \cdot \langle \vec{S}_{M1} \rangle r^2 = \frac{c}{8\pi} \frac{k^4 m_\omega^2 \sin^2\theta}{r^2} \sim \omega^4 \sin^2\theta$$



$$P_{M1} = \int dS \frac{dP_{M1}}{dS} = \frac{c k^4 m_\omega^2}{3} = \frac{m_\omega^2 \omega^4}{3c^3}$$

$$\frac{P_{M1}}{P_{E1}} = \frac{m_\omega^2}{P_\omega^2} \sim \left(\frac{\omega}{c}\right)^2$$

$$m_\omega \sim \frac{df}{c}$$

$$P_\omega \sim \frac{df}{dg}$$

$$\sim dg \frac{v}{c}$$

## Electric Quadrupole radiation - radiation zone approx

$$\vec{A}_{E2} = \frac{e^{ikr}}{r} \left( \hat{r} - ik \right) \left( \frac{-i\omega}{60} \hat{r} \cdot \vec{Q}_w \right)$$

$$= -\frac{e^{ikr}}{r} \frac{k^3}{6} \hat{r} \cdot \vec{Q}_w \quad \text{in RZ approx}$$

$$\vec{B}_{E2} = \vec{\nabla} \times \vec{A}_{E2} = -(\vec{\nabla} e^{ikr}) \times \left[ \frac{k^2 \hat{r} \cdot \vec{Q}_w}{6r} \right]$$

$$= e^{ikr} \vec{\nabla} \times \left[ \frac{k^2 \hat{r} \cdot \vec{Q}_w}{6r} \right]$$

$$\boxed{\vec{B}_{E2} = -ik^3 \frac{e^{ikr}}{6r} \hat{r} \times (\hat{r} \cdot \vec{Q}_w)} \quad O\left(\frac{1}{r^2}\right) \text{ so ignore in RZ approx}$$

$$\vec{E}_{E2} = \frac{i}{k} \vec{\nabla} \times \vec{B}_{E2} = k^2 (\vec{\nabla} e^{ikr}) \times \left[ \frac{\hat{r} \times (\hat{r} \cdot \vec{Q}_w)}{6r} \right]$$

$$+ k^2 e^{ikr} \vec{\nabla} \times \left[ \frac{\hat{r} \times (\hat{r} \cdot \vec{Q}_w)}{6r} \right]$$

$O\left(\frac{1}{r^2}\right)$  so ignore in RZ approx

$$\vec{E}_{E2} = ik^3 \frac{e^{ikr}}{6r} \hat{r} \times \left[ \hat{r} \times (\hat{r} \cdot \vec{Q}_w) \right]$$

triple product rule

$$= ik^3 \frac{e^{ikr}}{6r} \left\{ \hat{r} [\hat{r} \cdot (\hat{r} \cdot \vec{Q}_w)] - (\hat{r} \cdot \vec{Q}_w) [\hat{r} \cdot \hat{r}] \right\}$$

$$\boxed{\vec{E}_{E2} = ik^3 \frac{e^{ikr}}{6r} \left\{ \hat{r} (\hat{r} \cdot \vec{Q}_w \cdot \hat{r}) - (\hat{r} \cdot \vec{Q}_w) \right\}}$$

## Poynting vector

$$\vec{S}_{E2} = +\frac{c}{4\pi} \operatorname{Re} \{ \vec{E}_{E2} \} \times \operatorname{Re} \{ \vec{B}_{E2} \}$$

$$= \frac{(-k)^6}{36r^2} \sin^2(kr-wt) \left\{ \hat{r} (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r}) - (\hat{r} \cdot \vec{\mathbb{Q}}_w) \right\} \times \left[ \hat{r} \times (\hat{r} \cdot \vec{\mathbb{Q}}_w) \right]$$

$$= \frac{(-k)^6}{36r^2} \sin^2(kr-wt) \left\{ \hat{r} [\hat{r} (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r}) \cdot (\hat{r} \cdot \vec{\mathbb{Q}}_w)] - (\hat{r} \cdot \vec{\mathbb{Q}}_w) [\hat{r} (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r}) \cdot \hat{r}] \right.$$

$$- \hat{r} [(\hat{r} \cdot \vec{\mathbb{Q}}_w) \cdot (\hat{r} \cdot \vec{\mathbb{Q}}_w)] \\ + (\hat{r} \cdot \vec{\mathbb{Q}}_w) [(\hat{r} \cdot \vec{\mathbb{Q}}_w) \cdot \hat{r}] \right\}$$

$$= \frac{(-k)^6}{36r^2} \sin^2(kr-wt) \left\{ (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r})^2 \hat{r} - (\hat{r} \cdot \vec{\mathbb{Q}}_w) (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r}) \right. \\ \left. - (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \vec{\mathbb{Q}}_w \cdot \hat{r}) \hat{r} \cdot (\hat{r} \cdot \vec{\mathbb{Q}}_w) (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r}) \right\}$$

$$\vec{S}_{E2} = \frac{-ck^6}{4\pi 36r^2} \sin^2(kr-wt) \left\{ (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r})^2 - (\hat{r} \cdot \vec{\mathbb{Q}}_w)^2 \right\} \hat{r}$$

$$\langle \vec{S}_{E2} \rangle = \frac{-ck^6}{4\pi 72r^2} \left\{ (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r})^2 - (\hat{r} \cdot \vec{\mathbb{Q}}_w)^2 \right\} \hat{r}$$

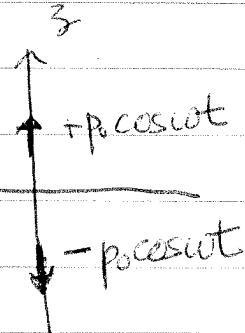
$$\frac{dP_{E2}}{d\Omega} = \hat{r} \cdot \langle \vec{S}_{E2} \rangle r^2 = \frac{ck^6}{4\pi 72} \left\{ (\hat{r} \cdot \vec{\mathbb{Q}}_w)^2 - (\hat{r} \cdot \vec{\mathbb{Q}}_w \cdot \hat{r})^2 \right\}$$

angular dependence of  $\frac{dP_{E2}}{d\Omega}$  depends  
on specific form of the tensor  $\vec{\mathbb{Q}}_w$

For example: suppose  $\Omega_{ij} = 0$  except for  $\Omega_{zz}$   
 $\Rightarrow \Omega_w = \Omega_{zz} \hat{z}\hat{z}$

$$(\hat{r} \cdot \hat{\Omega}_w \cdot \hat{r})^2 = (\Omega_{zz} \cos^2\theta)^2$$

$$(\hat{r} \cdot \hat{\Omega}_w)^2 = \Omega_{zz}^2 \cos^2\theta$$

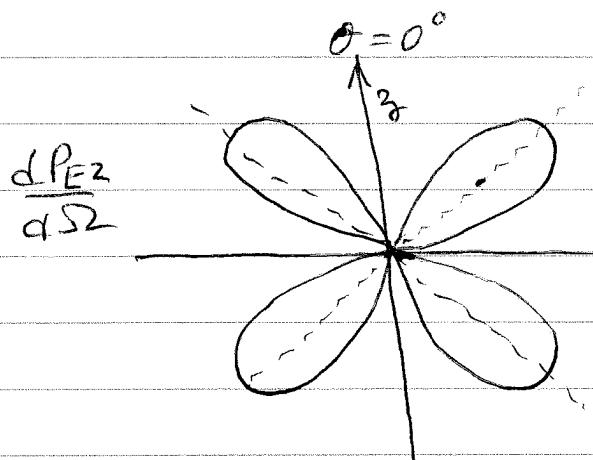


$$\frac{dP_{E2}}{dS2} = \frac{ck^6}{4\pi f_2} \Omega_{zz}^2 [\cos^2\theta - \cos^4\theta]$$

$$= \frac{ck^6}{4\pi f_2} \Omega_{zz}^2 \cos^2\theta \sin^2\theta$$

$$\cos\theta \sin\theta = \frac{1}{2} \sin 2\theta$$

$$= \frac{ck^6}{4\pi f_2 288} \Omega_{zz}^2 \sin^2 2\theta$$



peak at  $45^\circ$

$\theta = 90^\circ$  rotationally invariant  
about  $\hat{z}$  axis

$$\frac{P_{E2}}{P_{E1}} \sim \frac{k^6 Q^2}{k^4 p^2} \sim \frac{k^2 (qd^2)^2}{(qd)^2} \sim k^2 d^2 \sim \left(\frac{v}{c}\right)^2$$

$$P_{E2} \sim P_{M1}$$

For more general case, choose axes so that  $\overset{\leftarrow}{Q_w}$  is diagonal - can always do this since  $\overset{\leftarrow}{Q_w}$  is symmetric

$$(\overset{\leftarrow}{r} \cdot \overset{\leftarrow}{Q_w} \cdot \overset{\leftarrow}{r}) = \overset{\leftarrow}{r} \cdot \begin{pmatrix} Q_{xx} & 0 & 0 \\ 0 & Q_{yy} & 0 \\ 0 & 0 & Q_{zz} \end{pmatrix} \cdot \overset{\leftarrow}{r}$$

$$= \overset{\leftarrow}{r} \cdot \begin{pmatrix} Q_{xx} \sin\theta \cos\varphi \\ Q_{yy} \sin\theta \sin\varphi \\ Q_{zz} \cos\theta \end{pmatrix} = Q_{xx} \sin^2\theta \cos^2\varphi + Q_{yy} \sin^2\theta \sin^2\varphi + Q_{zz} \cos^2\theta$$

$$(\overset{\leftarrow}{r} \cdot \overset{\leftarrow}{Q_w})^2 = Q_{xx}^2 \sin^2\theta \cos^2\varphi + Q_{yy}^2 \sin^2\theta \sin^2\varphi + Q_{zz}^2 \cos^2\theta$$

$$\frac{dP_{E2}}{d\Omega} = \frac{k^6}{f^2} \left\{ Q_{zz}^2 (\cos^3\theta - \cos^4\theta) + Q_{xx}^2 (\sin^2\theta \cos^2\varphi - \sin^4\theta \cos^4\varphi) + Q_{yy}^2 (\sin^2\theta \sin^2\varphi - \sin^4\theta \sin^4\varphi) \right\}$$

$$\frac{dP_{E2}}{d\Omega} = \frac{k^6}{f^2} \left\{ Q_{zz}^2 \cos^2\theta \sin^2\theta + Q_{xx}^2 \sin^2\theta \cos^2\theta (1 - \sin^2\theta \cos^2\varphi) + Q_{yy}^2 \sin^2\theta \sin^2\theta (1 - \sin^2\theta \sin^2\varphi) \right\}$$

no special symmetries - varies with  $\theta$  and  $\varphi$

For arbitrary charge distributions - not pure harmonic

For  $\vec{p}_\omega e^{-i\omega t}$  pure harmonic oscillation, we found the radiated fields in electric dipole approx one

$$\vec{E} = \vec{E}_\omega e^{-i\omega t}, \quad \vec{B} = \vec{B}_\omega e^{-i\omega t}$$

$$\vec{E}_\omega = -k^2 \frac{e^{-ikr}}{r} \hat{r} \times (\hat{r} \times \vec{p}_\omega) = -\frac{\omega^2}{c^2} \frac{e^{i\omega r/c}}{r} \hat{r} \times (\hat{r} \times \vec{p}_\omega)$$

$$\vec{B}_\omega = k^2 \frac{e^{-ikr}}{r} (\hat{r} \times \vec{p}_\omega) = \frac{\omega^2}{c^2} \frac{e^{i\omega r/c}}{r} (\hat{r} \times \vec{p}_\omega)$$

$$\text{as } k = \frac{\omega}{c}$$

For an arbitrarily time varying charge distribution with electric dipole moment

$$\vec{p}(t) = \int \frac{d\omega}{2\pi} \vec{p}_\omega e^{-i\omega t}$$

Then solution for fields given by superposition

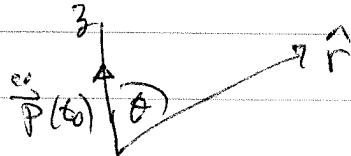
$$\begin{aligned} \vec{E}(\vec{r}, t) &= \int \frac{d\omega}{2\pi} \vec{E}_\omega e^{-i\omega t} \\ &= - \int \frac{d\omega}{2\pi} \frac{e^{-i\omega(t-r/c)}}{r} \left( \frac{\omega^2}{c^2} \right) \hat{r} \times (\hat{r} \times \vec{p}_\omega) \\ &= -\frac{1}{c^2 r} \hat{r} \times \left[ \hat{r} \times \int \frac{d\omega}{2\pi} e^{-i\omega(t-r/c)} \vec{p}_\omega \omega^2 \right] \\ &= \frac{1}{c^2 r} \hat{r} \times \left[ \hat{r} \times \frac{\partial^2}{\partial t^2} \int \frac{d\omega}{2\pi} e^{-i\omega(t-r/c)} \vec{p}_\omega \right] \end{aligned}$$

$$\boxed{\vec{E}(\vec{r}, t) = \frac{1}{c^2 r} \hat{r} \times [\hat{r} \times \ddot{\vec{p}}(t - r/c)]} \quad \ddot{\vec{p}} = \frac{d^2 \vec{p}}{dt^2}$$

define  $t_0 \equiv t - r/c$  = "retarded time"

in spherical coords, if  $\ddot{\vec{p}}(t_0)$  is along  $\hat{z}$

$$\vec{E}(\vec{r}, t) = \frac{\ddot{p}(t_0) \sin \theta}{c^2 r} \hat{\theta}$$



Similarly

$$\vec{B}(\vec{r}, t) = \int \frac{d\omega}{2\pi} \vec{B}_\omega e^{-i\omega t}$$

$$= \int \frac{d\omega}{2\pi} \frac{e^{-i\omega(t-r/c)}}{r} \left(\frac{\omega^2}{c^2}\right) (\hat{r} \times \vec{p}_\omega)$$

$$= \frac{-1}{c^2 r} \hat{r} \times \frac{\partial^2}{\partial t^2} \int \frac{d\omega}{2\pi} e^{-i\omega(t-r/c)} \vec{p}_\omega$$

$$\boxed{\vec{B}(\vec{r}, t) = -\frac{1}{c^2 r} \hat{r} \times \ddot{\vec{p}}(t_0)}$$

$$\vec{B}(\vec{r}, t) = \frac{\ddot{p}(t_0)}{c^2 r} \sin \theta \hat{\phi} \quad \text{in spherical coords}$$

Pointing vector

$$\vec{s} = \frac{c}{4\pi} \vec{E} \times \vec{B} = \frac{c}{4\pi} \left(\frac{1}{c^2 r}\right)^2 [\ddot{p}(t_0)]^2 \sin^2 \theta \hat{r}$$

Total power radiated through a sphere of radius  $r$  is

$$\begin{aligned} \Phi &= \oint d\Omega \hat{r} \cdot \vec{S} = 2\pi \int_0^\pi d\theta \sin\theta r^2 \hat{r} \cdot \vec{S} \\ &= \frac{[\ddot{P}(t_0)]^2}{2c^3} \underbrace{\int_0^\pi d\theta \sin^3\theta}_{4/3} \end{aligned}$$

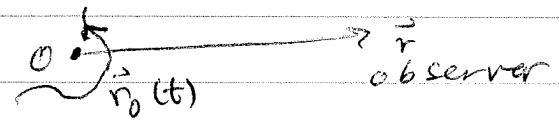
$$P = \frac{2 [\ddot{P}(t_0)]^2}{3c^3}$$

For a point charge moving along a trajectory  $\vec{r}_0(t)$

$$\vec{F}(t) = q \vec{r}_0(t)$$

$$\ddot{\vec{F}}(t) = q \ddot{\vec{r}}_0(t) = q \vec{a}(t)$$

$\vec{a}$  acceleration



$$P = \frac{2}{3} \frac{q^2 \dot{a}^2(t_0)}{c^3}$$

Larmor's formula

← total power passing through  
a sphere of radius  $r$  at time  $t$   
is due to acceleration at retarded  
time  $t_0 = t - r/c$

power radiated  $\propto (\text{acceleration})^2$

Larmor's formula above only holds in the  
non-relativistic limit since it is based on  
the electric dipole approx.