PHY 415 Solutions Problem Set 6 1) Top view of the wires $\hat{\Psi}_{1} \ll$ At the pant (x=0, y, z) on the plane that bisects the space between the Two wries, $\frac{\vec{B} = 2I_1}{C\sqrt{\left(\frac{d}{2}\right)^2 + y^2}} + \frac{\vec{\phi}}{C\sqrt{\left(\frac{d}{2}\right)^2 + y^2}} + \frac{2I_2}{C\sqrt{\left(\frac{d}{2}\right)^2 + y^2}} + \frac{\vec{\phi}}{C\sqrt{\left(\frac{d}{2}\right)^2 + y^2}}$ * X 2 wries out of plane of page where \$, ad \$, are the potor angle mit vectors with respect to when I ad 2 (see Lingian) $\hat{q}_1 = \cos \theta \hat{q} - \sin \theta \hat{x}$ $\hat{q}_2 = -\cos \theta \hat{q} - \sin \theta \hat{x}$ $\frac{cos Q = \frac{d/2}{\sqrt{(4_2)^2 + y^2}}$ $sm\theta = \frac{y}{\sqrt{\left(\frac{d}{2}\right)^2 + y^2}}$ So an the 18 plane caparaty the two wies $\vec{B}(y) = 2\cos\Theta \quad (I_1 - I_2)\hat{y} - 2\sin\Theta \quad (I_1 + I_2)\hat{x}$ $C\sqrt{\left(\frac{d}{2}\right)^2 + y^2} \quad C\sqrt{\left(\frac{d}{2}\right)^2 + y^2}$ $= \underbrace{d(I_1 - I_2)}_{c\left[\left(\frac{d}{2}\right)^2 + y^2\right]} \underbrace{g - zy(I_1 + I_2)}_{c\left[\left(\frac{d}{2}\right)^2 + y^2\right]}$ Ĩ. when I,= I, 1st tem vanishes! B = Bx when I, = - I2, and tens vanishes! B = BG

Maxweel stress times When E=0 Tig = 4TT { B, B, -280, B²} we want - Sda F. m = - Sdz Sdy T. x T.X = Txx X + Tyx & + T3x 3 $S_{mode} \vec{B} = \frac{d(T_1 - T_2)}{c[y^2 + (d_2)^2]} \vec{y} - \frac{2y(T_1 + T_2)}{c[y^2 + (d_2)^2]} \vec{x}$ has only x and y components, then Txz=Tyz=0 $T_{ij} = \frac{1}{4\pi} \left(\frac{1}{2} B_x - \frac{1}{2} B_y B_x B_y \right)$ $B_{\gamma}B_{\chi}$ $\pm B_{\gamma}^2 - \pm B_{\chi}^2$ д 0 0 $-\frac{1}{2}(B_x + B_y)$ Consider the time (dy Tyx $\int \frac{dy}{dy} = \frac{1}{4\pi} \int \frac{dy}{dy} = \frac{1}{4\pi} \int \frac{dy}{dy} = \frac{1}{4\pi} \int \frac{dy}{dy} = \frac{d(T_1 - T_2) + y(T_1 + T_2)}{C^2 \left[\frac{y^2}{2} + \frac{dy}{2} \right]^2}$ = 0 since integraid is odd in 14. So the only tem that and betes to Sola Tom is X Sda TXX

Sda Txx = stifdz fdg [Bx - By] $= \frac{1}{8\pi^{2}} \int dy \left[\frac{4y^{2}(I_{1}+I_{2})^{L}}{(y^{2}+(d_{1})^{2})^{2}} - \frac{d^{2}(I_{1}-I_{2})^{L}}{(y^{2}+(d_{2})^{2})^{2}} \right]$ prig substitution integrals do do the æ verlables $\sqrt{\frac{y^2}{2}} \frac{y^2}{\left(\frac{y}{2}\right)^2} = \frac{y^2}{\left[\frac{y^2}{2} + \frac{y^2}{2}\right]^2}$ $\frac{1}{[4^{2}+(2)^{2}]} = (\frac{2}{d})^{2} \cos^{2}\theta$ $y = \frac{d}{2} \tan \theta \implies dy = \frac{d}{2} \frac{d \tan \theta}{d\theta} d\theta = \frac{d}{2} \frac{d\theta}{\cos^2 \theta}$ $\int da T_{XX} = \frac{1}{9\pi c} \int d3 \int d0 \frac{d}{2\cos^2 0} \left[4(T_1 + T_2)^2 \sin^2 0 \frac{t^2}{d} \right]^2 \cos^2 0$ $-\frac{\pi}{2} \int d0 \frac{d}{2\cos^2 0} \left[4(T_1 + T_2)^2 \sin^2 0 \frac{t^2}{d} \right]^2 \cos^2 0$ $-d^{2}(I_{l}-I_{z})^{2}(z_{d})^{4}cos^{2}b$ $= \frac{1}{8\pi c^2 d} \int dz \int d\sigma \int (f_i + f_2)^2 \sin^2 \theta$ $-(I_1-I_2)^2\cos^2\theta$ $= \frac{1}{\pi dc} \int_{2}^{2} dz = \int_{2}^{2} \left((I_{1} + I_{2})^{2} - (I_{1} - I_{2})^{2} \right)$ $= \frac{L_{3}}{1!} \left[I_{1}^{2} + I_{2}^{2} + 2I_{1}I_{2} - (I_{1}^{2} + I_{2}^{2} - 2I_{1}I_{2}) \right]$ $\frac{2L_3}{L_2}I_1I_2$

So $\int -\int da = -2 L_3 \frac{T_1 T_2}{dc^2} \hat{\chi}$ yer all values of I, ad Iz The above is exactly equal to the magnetic static force on come 2 due to the magnetic field of unie 1. $\overline{F}_2 = h_3 \overline{I_2} \times \overline{B},$ d $= L_{3} \frac{J_{2}}{C} \hat{x} \left(\frac{2J_{1}}{dC} \hat{y} \right)$ $= -2 L_3 I_1 I_2 \hat{x}$ 12 \bot If I, and Iz are in the same direction, the force is attractive. If I, ad Iz are in opposite directions, the force is repulsive - Ida T. m is the flux of electromagnetic momentum passing through the plane which is the total electronagnetic force on the place (force = momentas / the) pushing from side I to side 2. We would get the same result is we conjuted - Jala T. m over any y-3 plane mbetween The two whis (not just midway). We can view this like an electronregatic pressure that fills space and transmits the force from when to whe 2

2.1 2) <u>س حل</u> I've net change e iniformly distributed on ŝÎ surface. fust we find E + B fields. The above is an electrostatic and magnetostatic situation, since if =0 and T.J =0 $\vec{E}(\vec{r}) = \int_{r}^{r} e \frac{\hat{r}}{r^{2}}$ outside sphenical shell, r>R E looks just like pt r<R 0 charge at the origin From problem Set 5, problem 1(b), we had B = STOWR 2 rsr -here $\sigma = \frac{e}{4\pi R^2}$ $\frac{1}{B} = \frac{2eW}{3CR} \hat{z}$ $\vec{B} = \frac{4}{3c}\pi\sigma\omega R^4 = \frac{2\cos\sigma \hat{r} + \sin\sigma\hat{\sigma}}{r^3}$ rz R $= \frac{e \omega R^2}{3C r^3} \left[2 \cos \theta \hat{r} + \sin \theta \hat{\theta} \right]$ 100 This is just a cupole field with total dipole moment [m] = ewR = 4TWOR4

2.2 a) The total electromagnetic energy is $W = \frac{1}{8\pi} \int d^2 \left[E^2 + B^2 \right]$ Smile this is an electrostatic and magnetostatic schustion it is easier to me $W = \frac{1}{2} \int d^3r \, p \, \phi + \frac{1}{2c} \int d^3r \, \vec{j} \cdot \vec{A}$ sie then the integrals are confined to the shell r= R where f and j are not zero. We have for electrostatic potential $\phi(r) = \int_{r}^{e} r \partial R$ OR rsp Weller = $\frac{1}{2}\int d^3r \,\rho \,\phi = \frac{1}{2} 4\pi R^2 \sigma \,\phi(R) = \frac{1}{2} \frac{e^2}{R}$ For the magneto static vector potential we have $\vec{A}(\vec{r}) = \frac{\vec{m} \times \hat{r}}{r^2} = \frac{m \sin \theta}{r^2} \hat{\varphi} \quad r \ge R$ since \vec{B} is a pune dipole field. The magnetic dipole moment is $\vec{m} = \frac{4\pi}{3c} \omega \sigma R^4$

$$\begin{split} \vec{A}(r=R) &= \prod_{3c} \omega \, \mathcal{G} \, R^2 \, \sin \theta \, \hat{\phi} \\ W_{mey} &= \frac{1}{2c} \int d^3r \, \vec{j} \cdot \vec{A} = \frac{1}{2} \int da \quad \vec{K} \cdot \vec{A} \\ \vec{K} &= \sigma \, \omega \, R \, \sin \theta \, \hat{\phi} \quad (see \ \rho \, \omega \, \ell \ / a, \, Sct 5 \,) \\ W_{mey} &= \frac{1}{2c} \int da \quad \frac{4\pi}{3c} \, \omega^2 \sigma^2 R^3 \, \int d\theta \, \sin \theta \, \sin^2 \theta \\ &= 2\pi \, R^2 \, \frac{4\pi}{3c} \, \omega^2 \sigma^2 R^3 \, \int d\theta \, \sin \theta \, \sin^2 \theta \\ &= 2\pi \, R^2 \, \omega^2 \sigma^2 R^5 \, \int \theta \, \sin \theta \, \left[1 - \omega \, s^3 \theta \right] \\ &= \left[- \omega \, s \, \theta \, + \frac{\cos^3 \theta}{3} \right]_{0}^{R} \\ &= \left(1 - \frac{1}{3} \right) - \left(-1 + \frac{1}{3} \right) = \frac{4}{3} \\ W_{mag} &= \left(\frac{4\pi}{9c^2} \, \omega^2 \, \sigma^3 R^5 \, = \left(\frac{4\pi}{9c^2} \, \omega^2 \, \frac{e^2 \, R^5}{9c^2} \, R^5 \right) \\ &W &= W \, dec + W \, aog \\ \hline W &= \frac{1}{2} \frac{e^2}{R} \left[1 + \frac{2}{q} \, \frac{R^2 \, \omega^2}{c^2} \, \right] \end{split}$$

2.3

5) Angula monontum density $\vec{J} = \vec{r} \times T = \frac{1}{4TC} \vec{r} \times (\vec{e} \times \vec{B})$ d = 0 miside shell since $\vec{E} = 0$ miside outside shell $\vec{E} \times \vec{B} = \underbrace{e}_{r^2} \hat{r} \times \left[\underbrace{e \omega k^2}_{3 C r^3} (2 \omega s \Theta \hat{r} + s m \Theta \hat{O} \right]$ $= \frac{e^2 \omega R^2}{2 \sigma r^5} \sin \theta \hat{\phi}$ as $\hat{\tau} \times \hat{\theta} = \hat{\varphi}$ $\overline{\int}_{4\pi c}^{2} = \frac{1}{4\pi c} \widehat{r} \times (\overline{E} \times \overline{B}) = -\frac{e^{2} \omega R^{2}}{12\pi c^{2} r^{4}} \sin \theta \hat{\theta} \cos \hat{r} \times \hat{q} = \hat{\theta}$ $\vec{J} = -\frac{e^2 w R^2}{12 \pi c^2 r^4} \sin \theta \hat{\theta} r \hat{\gamma} R, \quad J = 0, \quad r \leq R$ integrate over all space to get total angula momentum $\overline{L} = \int d^{3}r \, \overline{d} = 2\pi \int d\theta \sin \theta \int dr r^{2} \, \overline{d}$ By symmetry we only need to integrate Iz since other components will vanish.

2.4

$$\hat{\theta} \cdot \hat{g} = -\sin \theta \implies J_{g} = \frac{e^{2} \omega R^{2}}{12 \pi c^{2} r^{4}} \sin \theta$$

$$L_{g} = \frac{e^{2} \omega R^{2}}{12 \pi c^{2}} 2\pi \int_{0}^{\pi} \frac{d\theta \sin^{3}\theta}{d\theta \sin^{3}\theta} \int_{0}^{\infty} dr \frac{1}{r^{2}}$$

$$R \stackrel{\text{since } d=0}{r} \frac{de^{2} \omega R}{r^{2}}$$

$$\frac{u}{3} \stackrel{\text{l}}{r}$$

$$L_{g} = \frac{2}{q} \frac{e^{2} \omega R}{c^{2}}$$

2.5

c)

 $L_{z} = \frac{z}{q} \frac{e^{2} \omega R}{c^{2}} = \frac{\pi}{2}$ $\Rightarrow \frac{\omega R}{c} = \frac{q}{4} \frac{\pi c}{\rho^2}$ Use $\frac{hc}{a^2} = 137 = \frac{1}{\alpha}$ fine structure constant $\frac{\omega R}{C} = \left(\frac{9}{4}\right)(137) = 308.25$ so the velocity at the equator is WR = 308.25 c much faster than the speed of light. Clearly the model count be correct.

2-6 d) $W = \frac{1}{2} \frac{e^2}{R} \left[1 + \frac{2}{q} \left(\frac{R\omega}{c} \right)^2 \right]$ $= \frac{1}{2} \frac{e^{2}}{R} \left[\left(+ \frac{2}{q} \left(\frac{q}{4} \right)^{2} \left(137 \right)^{2} \right] \right]$ $W = \frac{1}{2} \frac{e^2}{R} \left[2.1116 \times 10^4 \right] = mc^2$ $R = \frac{1}{2} \frac{e^2}{m_0} 2 \left(2.116 \times 10^4 \right)$ if as is Bohr radios $\frac{R}{q_0} = \frac{1}{2} \frac{e^2}{q_0} \frac{1}{mr^2} \left(2.1116 \times 10^4 \right)$ $u_{e} = \frac{e^2}{2a_0} = 13.6 ev$, $mc^2 = 0.511 \times 10^6 eV$ $\frac{R}{a_0} = \frac{(13.6)(2.1116 \times 10^4)}{0.511 \times 10^6} = \left[0.562 - \frac{R}{a_0}\right]$ electron radius is n' Bohn soudin! Clearly not convect! Much for big! Puttery in numbers : R = (0.562)(0.529×10° cm) = 0.297 Å $\omega = (308.25)(3\times10^{10} \text{ om/sec}) = 3.11\times10^{10} \text{ rads}$ (0.297×10-300) sec

Problem Set 8, problem 2

In my solution to problem 2 on Problem Set 8, I computed the electromagnetic energy of the spinning charged sphere,

$$W = \frac{1}{8\pi} \int d^3r \left[|\mathbf{E}|^2 + |\mathbf{B}|^2 \right] \tag{1}$$

as

$$W = \frac{1}{2} \int d^3 r \,\rho \phi + \frac{1}{2c} \int d^2 r \,\mathbf{j} \cdot \mathbf{A} \tag{2}$$

which should be correct in an electrostatic and magnetostatic situation such as we have in the problem.

I then wrote the above as,

$$W = \frac{1}{2} \oint_{S} da \,\sigma \phi + \frac{1}{2c} \oint_{S} da \,\mathbf{K} \cdot \mathbf{A} \qquad \text{with } S \text{ the surface of the sphere of radius } R \tag{3}$$

since all the charge and current are restricted to the surface of the sphere. Here $\sigma = e/(4\pi R^2)$ is the surface charge density, and $\mathbf{K} = \sigma \omega R \sin \theta \hat{\boldsymbol{\varphi}}$ is the surface current density.

But you may have wondered, is that OK to do? When I compute the above surface integrals, should I use the values of ϕ and **A** just on the outside of the surface of the sphere, or should I use their values just on the inside of the surface of the sphere? There could possibly be a difference if either ϕ or **A** were discontinuous at the surface.

The answer is, YES! it is OK to do. One reason is that one can choose potentials ϕ and **A** that are continuous at the surface S, so it does not matter if we use the values just outside or just inside. We can also directly compute the energy W using Eq. (1) and confirm we get the same answer.

Electrostatic Energy

For the electrostatic part of the problem, we know that ϕ can be chosen to be continuous at the surface. The electric field due to the uniform surface charge on the sphere is,

$$\mathbf{E} = \begin{cases} \frac{q}{r^2} \mathbf{\hat{r}} & r > R \quad \text{outside} \\ 0 & r < R \quad \text{inside} \end{cases}$$
(4)

and we can take as the potential

$$\phi = \begin{cases} \frac{q}{r} & r > R \text{ outside} \\ \frac{q}{R} & r < R \text{ inside} \end{cases}$$
(5)

You can confirm that the above satisfies $\mathbf{E} = -\nabla \phi$. Since $\phi(r)$ is continuous at r = R, there is no problem defining the integral $\oint da \, \sigma \phi$. But we can also explicitly compute W_{elec} from \mathbf{E} ,

$$W_{\text{elec}} = \frac{1}{8\pi} \int d^2 r \, |\mathbf{E}|^2 = \frac{1}{8\pi} \, 4\pi \int_R^\infty dr \, r^2 E^2 = \frac{q^2}{2} \int_0^\infty dr \, \frac{1}{r^2} = \frac{q^2}{2R} \tag{6}$$

which is exactly the same as we found from $\frac{1}{2} \oint da \, \sigma \phi$.

Magnetostatic Energy

For the magnetostatic part of the problem, we have the magnetic field from problem 1 of Problem Set 5,

$$\mathbf{B} = \begin{cases} \frac{m}{r^3} [2\cos\theta \,\hat{\mathbf{r}} + \sin\theta \,\hat{\theta}] & r > R \quad \text{outside} \\ \frac{2m}{R^3} \,\hat{\mathbf{z}} & r < R \quad \text{inside} \end{cases}$$
(7)

where the magnetic dipole moment is $\mathbf{m} = m\hat{\mathbf{z}}$ with $m = \frac{e\omega R^2}{3c}$. Outside the magnetic field is just like a point magnetic dipole with dipole moment \mathbf{m} , while inside the magnetic field is constant.

The vector potential for the field outside is therefore just the dipole vector potential,

$$\mathbf{A}^{\text{out}} = \frac{\mathbf{m} \times \hat{\mathbf{r}}}{r^3} = \frac{m \sin \theta}{r^2} \,\hat{\mathbf{\phi}} \tag{8}$$

To get the vector potential inside, we note that a constant magnetic field $B\hat{z}$ can be given by the potential

$$\mathbf{A} = \frac{B}{2} [x \hat{\mathbf{y}} - y \hat{\mathbf{x}}] \qquad \Rightarrow \qquad \boldsymbol{\nabla} \times \mathbf{A} = B \hat{\mathbf{z}}$$
⁽⁹⁾

Now $\hat{\mathbf{z}} \times \mathbf{r} = \hat{\mathbf{z}} \times (x\hat{\mathbf{x}} + y\hat{\mathbf{y}} + z\hat{\mathbf{z}}) = (x\hat{\mathbf{y}} - y\hat{\mathbf{x}})$, so we can rewrite the above as, $\mathbf{A} = \frac{B}{2}\hat{\mathbf{z}} \times \mathbf{r} = \frac{Br\sin\theta}{2}\hat{\mathbf{\phi}}$. So we therefore have,

$$\mathbf{A}^{\rm in} = \frac{2m}{R^3} \, \frac{r \sin \theta}{2} \, \hat{\boldsymbol{\varphi}} = \frac{mr \sin \theta}{R^3} \, \hat{\boldsymbol{\varphi}} \tag{10}$$

Evaluating on the surface at r = R we have,

$$\mathbf{A}^{\text{out}} = \frac{m\sin\theta}{R^2}\,\hat{\boldsymbol{\varphi}}, \qquad \mathbf{A}^{\text{in}} = \frac{mR\sin\theta}{R^3}\,\hat{\boldsymbol{\varphi}} = \frac{m\sin\theta}{R^2}\,\hat{\boldsymbol{\varphi}} \tag{11}$$

and so on the surface of the sphere $\mathbf{A}^{\text{out}} = \mathbf{A}^{\text{in}}$ and the vector potential is continuous. Hence there is no problem defining the integral $\oint da \mathbf{K} \cdot \mathbf{A}$. You might ask, what if we chose a different form for \mathbf{A}^{in} than that of Eq. (9). But we have shown in an earlier discussion question that $\int d^3r \mathbf{j} \cdot \mathbf{A}$ is independent of the gauge of \mathbf{A} , and one can show that the same is true of a surface term $\oint da \mathbf{K} \cdot \mathbf{A}$.

We can also check that we have the correct answer by directly computing W_{mag} from **B**.

$$W_{\rm mag} = \frac{1}{8\pi} \int d^3 r \, |\mathbf{B}|^2 = \frac{1}{8\pi} \int_{\rm outside} d^3 r \, |\mathbf{B}^{\rm out}|^2 + \frac{1}{8\pi} \int_{\rm inside} d^3 r \, |\mathbf{B}^{\rm in}|^2 \tag{12}$$

Since $\mathbf{B}^{in} = (2m/R^3)\hat{\mathbf{z}}$ is constant, the contribution from the inside is,

$$W_{\text{mag,in}} = \frac{1}{8\pi} \frac{4\pi R^3}{3} \left(\frac{2m}{R^3}\right)^2 = \frac{2m^2}{3R^3}$$
(13)

Since $\mathbf{B}^{\text{out}} = m(2\cos\theta\,\hat{\mathbf{r}} + \sin\theta\,\hat{\theta})/r^3$, the contribution from the outside is,

$$W_{\rm mag,out} = \frac{1}{8\pi} m^2 \int_0^{2\pi} d\varphi \int_0^{\pi} d\theta \sin \theta \int_R^{\infty} dr \, r^2 \, \frac{4\cos^2\theta + \sin^2\theta}{r^6} \tag{14}$$

$$=\frac{1}{8\pi}m^2 2\pi \int_0^{\pi} d\theta \,\left(4\sin\theta\cos^2\theta + \sin^3\theta\right) \int_R^{\infty} dr \,\frac{1}{r^4} \tag{15}$$

$$= \frac{m^2}{4} \frac{1}{3R^3} \int_0^{\pi} d\theta \left(4\sin\theta\cos^2\theta + \sin\theta[1 - \cos^2\theta] \right)$$
(16)

$$= \frac{m^2}{12R^3} \left[-\frac{4}{3}\cos^3\theta - \cos\theta + \frac{1}{3}\cos^3\theta \right]_0^\pi = \frac{m^2}{12R^3} \left[4 \right] = \frac{m^2}{3R^3}$$
(17)

Note, $W_{\text{mag,in}} = 2W_{\text{mag,out}}$. So the total magnetostatic energy stored in the magnetic fields is,

$$W_{\rm mag} = W_{\rm mag,in} + W_{\rm mag,out} = \frac{2m^2}{3R^3} + \frac{m^2}{3R^3} = \frac{m^2}{R^3}$$
(18)

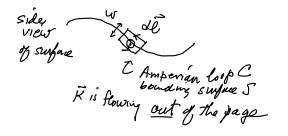
Using $m = e\omega R^2/3c$, we then get,

$$W_{\rm mag} = \frac{e^2 \omega^2 R^4}{9c^2 R^3} = \frac{e^2 \omega^2 R}{9c^2} \tag{19}$$

This is the same result as we obtained from $(1/2c) \oint da \mathbf{K} \cdot \mathbf{A}$.

General Comments

One can ask, is it just some peculiarity of this particular example that caused $A^{in} = A^{out}$ on the surface of the sphere? Might it be different for some other situation?



Let us consider the behavior of **A** at a surface current. Take a small Amperian loop C of length $d\ell$ parallel to the surface, and width w perpendicular to the surface. By Stokes law we can write for the magnetic flux through that loop,

$$\int_{S} da \, \mathbf{\hat{n}} \cdot \mathbf{B} = \int_{S} da \, \mathbf{\hat{n}} \cdot (\mathbf{\nabla} \times \mathbf{A}) = \oint_{C} dd \boldsymbol{\ell} \cdot \mathbf{A}$$
(20)

where S is the surface bounded by the Amperian loop C and $\hat{\mathbf{n}}$ is the normal to that surface.

As we take $w \to 0$, the flux through the loop vanishes as **B** stays finite. The contribution of the sides of width w to the circulation of **A** around the loop also vanishes. We thus get,

$$0 = d\ell \cdot (\mathbf{A}^{\text{above}} - \mathbf{A}^{\text{below}}) \tag{21}$$

Since $d\ell$ can point in any direction within the tangent plane of the surface, we conclude that the tangential component of **A** must always be continuous at a surface current.

Now consider,

$$\int da \,\mathbf{K} \cdot \mathbf{A} \tag{22}$$

where the integral is over a surface containing the surface current **K**. Since **K** necessarily lies in a direction tangent to the surface, it is only the tangential component of **A** that contributes to this integral. Since the tangential components of $\mathbf{A}^{\text{above}}$ and $\mathbf{A}^{\text{below}}$ must be equal, it therefore does not matter whether we use $\mathbf{A}^{\text{above}}$ or $\mathbf{A}^{\text{below}}$ when computing this integral – we must get the same answer in either case.