

Bose-Einstein Condensation in laser cooled gases

Gases of alkali atoms Li, Na, K, Rb, Cs

- all have a single s-electron in outermost shell.
 - important for efficiency of laser cooling
- use isotopes such that total intrinsic spin of all electrons and nucleons add up to an integer T_0
 - \Rightarrow atoms are bosons
- all have a net magnetic moment - used to confine dilute gas of atoms in a "magnetic trap"
- use "laser cooling" to get very low temperatures in low density gases, to try and see BEC

magnetic trap \rightarrow effective harmonic potential for atoms

$$V(r) = \frac{1}{2} m \omega_0^2 r^2 \quad \omega_0 \approx 2\pi \times 100 \text{ Hz}$$

1995 - 10^3 atoms with $T_0 \sim 100 \text{ nK}$

1999 - 10^8 atoms with $T_0 \sim \mu\text{K}$ gas size \sim many microns

How was BEC in these systems observed?

energy levels of ideal (non-interacting) bosons in harmonic trap

$$E(n_x, n_y, n_z) = (n_x + n_y + n_z + 3/2) \hbar \omega_0$$

n_x, n_y, n_z integers

ground state condensate wavefunction

$$\psi_0(r) \sim e^{-r^2/2a^2} \quad \text{with } a = \left(\frac{\hbar}{m\omega_0}\right)^{1/2}$$

$a \sim 1 \mu\text{m}$ for current traps

⇒ condensate has spatial extent $\sim a$

The spatial extent of the n^{th} excited energy level is roughly

$$m\omega_0^2 \langle r^2 \rangle \sim E(n) \approx n\hbar\omega_0$$

$$\Rightarrow \langle r^2 \rangle \sim \frac{n\hbar}{m\omega_0} \quad \text{or} \quad \sqrt{\langle r^2 \rangle} = \left(\frac{n\hbar}{m\omega_0} \right)^{1/2}$$

For $k_B T \gg \hbar\omega_0$, the atoms are excited up to level $n \sim \frac{k_B T}{\hbar\omega_0}$

⇒ spatial extent of the normal component of the gas is

$$R \sim \left(\frac{n\hbar}{m\omega_0} \right)^{1/2} \sim \left(\frac{\hbar k_B T}{\hbar m \omega_0^2} \right)^{1/2} = \left(\frac{k_B T}{m \omega_0^2} \right)^{1/2}$$

$$R \sim a \left(\frac{k_B T}{\hbar\omega_0} \right)^{1/2} \gg a$$

If T_c is the BEC transition temperature, then for $T > T_c$ one sees a more or less uniform cloud of atoms with radius $R \sim a \left(\frac{k_B T}{\hbar\omega_0} \right)^{1/2} \gg a$. But when one cools to $T < T_c$, one now has a finite fraction of the atoms condensed in the ground state, \rightarrow superimposed on the atomic cloud of radius R one sees the growth of a sharp peak in density at the center of cloud - this peak has a radius $a \ll R$.

To find the Bose-Einstein condensation temperature

The number of particles in the system is

$$N = \sum_{n_x, n_y, n_z} \left[\frac{1}{z^{-1} e^{\epsilon(n_x, n_y, n_z)/k_B T} - 1} \right] \leftarrow \text{Bose occupation function}$$

Let $\epsilon_0 = \epsilon(0, 0, 0) = \frac{3}{2} \hbar \omega_0$ the ground state energy

that the Bose occupation function can not be negative $\Rightarrow z^{-1} e^{\epsilon_0/k_B T} \geq 1 \Rightarrow z \leq e^{\epsilon_0/k_B T}$
 $\Rightarrow \mu \leq \epsilon_0$

For the Bose condensed state, z assumes its upper limit, i.e. $\mu = \epsilon_0$ or $z = e^{\epsilon_0/k_B T}$
 - this gives the greatest possible density in the excited states

$$\Rightarrow \text{for } T \leq T_c, N = \sum_{n_x, n_y, n_z} \left[\frac{1}{e^{(n_x + n_y + n_z) \hbar \omega_0 / k_B T} - 1} \right]$$

$$\Rightarrow N = N_0 + \int_0^\infty \int_0^\infty \int_0^\infty dx dy dz \left[\frac{1}{e^{(x+y+z) \hbar \omega_0 / k_B T} - 1} \right]$$

\uparrow
 number in ground state
 $n_x = n_y = n_z = 0$

\uparrow
 number in excited states.
 Make integral approx from sum using $\Delta n_x = \Delta n_y = \Delta n_z = 1$

$$N = N_0 + \left(\frac{k_B T}{\hbar \omega_0} \right)^3 \int_0^\infty \int_0^\infty \int_0^\infty dx dy dz \left[\frac{1}{e^{(x+y+z)} - 1} \right]$$

$$= N_0 + \left(\frac{k_B T}{\hbar \omega_0} \right)^3 \zeta(3) \quad \zeta(3) = 1 + \frac{1}{2^3} + \frac{1}{3^3} + \frac{1}{4^3} + \dots$$

At T_c , $N_0 = 0 \Rightarrow k_B T_c = \hbar \omega_0 \left(\frac{N}{\zeta(3)} \right)^{1/3}$
 for $T < T_c$, $N_0(T) = N \left(1 - (T/T_c)^3 \right)$

power of T/T_c term is different from ideal free gas due to presence of magnetic fields

Classical spin models

$$U = -J \sum_{\langle ij \rangle} \vec{S}_i \cdot \vec{S}_j$$

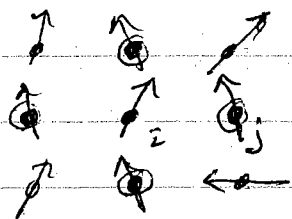
simple model of interacting magnetic

classical spins \vec{S}_i of unit magnitude $|\vec{S}_i| = 1$ on sites i of a periodic d -dimensional lattice.

\vec{S}_i interacts only with its neighbors \vec{S}_j

$\langle ij \rangle$ indicates nearest neighbor bonds of the lattice.

If coupling $J > 0$, then ferromagnetic interaction i.e. spins are in lower energy state when they are aligned.



\vec{S}_i interacts with spins on sites labeled by \odot .

Behavior of model depends significantly on dimensionality of lattice d , and number of components of the spin \vec{S} n .

Examples: $\vec{S} = (S_x, S_y, S_z)$ points in 3-dimensional space
 $n=3$ called the Heisenberg model

$\vec{S} = (S_x, S_y)$ restricted to lie in a plane
 $n=2$ called the XY model

$S = S_z = \pm 1$ restricted to lie in one direction
 $n=1$ called the Ising model

less obvious possibilities $\left\{ \begin{array}{l} n=0 \\ n=\infty \end{array} \right.$ called the polymer model
called the spherical model

We will focus on the Ising model (1925)

$$S = \pm 1$$

Ensembles

① fixed magnetization $M = \sum_i s_i$ M is total magnetization

partition function $\tilde{Z}(T, M) = \sum_{\{s_i\}} e^{-\beta \mathcal{H}[s_i]}$
s.t. $\sum_i s_i = M$

sum over all spin configurations

obeying the constraint $\sum_i s_i = M = N^+ - N^-$

\uparrow # up spins \uparrow # down spins

(similar to canonical ensemble with $\sum_i n_i = N$ total # particles)

Helmholtz free energy $F(T, M) = -k_B T \ln \tilde{Z}(T, M)$

② fixed magnetic field

to remove constraint of fixed M we can Legendre transform to a conjugate variable h , ~~the magnetic field~~. We will see that h is just the magnetic field.

Gibbs free energy $G(T, h) = F(T, M) - h M$

where ~~the~~ $\left(\frac{\partial F}{\partial M}\right)_T = h \quad \Rightarrow \quad \left(\frac{\partial G}{\partial h}\right)_T = -M$

$$dF = -SdT + h dM \quad \text{and} \quad dG = -SdT - Mdh$$

\uparrow entropy \uparrow entropy

To get partition function for G , take Laplace transform of \tilde{Z}

$$Z(T, h) = \sum_M e^{\beta h M} \tilde{Z}(T, M)$$

$$= \sum_M e^{\beta h M} \sum_{\{s_i\}} e^{-\beta H[s_i]} \quad \text{use } M = \sum_i s_i$$

$s.t. \sum_i s_i = M$

$$Z(T, h) = \sum_{\{s_i\}} e^{-\beta [H[s_i] - h \sum_i s_i]}$$

↙ looks like interaction of magnetic field h with total magnetization $M = \sum_i s_i$

↖ unconstrained sum over all spin configs $\{s_i\}$
 (similar to grand canonical ensemble with $\sum_i n_i = N$ unconstrained)

$$G(T, h) = -k_B T \ln Z(T, h)$$

Check:

$$\frac{\partial G}{\partial h} = -k_B T \frac{\partial Z}{Z \partial h} = -k_B T \sum_{\{s_i\}} \frac{\partial}{\partial h} \left(e^{-\beta [H - h \sum_i s_i]} \right)$$

$$= -k_B T \sum_{\{s_i\}} e^{-\beta [H - h \sum_i s_i]} \left(\beta \sum_i s_i \right)$$

$$= - \frac{\sum_{\{s_i\}} e^{-\beta [H - h \sum_i s_i]} \left(\sum_i s_i \right)}{\sum_{\{s_i\}} e^{-\beta [H - h \sum_i s_i]}}$$

$$= - \left\langle \sum_i s_i \right\rangle = -M$$

so $\frac{\partial G}{\partial h} = -M$ as required

we can work in fixed magnetization or fixed magnetic field ensemble according to our convenience. Usually it is easiest to work with fixed magnetic field. In this case we usually write

$$H = -J \sum_{\langle ij \rangle} S_i S_j - h \sum_i S_i$$

including the magnetic field part in the definition of H .

$$Z = \sum_{\{S_i\}} e^{-\beta H} \quad \uparrow \text{ includes } h \text{ term}$$

define magnetization density

$$m = \frac{M}{N} = \frac{1}{N} \left\langle \sum_i S_i \right\rangle \quad N = \text{total number spins}$$

Helmholtz free energy density: In limit $N \rightarrow \infty$, $F(T, M) = N f(T, m)$

$\frac{F}{N} \equiv f(T, m)$ depends on magnetization density

$$df = -s dT + h dm$$

$s = \frac{S}{N}$ entropy per spin

Gibbs free energy density: In limit $N \rightarrow \infty$, $G(T, h) = N g(T, h)$

$$\frac{G}{N} \equiv g(T, h)$$

$$dg = -s dT - m dh$$

$$\left(\frac{\partial f}{\partial m} \right)_T = h, \quad \left(\frac{\partial g}{\partial h} \right)_T = -m$$