

Energy in the degenerate limit $T=0$

$$\frac{E}{V} = \int_0^{\epsilon_F} d\epsilon g(\epsilon) \epsilon$$

↑
density of states

$$g(\epsilon) = C \sqrt{\epsilon}$$

with $C = \left(\frac{2\pi m}{h^2}\right)^{3/2} \frac{2g_s}{\sqrt{\pi}}$

$$m = \frac{N}{V} = \int_0^{\epsilon_F} d\epsilon g(\epsilon)$$

$$\Rightarrow \frac{E}{V} = C \int_0^{\epsilon_F} d\epsilon \epsilon^{3/2} = \frac{2}{5} C \epsilon_F^{5/2}$$

$$m = \frac{N}{V} = C \int_0^{\epsilon_F} d\epsilon \epsilon^{1/2} = \frac{2}{3} C \epsilon_F^{3/2}$$

$\Rightarrow \frac{E}{V} = \frac{3}{5} \frac{N}{V} \epsilon_F$

$$\frac{E}{V} = \frac{3}{5} m \epsilon_F \quad \text{or} \quad \boxed{\frac{E}{N} = \frac{3}{5} \epsilon_F}$$

↑ energy per volume

↑ energy per particle

Above gives $T=0$ results. To get behavior at low $T > 0$, or to get quantities such as $C_V = \left(\frac{\partial E}{\partial T}\right)_V$, we need to get the next order terms in a low temperature expansion.

In general we need to do integrals of the form

$$\int d\epsilon \frac{\tilde{\phi}(\epsilon)}{z^{\gamma} e^{\beta\epsilon} + 1} = \int d\epsilon \tilde{\phi}(\epsilon) m(\epsilon) \quad , \quad \tilde{\phi}(\epsilon) \text{ some function}$$

ex: to compute m , $\tilde{\phi}(\epsilon) = g(\epsilon)$; to compute $\frac{E}{V}$, $\tilde{\phi}(\epsilon) = g(\epsilon) \epsilon$

transform variables to $X = \beta t$.

Then we want to do integrals of the form

$$\Phi \equiv \int_0^\infty dx \frac{\phi(x)}{z^{-1} e^x + 1} \quad \phi(x) \text{ is any function of } x.$$

For example, to get the "standard" function $f_n(z)$, we use $\phi(x) = \frac{1}{n!} x^{n-1}$

Define $\xi = \beta \mu = \ln z$

$$\Phi = \int_0^\infty dx \frac{\phi(x)}{e^{x-\xi} + 1}$$

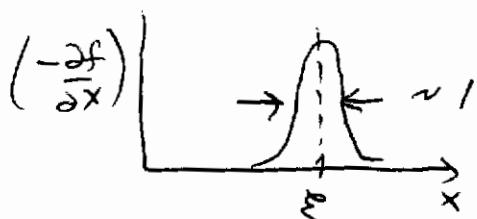
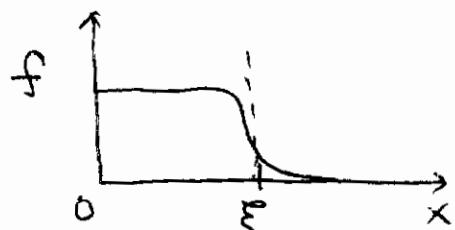
Define $\psi(x) = \int_x^\infty \phi(x') dx'$, $f(x) = \frac{1}{[e^{x-\xi} + 1]}$ fermi function

$$\Phi = \int_0^\infty dx \left(\frac{\partial \psi}{\partial x} \right) f(x) \quad \text{integrate by parts}$$

$$= \psi(x) f(x) \Big|_0^\infty + \int_0^\infty dx \psi(x) \left(-\frac{\partial f}{\partial x} \right)$$

$$= \int_0^\infty dx \psi(x) \left(-\frac{\partial f}{\partial x} \right) \quad \text{since } \psi(0) = 0 \text{ and } f(\infty) = 0 \\ \text{1st term vanishes}$$

Now we use the fact that at low T, $\left(-\frac{\partial f}{\partial x} \right)$ is strongly peaked about $x = \xi$



$\xi \gg 1$
 $\xi \sim \frac{E_F}{kT}$ large

expand $\psi(x)$ about $x = \xi$

$$\psi(x) = \sum_{n=0}^{\infty} \frac{d^n \psi}{dx^n} \Big|_{x=\xi} \frac{(x-\xi)^n}{n!}$$

$$\Rightarrow \Phi = \sum_{n=0}^{\infty} \frac{d^n \psi}{dx^n} \Big|_{x=\xi} \int_0^{\infty} dx \frac{(x-\xi)^n}{n!} \left(-\frac{\partial f}{\partial x} \right)$$

since $\left(-\frac{\partial f}{\partial x} \right)$ is zero except for a region of order 1 about $x = \xi \gg 1$, we can replace the lower limit of the integral by $-\infty$ without any noticeable change

Then we can make a change of variable $y = x - \xi$ and the integrals become

$$\int_{-\infty}^{\infty} dy \frac{y^n}{n!} \left(-\frac{\partial f}{\partial y} \right) \quad \text{where } f(y) = \frac{1}{e^y + 1}$$

$$\text{Now } -\frac{\partial f}{\partial y} = \frac{e^y}{(e^y + 1)^2} = \frac{e^y}{e^{2y} + 2e^y + 1} = \frac{1}{e^y + 2 + e^{-y}}$$

is symmetric about $y = 0$.

\Rightarrow all the integrals for n odd vanish!

To sum over only n even terms, let $n \rightarrow 2n$

$$\Phi = \sum_{n=0}^{\infty} \frac{d^{2n}\Phi}{dx^{2n}} \Big|_{x=\xi} \int_{-\infty}^{\infty} dy \frac{y^{2n}}{(2n)!} \left(-\frac{\partial f}{\partial y} \right)$$

$$\text{let } a_n = \int_{-\infty}^{\infty} dy \frac{y^{2n}}{(2n)!} \left(-\frac{\partial f}{\partial y} \right) \rightarrow a_0 = \int_{-\infty}^{\infty} dy \left(-\frac{\partial f}{\partial y} \right) = 1$$

The a_n are just numbers that we computed.
They contain no system parameters whatsoever

For $n \geq 1$ one can show

$$a_n = 2 \left(1 - \frac{1}{2^{2n}} + \frac{1}{3^{2n}} - \frac{1}{4^{2n}} + \frac{1}{5^{2n}} - \dots \right)$$

$$= \left(2 - \frac{1}{2^{2(n-1)}} \right) \zeta(2n)$$

where $\zeta(n) = 1 + \frac{1}{2^n} + \frac{1}{3^n} + \frac{1}{4^n} + \dots$ is the Riemann zeta function

$$\text{In particular } a_1 = \frac{\pi^2}{6}, \quad a_2 = \frac{7\pi^4}{360}$$

$$\Phi = \sum_{n=0}^{\infty} a_n \frac{d^{2n}\Phi}{dx^{2n}} \Big|_{x=\xi} = \Phi(\xi) + \sum_{n=1}^{\infty} a_n \frac{d^{2n}\Phi}{dx^{2n}} \Big|_{x=\xi}$$

use $\frac{d\Phi}{dx} = \phi$ to finally get

$$\Phi(x) = \int_0^x dx' \phi(x')$$

$$\Phi = \int_0^{\xi} dx \phi(x) + \sum_{n=1}^{\infty} a_n \frac{d^{2n-1}\phi}{dx^{2n-1}} \Big|_{x=\xi}$$

$$= \int_0^{\xi} dx \phi(x) + \frac{\pi^2}{6} \frac{d\phi}{dx} \Big|_{x=\xi} + \frac{7\pi^4}{360} \frac{d^3\phi}{dx^3} \Big|_{x=\xi} + \dots$$

This gives a power series in temperature.

To see this, transform back to the energy variable

$$x = \beta \epsilon, \quad \epsilon = k_B T x$$

$$\Phi = \int_0^\infty d\epsilon \frac{\phi(\epsilon)}{Z^{-1} e^{\beta \epsilon} + 1} = k_B T \left\{ \int_0^\infty dx \frac{\phi(k_B T x)}{Z^{-1} e^{x} + 1} \right\}$$

$$\text{Using } k_B T \int_0^\infty dx \phi(k_B T x) = \int_0^\infty d\epsilon \phi(\epsilon)$$

$$\text{and } \frac{d\phi}{dx} = \frac{d\phi}{d\epsilon} \frac{d\epsilon}{dx} = \frac{d\phi}{d\epsilon} k_B T$$

we get

$$\Phi = \int_0^\infty d\epsilon \phi(\epsilon) m(\epsilon)$$

$$\boxed{\bar{\Phi} = \int_0^\mu d\epsilon \phi(\epsilon) + \frac{\pi^2 (k_B T)^2}{6} \frac{d\phi}{d\epsilon} \Big|_{\epsilon=\mu} + \frac{7\pi^4 (k_B T)^4}{360} \frac{d^3\phi}{d\epsilon^3} \Big|_{\epsilon=\mu} + \dots}$$

Example

$$\textcircled{1} \text{ density } m = \frac{N}{V} = \int_0^\infty d\epsilon g(\epsilon) m(\epsilon) \Rightarrow \phi(\epsilon) = g(\epsilon)$$

$$m = \int_0^\mu d\epsilon g(\epsilon) + \frac{\pi^2 (k_B T)^2}{6} \frac{dg}{d\epsilon} \Big|_{\epsilon=\mu} + \dots$$

Now as $T \rightarrow 0$, $\mu \rightarrow E_F$ the fermi energy

$$n = \int_0^{\epsilon_F} d\epsilon g(\epsilon) + \int_{\epsilon_F}^{\mu} d\epsilon g(\epsilon) + \frac{\pi^2}{6} (k_B T)^2 \left. \frac{dg}{d\epsilon} \right|_{\epsilon=\mu}$$

But ϵ_F was determined by $n = \int_0^{\epsilon_F} d\epsilon g(\epsilon)$

$$\Rightarrow \int_{\epsilon_F}^{\mu} d\epsilon g(\epsilon) = -\frac{\pi^2}{6} (k_B T)^2 \left. \frac{dg}{d\epsilon} \right|_{\epsilon=\mu}$$

since left hand side is $O(kT)^2$ is small, we can approx
~~the right hand side~~ as it as

$$\int_{\epsilon_F}^{\mu} d\epsilon g(\epsilon) \approx (\mu - \epsilon_F) g(\epsilon_F)$$

$$\Rightarrow (\mu - \epsilon_F) \approx -\frac{\pi^2}{6} (k_B T)^2 \left. \frac{dg}{d\epsilon} \right|_{\epsilon=\mu}$$

so $\mu - \epsilon_F \sim O(k_B T)^2$ is small, so to lowest order
 can evaluate $\frac{dg}{d\epsilon}$ on right hand side at $\epsilon = \epsilon_F$

instead of $\epsilon = \mu$

$$\boxed{\mu(T) \approx \epsilon_F - \frac{\pi^2}{6} (k_B T)^2 \frac{g'(\epsilon_F)}{g(\epsilon_F)}}$$

$$g' = \frac{dg}{d\epsilon}$$

Shows that chemical potential μ decreases from ϵ_F
 by $O(kT)^2$ at low T

For free electrons where $g(\epsilon) = C \sqrt{\epsilon}$

$$g'(\epsilon) = \frac{1}{2} C \frac{1}{\sqrt{\epsilon}}$$

$$\mu(T) \approx \epsilon_F - \frac{\pi^2}{6} (k_B T)^2 \frac{1}{2\epsilon_F} = \epsilon_F - \frac{\pi^2}{12} \frac{(k_B T)^2}{\epsilon_F}$$

$$\boxed{\mu(T) \approx \epsilon_F \left(1 - \frac{1}{3} \left(\frac{\pi k_B T}{2\epsilon_F}\right)^2\right) = \epsilon_F \left(1 - \frac{1}{3} \left(\frac{\pi T}{2T_F}\right)^2\right)}$$

Correction is small for metals at room temp as $T_F \sim 10,000^\circ K$

② energy $\frac{E}{V} = \int_0^\infty d\epsilon g(\epsilon) \epsilon n(\epsilon) \Rightarrow \phi(\epsilon) = g(\epsilon) \epsilon$

$$U = \frac{E}{V} = \int_0^\mu d\epsilon g(\epsilon) \epsilon + \frac{\pi^2}{6} (k_B T)^2 [g(\mu) + \mu g'(\mu)]$$

$$= \underbrace{\int_0^{\epsilon_F} d\epsilon g(\epsilon) \epsilon}_{= U(0)} + \underbrace{\int_{\epsilon_F}^\mu d\epsilon g(\epsilon) \epsilon}_{\text{as before}} + \frac{\pi^2}{6} (k_B T)^2 [g(\mu) + \mu g'(\mu)]$$

ground state $\simeq (\mu - \epsilon_F) g(\epsilon_F) \epsilon_F$ replace $\mu \approx \epsilon_F$
 energy density as before as before

$$U(T) = U(0) + (\mu - \epsilon_F) g(\epsilon_F) \epsilon_F + \frac{\pi^2 (k_B T)^2}{6} [g(\epsilon_F) + \epsilon_F g'(\epsilon_F)]$$

$$= U(0) + \left[-\frac{\pi^2}{6} (k_B T)^2 \frac{g'(\epsilon_F)}{g(\epsilon_F)} \right] g(\epsilon_F) \epsilon_F + \frac{\pi^2 (k_B T)^2}{6} [g(\epsilon_F) + \epsilon_F g'(\epsilon_F)]$$

$$\boxed{U(T) = U(0) + \frac{\pi^2 (k_B T)^2}{6} g(\epsilon_F)}$$

specific heat per volume

$$C_V = \frac{C_V}{V} = \frac{1}{V} \left(\frac{\partial E}{\partial T} \right)_V = \left(\frac{\partial U}{\partial T} \right)_V$$

$$C_V = \frac{\pi^2 k_B^2}{3} T g(\epsilon_F)$$

for free electrons we can write $g(\epsilon) = C\sqrt{\epsilon}$

$$m = \int_0^{\epsilon_F} \epsilon g(\epsilon) d\epsilon = \frac{2}{3} C \epsilon^{3/2} \Rightarrow C = \frac{3}{2} \frac{m}{\epsilon_F^{3/2}}$$

$$\Rightarrow g(\epsilon_F) = \frac{3}{2} \frac{m}{\epsilon_F^{3/2}} \cdot \epsilon_F^{1/2} = \frac{3}{2} \frac{m}{\epsilon_F}$$

density of states
at fermi energy

$$C_V = \frac{\pi^2}{2} \left(\frac{k_B T}{\epsilon_F} \right) m k_B$$

or total specific heat $C_V = V C_v$ $mV = N$

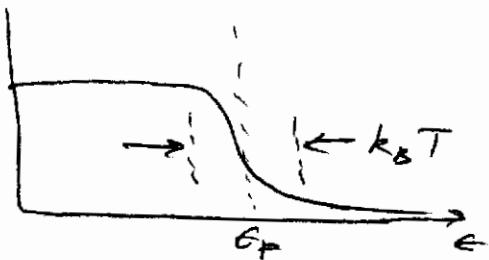
$$C_V = \frac{\pi^2}{2} \left(\frac{k_B T}{\epsilon_F} \right) N k_B$$

\Rightarrow specific heat due to fermi gas of electrons in a conductor is $C_V \sim T$ at low temperatures

We already saw that specific heat due to ionic vibrations (phonons) in a solid went like $C_V \sim T^3$ at low temperatures (Debye model)

\Rightarrow electronic contribution to C_V dominates at sufficiently low T .

Simple estimate of C_V



When we increase temperature to $k_B T$, the electrons near the Fermi energy ϵ_F will increase their energy by an amount $\sim k_B T$. The number of such electrons ~~is roughly~~ per unit volume is roughly

$$g(\epsilon_F)(k_B T)$$

\uparrow \uparrow
 density of states energy interval about ϵ_F of
 states which ~~may~~ get excited
 at ϵ_F

\Rightarrow increase in energy per unit volume is

$$\Delta U \sim (g(\epsilon_F) k_B T) (k_B T) \sim g(\epsilon_F) (k_B T)^2$$

\uparrow \uparrow
 # electrons excitation
 excited energy per
 for free
 electrons

$$\Rightarrow C_V = \frac{1}{T} \frac{\Delta U}{\Delta T} \sim g(\epsilon_F) k_B^2 T = \frac{3}{2} \frac{m}{\epsilon_F} k_B^2 T = \frac{3}{2} M k_B \left(\frac{T}{T_F} \right)$$

The previous calculation gives the precise numerical coefficient

electronic specific heat per volume

$$C_V^{\text{elec}} = \frac{\pi^2}{2} \left(\frac{k_B T}{\epsilon_F} \right) \frac{N k_B}{V} \left(1 + o \left(\frac{k_B T}{\epsilon_F} \right)^2 \right)$$

compare to classical result $C_V^{\text{classical}} = \frac{N k_B}{V}$

The correct result for degenerate fermi gas is a factor

$$\frac{\pi^2}{2} \left(\frac{k_B T}{\epsilon_F} \right) = \frac{\pi^2}{2} \left(\frac{T}{T_F} \right) \text{ smaller than classical result by factor } \sim \frac{10^2}{10^4} = 10^{-2} \text{ at room temperature}$$

also, classical C_V is indep of T , whereas fermi gas result is $\propto T$.

At low T , the ionic contribution to C_V is

$$C_V^{\text{ion}} = \frac{12\pi^4}{5} \left(\frac{T}{\Theta_D} \right) \frac{3}{V} N k_B$$

$$\frac{C_V^{\text{elec}}}{C_V^{\text{ion}}} = \frac{\pi^2}{2} \left(\frac{k_B T}{\epsilon_F} \right) \frac{5}{12\pi^4} \left(\frac{\Theta_D}{T} \right)^3 \approx \frac{5}{24\pi^2} \left(\frac{\Theta_D}{T_F} \right) \left(\frac{\Theta_D}{T} \right)^2$$

$$\approx 1 \quad \text{when} \quad T^* = \sqrt{\frac{5}{24\pi^2} \left(\frac{\Theta_D}{T_F} \right)} \Theta_D \approx 0.15 \left(\frac{\Theta_D}{T_F} \right)^{1/2} \Theta_D$$

for metals, $T_F \sim 10^4 \text{ K}$, $\Theta_D \sim 10^2 \text{ K}$

$$T^* = 0.15 \sqrt{10^{-2}} \Theta_D \approx 0.015 \Theta_D$$

so ionic contrib to C_V dominates over electronic contrib until $T \lesssim 0.01 \Theta_D$ ie at $0(1) \text{ K}$. The electronic contrib dominates at lower temperatures.