Average energy $\langle E \rangle$ vs. the most probable energy $\bar{E}$ in the canonical ensemble.

In our earlier discussion of fluctuations in the canonical ensemble we expanded

$$E - TS(E) \cong A_{micro}(T) + \frac{\delta E^2}{2TC_V}$$

now we continue the expansion to $O(\delta E^3)$

$$E - TS(E) \cong A_{micro}(T) + \frac{\delta E^2}{2TC_V} - \frac{1}{3!} T \frac{\partial^3 S}{\partial E^3}\bigg|_{E=\bar{E}} \delta E^3$$

Note $\frac{\partial^3 S}{\partial E^3} \sim \frac{1}{N^2}$ since $S \sim N$ and $E \sim N$ are both extensive

so we can write

$$E - TS(E) \cong A_{micro}(T) + \frac{\delta E^2}{2TC_V} - \frac{\gamma}{N^2} \delta E^3$$

where γ is some constant that does not increase with N (it can depend on T)

Now compute $\langle \delta E \rangle = \langle E \rangle - \bar{E}$

↑ difference between average value and most probable value

$$\langle \delta E \rangle = \frac{\int \frac{d\delta E}{\Delta} e^{-(E - TS(E))/k_B T} \delta E}{\int \frac{d\delta E}{\Delta} e^{-(E - TS(E))/k_B T)}}$$

$$\simeq \frac{\int d\delta E \; e^{-\frac{\delta E^2}{2k_B T^2 C_V} + \frac{\gamma \delta E^3}{N^2 k_B T}} \delta E}{\int d\delta E \; e^{-\frac{\delta E^2}{2k_B T^2 C_V} + \frac{\gamma}{N^2}\frac{\delta E^3}{k_B T}}}$$

$$\simeq \frac{\int d\delta E \; e^{-\frac{\delta E^2}{2k_B T^2 C_V}} \left(1 + \frac{\gamma \delta E^3}{N^2 k_B T}\right) \delta E}{\int d\delta E \; e^{-\frac{\delta E^2}{2k_B T^2 C_V}} \left(1 + \frac{\gamma \delta E^3}{N^2 k_B T}\right)}$$

where we expanded $e^{\frac{\gamma \delta E^3}{N^2 k_B T}} \approx 1 + \frac{\gamma \delta E^3}{N^2 k_B T}$ for $N \to \infty$

For a Gaussian distribution, only the even moments are non-vanishing

$$\langle \delta E \rangle \simeq \frac{\int d\delta E \; e^{-\frac{\delta E^2}{2k_B T^2 C_V}} \cdot \left(\frac{\gamma}{N^2 k_B T}\right) \delta E^4}{\int d\delta E \; e^{-\frac{\delta E^2}{2k_B T^2 C_V}}}$$

$$= \left(\frac{\gamma}{N^2 k_B T}\right) \cdot \left(k_B T^2 C_V\right)^2 3$$

where we used $$\frac{\int_{-\infty}^{\infty} dx\, e^{-\frac{1}{2}\frac{x^2}{\sigma^2}} x^4}{\int_{-\infty}^{\infty} dx\, e^{-\frac{1}{2}\frac{x^2}{\sigma^2}}} = 3\sigma^4$$

But the main point is $C_V \sim N$

so $$\langle \delta E \rangle \sim \frac{1}{N^2} \cdot N^2 \sim O(1)$$

The relative difference between average and most probable energy therefore scales as

$$\frac{\langle E \rangle - \bar{E}}{\langle E \rangle} = \frac{\langle \delta E \rangle}{\langle E \rangle} \sim \frac{1}{N} \to 0 \quad \text{as } N \to \infty$$

We introduced the canonical distribution as a means of describing a physical system in contact with a heat bath.

The canonical distrib gives the same result as the microcannonical because in the $N \to \infty$ (thermodynamic) limit, the canonical probability distribution

$$p(E) = \frac{\Omega(E)\, e^{-E/k_B T}}{\Delta\, Q_N(V,T)}$$

approaches a delta-function* at the most probable energy = average energy, as set by the temperature T.

We could alternatively introduced the canonical ensemble just as a mathematical trick for computing $\Omega(E)$,

* since $E \sim N$ increases as $N \to \infty$
and $\langle E^2\rangle - \langle E\rangle^2$ increases as $\sqrt{N}$
it is not really $P(E)$ that approaches a well defined function as $N \to \infty$. Rather it is the distribution
$p(u \equiv E/N)$, the probability density to have an energy per particle u, that approaches a delta function as $N \to \infty$

Stirling's Formula

In lecture we used the saddle point approx to discuss the relation between the Helmholtz free energy in the canonical vs. the micro canonical ensemble. The saddle pt approx is also how one derives Stirling's approx for $n!$

Consider the integral

$$I = \int_0^\infty dx \; x^n e^{-x}$$

integrate by parts

$$I = -x^n e^{-x}\Big]_0^\infty + \int_0^\infty n x^{n-1} e^{-x} dx$$

boundary term vanishes at its limits so

$$I = \int_0^\infty dx \; n x^{n-1} e^{-x}$$

integrate by parts again

$$I = \int_0^\infty dx \; n(n-1) x^{n-2} e^{-x}$$

and so on to get

$$I = \int_0^\infty dx \; n(n-1)(n-2)\cdots(1) e^{-x} = n!$$

Now evaluate I in saddle pt approx.

Define $u(x) = -x + n \ln x$

$$I = \int_0^\infty dx\, e^{u(x)}$$

expand $u(x)$ about its maximum

$$u'(x) = -1 + \frac{n}{x} \Rightarrow \begin{cases} u(\bar{x}) = -n + n\ln n \\ \bar{x} = n \quad \text{is the maximum} \end{cases}$$

$$u''(x) = -\frac{n}{x^2} \Rightarrow u''(\bar{x}) = -\frac{1}{n}$$

$$u'''(x) = \frac{2n}{x^3} \qquad u'''(\bar{x}) = 2/n^2$$

$$u''''(x) = -\frac{6n}{x^4} \qquad u''''(\bar{x}) = -6/n^3$$

For $\delta x = x - \bar{x}$,

$$u(x) \simeq -n + n\ln n - \frac{\delta x^2}{2n} + \frac{1}{6}\frac{2\,\delta x^3}{n^2} - \frac{1}{24}\frac{6\,\delta x^4}{n^3} + \cdots$$

$$= -n + n\ln n - \frac{\delta x^2}{2n} + \frac{\delta x^3}{3n^2} - \frac{\delta x^4}{4n^3} + \cdots$$

$$I = \int_0^\infty dx\, e^{-n+n\ln n}\, e^{-\delta x^2/2n}\, \underbrace{e^{\frac{\delta x^3}{3n^2} - \frac{\delta x^4}{4n^3}}}_{\text{expand for small } \delta x}$$

$$\simeq \int_{-\infty}^{\infty} d\delta x\, e^{-n+n\ln n}\, e^{-\delta x^2/2n} \left[1 + \frac{\delta x^3}{3n^2} - \frac{\delta x^4}{4n^3} + O(\delta x^6)\right]$$

$$= e^{-n+n\ln n} \int_{-\infty}^{\infty} d\delta x\, e^{-\delta x^2/2n} \left[1 + \frac{\delta x^3}{3n^2} - \frac{\delta x^4}{4n^3} + \cdots\right]$$

$$= e^{-n+n\ln n} \sqrt{2\pi n} \left[1 + \frac{\langle \delta x^3\rangle}{3n^2} - \frac{\langle \delta x^4\rangle}{4n^3} + \cdots\right]$$

(3)

Now $\langle \delta x^3 \rangle = 0$, $\langle \delta x^4 \rangle \sim n^2$, so

$$I = n! = e^{-n + n\ln n} \sqrt{2\pi n} \left[1 + O\left(\frac{1}{n}\right) \right]$$

$$\ln n! = n\ln n - n + \frac{1}{2}\ln n + \frac{1}{2}\ln 2\pi + \ln\left(1 + O\left(\frac{1}{n}\right)\right)$$

$$= \underbrace{n\ln n - n}_{\text{these are the leading terms}} + \underbrace{\frac{1}{2}\ln n + \frac{1}{2}\ln 2\pi}_{\text{these are next order corrections}} + O\left(\frac{1}{n}\right)$$

Factorization of canonical partition function – the ideal gas

Consider a system of N noninteracting particles

$$\Rightarrow \mathcal{H}[\vec{q}_i, \vec{p}_i] = \sum_{i=1}^{N} \mathcal{H}^{(1)}(\vec{q}_i, \vec{p}_i)$$

where $\mathcal{H}^{(1)}$ is the single particle Hamiltonian that depends only on the three coordinates $\vec{q}_i$ and three momenta $\vec{p}_i$ of particle i.

$$Q_N = \frac{1}{N!\, h^{3N}} \left(\prod_{i=1}^{N} \int d\vec{q}_i \, d\vec{p}_i \right) e^{-\beta \mathcal{H}}$$

$$= \frac{1}{N!} \left(\prod_{i=1}^{N} \int \frac{d\vec{q}_i \, d\vec{p}_i}{h^3} \right) e^{-\beta \sum_j \mathcal{H}^{(1)}(\vec{q}_j, \vec{p}_j)}$$

factor the exponential

$$= \frac{1}{N!} \prod_{i=1}^{N} \left(\int \frac{d\vec{q}_i \, d\vec{p}_i}{h^3} e^{-\beta \mathcal{H}^{(1)}(\vec{q}_i, \vec{p}_i)} \right)$$

factor for particle i is identical to factor for particle j

$$\Rightarrow \boxed{Q_N = \frac{1}{N!} (Q_1)^N}$$

for noninteracting particles

where Q_1 is the one particle partition function

$$Q_1 = \int \frac{d\vec{q}\,d\vec{p}}{h^3}\, e^{-\beta H^{(1)}(\vec{q},\vec{p})}$$

Apply to the ideal gas.

$$\mathcal{H}^{(1)}(\vec{q},\vec{p}) = \frac{p^2}{2m}$$

$$Q_1 = \int \frac{d\vec{q}}{h^3} \int d\vec{p}\, e^{-\beta \frac{p^2}{2m}}$$

$$\int d\vec{q} = V \quad \text{volume of system}$$

$$\int d\vec{p}\, e^{-\beta p^2/2m} = \left(\frac{2\pi m}{\beta}\right)^{3/2} \quad \text{3D Gaussian integral}$$

$$Q_1 = \frac{V}{h^3}\left(2\pi m k_B T\right)^{3/2}$$

$$\Rightarrow Q_N = \frac{1}{N!}\left(\frac{V}{h^3}\right)^N \left(2\pi m k_B T\right)^{3N/2}$$

$\ln N! = N\ln N - N$

$$A(T,V,N) = -k_B T \ln Q_N \quad \text{using Stirling's formula}$$

$$= -k_B T\left\{ N \ln\left[\frac{V}{h^3}\left(2\pi m k_B T\right)^{3/2}\right] \underbrace{- N\ln N + N}\right\}$$

$$A(T,V,N) = -k_B T N - k_B T N \ln\left[\frac{V}{h^3 N}\left(2\pi m k_B T\right)^{3/2}\right]$$

Compute average energy

$$\langle E\rangle = -\frac{\partial}{\partial\beta}(\ln Q_N) = -\frac{\partial}{\partial\beta}(-\beta A)$$

$$= -\frac{\partial}{\partial\beta}\left(N + N\ln\left[\frac{V}{h^{3N}}\left(\frac{2\pi m}{\beta}\right)^{3/2}\right]\right)$$

$$= -N\frac{\partial}{\partial\beta}\left(\ln\beta^{-3/2}\right) = \frac{3}{2}N\frac{\partial}{\partial\beta}\ln\beta = \frac{3}{2}N\frac{1}{\beta}$$

$\langle E\rangle = \frac{3}{2}Nk_BT$ as expected

entropy

$$S = -\left(\frac{\partial A}{\partial T}\right)_{V,N} = k_BN + k_BN\ln\left[\frac{V}{h^{3N}}(2\pi m k_B T)^{3/2}\right]$$
$$+ k_BTN\,\frac{3}{2}\left(\frac{1}{T}\right)$$

from derivative of log

$$S = \frac{5}{2}Nk_B + Nk_B\ln\left[\frac{V}{h^{3N}}(2\pi m k_BT)^{3/2}\right]$$

substitute in $k_BT = \frac{2}{3}\frac{E}{N}$ to get

$$\Rightarrow\ S(E,V,N) = \frac{5}{2}Nk_B + Nk_B\ln\left[\frac{V}{h^{3N}}\left(\frac{4\pi mE}{3N}\right)^{3/2}\right]$$

we have recovered the Sackur-Tetrode equation which we earlier derived from the microcanonical ensemble! Canonical and microcanonical approaches are equivalent.

Because in computing Q_N we sum over all states with any energy, as compared to computing Ω where we restrict the sum to states in a particular energy shell E, it is usually easier to compute Q_N, rather than Ω.

Note:

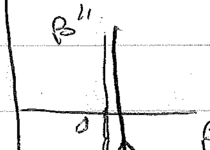
$$Q_N(\beta) = \int \frac{dE}{\Delta}\, \Omega(E)\, e^{-\beta E}$$

$Q_N(\beta)$ is Laplace transform of $\frac{\Omega}{\Delta}(E)$

$\Rightarrow \frac{\Omega}{\Delta}$ is inverse Laplace transform of Q_N

$$\frac{\Omega(E)}{\Delta} = \frac{1}{2\pi i} \int_{\beta'-i\infty}^{\beta'+i\infty} e^{\beta E}\, Q_N(\beta)\, d\beta \qquad (\beta' \geq 0)$$

$$= \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{i(\beta'+i\beta'')}\, Q_N(\beta'+i\beta'')\, d\beta''$$



where $\beta' = \mathrm{Re}(\beta) = 0^+$

contour of integration lies to right of imaginary axis

entropy $S = k_B \ln \Omega$

Helmholtz $-\frac{A}{T} = k_B \ln Q_N$

$$-\frac{A}{T} = S - \frac{E}{T}$$

Helmholtz free energy is Legendre transf of S with respect to E

Thermodynamic potentials which are Legendre transforms of each other, have ensemble partition functions that are Laplace transforms of each other.