

Solutions Problem Set 2

1) Prove $\left(\frac{\partial C_p}{\partial p}\right)_T = -TV \left[\alpha^2 + \left(\frac{\partial \alpha}{\partial T}\right)_p\right]$

$$C_p = T \left(\frac{\partial S}{\partial T}\right)_p = \text{definition of } C_p$$

$$S(T, p) = -\left(\frac{\partial G(T, p)}{\partial T}\right)_p$$

$$C_p = -T \left(\frac{\partial^2 G}{\partial T^2}\right)_p$$

$$\left(\frac{\partial C_p}{\partial p}\right)_T = -T \left(\frac{\partial^3 G}{\partial T^2 \partial p}\right) = -T \left(\frac{\partial^2}{\partial T^2} \left(\frac{\partial G}{\partial p}\right)_T\right)_p$$

$$= -T \left(\frac{\partial^2 V}{\partial T^2}\right)_p = -T \left(\frac{\partial}{\partial T} \left(\frac{\partial V}{\partial T}\right)_p\right)_p$$

$$= -T \frac{\partial}{\partial T} (V\alpha)_p \quad - \text{using definition of } \alpha$$

$$= -TV \left(\frac{\partial \alpha}{\partial T}\right)_p - T\alpha \left(\frac{\partial V}{\partial T}\right)_p$$

$$= -TV \left(\frac{\partial \alpha}{\partial T}\right)_p - T\alpha V\alpha$$

$$\left(\frac{\partial C_p}{\partial p}\right)_T = -TV \left[\alpha^2 + \left(\frac{\partial \alpha}{\partial T}\right)_p\right]$$

2) $T = \left(\frac{V}{V_0}\right)^\gamma T_0$ gives temperature as system changes volume V

a) work done on the gas

$$dW = -p dV$$

use ideal gas $pV = Nk_B T \Rightarrow p = \frac{Nk_B T}{V}$

$$dW = -Nk_B T \frac{dV}{V}$$

total work done is

$$W = \int dW = -Nk_B \int_{V_0}^{V_1} \frac{dV}{V} T(V)$$

$$= Nk_B \int_{V_0}^{V_1} \frac{dV}{V} \left(\frac{V}{V_0}\right)^\gamma T_0 = -Nk_B T_0 \int_{V_0}^{V_1} dV \frac{V^{\gamma-1}}{V_0^\gamma}$$

$$= -\frac{Nk_B T_0}{V_0^\gamma} \frac{V_1^\gamma - V_0^\gamma}{\gamma} = \frac{Nk_B T_0}{\gamma} \left[1 - \left(\frac{V_1}{V_0}\right)^\gamma\right]$$

$$W = \frac{Nk_B T_0}{\gamma} \left[1 - \left(\frac{V_1}{V_0}\right)^\gamma\right]$$

b) use $E = \frac{3}{2} Nk_B T$ for ideal gas

$$\Delta E = \frac{3}{2} Nk_B (T_1 - T_0) = \frac{3}{2} Nk_B T_0 \left[\left(\frac{V_1}{V_0}\right)^\gamma - 1\right]$$

c) Heat transfer to the gas

$$dQ = TdS = dE - dW \quad (\text{from } dE = TdS - p dV)$$
$$\Rightarrow \Delta Q = \Delta E - \Delta W \quad \text{and } dW = -p dV$$

$$\Delta Q = Nk_B T_0 \left[\left(\frac{V_1}{V_0} \right)^\eta - 1 \right] \frac{3}{2} - \frac{Nk_B T_0}{\eta} \left[1 - \left(\frac{V_1}{V_0} \right)^\eta \right]$$

$$\Delta Q = Nk_B T_0 \left[\left(\frac{V_1}{V_0} \right)^\eta - 1 \right] \left[\frac{3}{2} + \frac{1}{\eta} \right]$$

$$\Delta Q = 0 \quad \text{when} \quad \frac{3}{2} + \frac{1}{\eta} = 0 \quad \Rightarrow \quad \boxed{\eta = -\frac{2}{3}}$$

From problem ① if the gas expands adiabatically
(i.e. $\Delta S = 0 \Rightarrow \Delta Q = 0$) then

$$p V^{5/3} = \text{const}$$

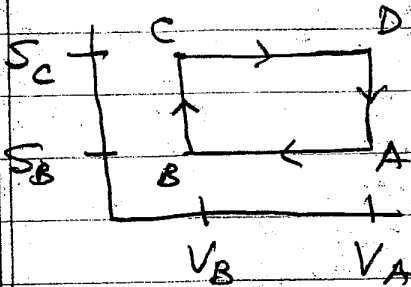
using $pV = Nk_B T$ the above becomes

$$\frac{Nk_B T}{V} V^{5/3} = Nk_B T V^{2/3} = \text{const}$$

$$\text{or } T \sim V^{-2/3}$$

which gives $\eta = -2/3$ in
agreement with part (c)

3) Otto cycle



heat extracted in $B \rightarrow C$
 work done by system in $C \rightarrow D$
 heat returned in $D \rightarrow A$
 work done on system in $A \rightarrow B$

$$\text{efficiency } \epsilon \equiv \frac{Q_{BC} - Q_{AD}}{Q_{BC}} = 1 - \frac{Q_{AD}}{Q_{BC}}$$

using $dQ = TdS$, the heat extracted from reservoir and added to the system in step $B \rightarrow C$

$$Q_{BC} = \int_B^C T dS \quad \text{integral is taken at constant } V$$

$$\text{Now we know } T = \left(\frac{\partial E}{\partial S} \right)_{V, N} \quad \text{so}$$

$$Q_{BC} = \int_B^C \left(\frac{\partial E}{\partial S} \right)_{V, N} dS = E_C - E_B$$

(we can do the integral since trajectory of integration is at constant V , or more generally,
 $TdS = dE + pdV$, but here $dV = 0$)

For ideal gas, as assumed here, $E = \frac{3}{2} N k_B T$

$$\text{So } Q_{BC} = \frac{3}{2} N k_B (T_C - T_B)$$

similarly

$$Q_{AD} = \frac{3}{2} N k_B (T_D - T_A)$$

efficiency

$$\epsilon = 1 - \frac{(T_D - T_A)}{(T_C - T_B)}$$

Finally we need to relate T_A, T_B, T_C, T_D to the volumes V_A and V_B

From Set 1, problem 4 we had for the adiabatic expansion of the ideal gas,

$$p V^{5/3} = \text{const}$$

$$\text{using } p V = N k_B T \text{ gives } N k_B T V^{2/3} = \text{const}$$

So along path $C \rightarrow D$ we get

$$T_C V_B^{2/3} = T_D V_A^{2/3} \Rightarrow T_D = T_C \left(\frac{V_B}{V_A} \right)^{2/3}$$

and along path $B \rightarrow A$ we get

$$T_B V_B^{2/3} = T_A V_A^{2/3} \Rightarrow T_A = T_B \left(\frac{V_B}{V_A} \right)^{2/3}$$

So

$$\epsilon = 1 - \frac{T_C \left(\frac{V_B}{V_A} \right)^{2/3} - T_B \left(\frac{V_B}{V_A} \right)^{2/3}}{T_C - T_B}$$

$$\epsilon = 1 - \left(\frac{V_B}{V_A} \right)^{2/3}$$

Now for ideal gas

$$C_V = \frac{3}{2} N k_B$$

$$C_P = \frac{5}{2} N k_B$$

$$\frac{C_P - C_V}{C_V} = \frac{5-3}{3} = \frac{2}{3}$$

$$\epsilon = 1 - \frac{N_B}{V_A} \quad \text{---} \quad (C_P - C_V) / C_V$$

4) N particles in a box of volume V travel with extremely relativistic speeds such that we have $|\vec{v}| \approx c$ and their energy can be taken to be $\epsilon = pc$, with p the magnitude of the relativistic momentum.

Consider such a particle colliding with the wall. We assume the collision is elastic.



Then, just like in notes 2-1 for non-relativistically moving particles, we can write the pressure on the wall as

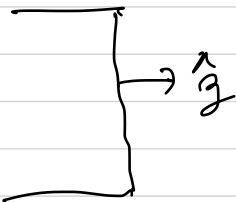
$$P = \left\langle \frac{(\Delta p_{\perp})_{\text{rate}}}{A} \right\rangle$$

P is pressure
 p is momentum.

where $\Delta p_{\perp} = 2p_{\perp}$ is the change in momentum of the particle during the collision, p_{\perp} is the component of momentum in the direction perpendicular to the wall, A is the wall area, and the rate of collisions is

$$\text{rate} = \frac{1}{2} \frac{N}{V} A v_{\perp}$$

with v_{\perp} the component of the velocity in the direction perpendicular to the wall.



If the normal vector to the wall is \hat{n} , then $p_{\perp} = p_z$ and $v_{\perp} = v_z$

$\langle \dots \rangle$ represents the average over all particles and all collisions

$$\Rightarrow P = \frac{N}{V} \langle P_{\perp} v_{\perp} \rangle$$

Now, by our assumption that particles travel equally likely in all directions we can write

$$\begin{aligned} \langle P_{\perp} v_{\perp} \rangle &= \langle P_x v_x \rangle = \langle P_y v_y \rangle = \langle P_z v_z \rangle \\ &= \frac{1}{3} \langle \vec{p} \cdot \vec{v} \rangle = \frac{1}{3} \langle p v \rangle \end{aligned}$$

The last step, $\vec{p} \cdot \vec{v} = p v$, is because \vec{p} and \vec{v} are colinear. Now we use $v \approx c$ and we get

$$P = \frac{1}{3} \frac{N}{V} \langle p c \rangle = \frac{1}{3} \frac{N}{V} \langle E \rangle$$

$$\Rightarrow \langle E \rangle = \frac{3 P V}{N}$$

now use the ideal gas law $P V = N k_B T$
to get

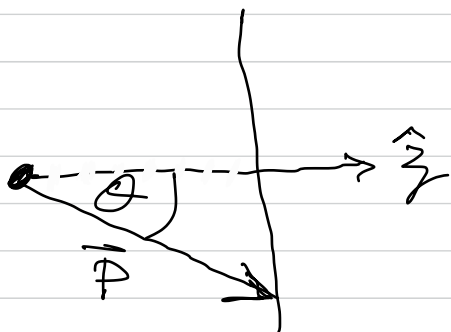
$$\boxed{\langle E \rangle = 3 k_B T}$$

compare that to the non-relativistic result $\langle E \rangle = \frac{3}{2} k_B T$

To further convince you that

$$\langle p_{\perp} v_{\perp} \rangle = \frac{1}{3} \langle p v \rangle$$

we can write



$$p_{\perp} = p \cos \theta$$

$$v_{\perp} = v \cos \theta$$

$$\langle p_{\perp} v_{\perp} \rangle = \langle p v \cos^2 \theta \rangle$$

The average $\langle \dots \rangle$ is an average over all particle speeds v , and all velocity directions, specified by the spherical angles (φ, θ) .

We can first do the average over all directions to get

$$\langle \cos^2 \theta \rangle = \frac{1}{4\pi} \int_0^{2\pi} d\varphi \int_0^{\pi} \sin \theta \cos^2 \theta$$

average over all spherical angles

$$\langle \cos^2 \theta \rangle = \frac{1}{2} \left[\frac{-\cos^3 \theta}{3} \right]_0^{\pi} = \frac{1}{2} \left[\frac{1}{3} + \frac{1}{3} \right] = \frac{1}{3}$$

$$\text{So } \langle p_{\perp} v_{\perp} \rangle = \langle p v \cos^2 \theta \rangle = \frac{1}{3} \langle p v \rangle$$

Note: Nowhere in our calculation did we need to use the relativistic expression for momentum

$$p = m \gamma v \quad \text{with} \quad \gamma = \frac{1}{\sqrt{1 - (v/c)^2}}$$