

1) photon occupation number

$$\langle n \rangle = \frac{\sum_n e^{-\beta \hbar \omega n} n}{\sum_n e^{-\beta \hbar \omega n}} = \frac{\sum_n e^{-\beta \hbar \omega n} n}{\omega}$$

where $\omega = \sum_n e^{-\beta \hbar \omega n}$

$$a) \frac{\partial \langle n \rangle}{\partial \beta} = \frac{1}{\omega} \sum_n e^{-\beta \hbar \omega n} n (-\hbar \omega n)$$

$$- \frac{1}{\omega^2} \left(\sum_n e^{-\beta \hbar \omega n} n \right) \frac{\partial \omega}{\partial \beta}$$

$$= -\hbar \omega \frac{\sum_n e^{-\beta \hbar \omega n} n^2}{\omega}$$

$$- \frac{1}{\omega^2} \left(\sum_n e^{-\beta \hbar \omega n} n \right) \left(\sum_n e^{-\beta \hbar \omega n} (-\hbar \omega n) \right)$$

$$= -\hbar \omega \langle n^2 \rangle + \hbar \omega \langle n \rangle^2$$

So $\boxed{-\frac{1}{\hbar \omega} \frac{\partial \langle n \rangle}{\partial \beta} = \langle n^2 \rangle - \langle n \rangle^2}$

b) For photons, $\langle n \rangle = \frac{1}{e^{\beta \hbar \omega} - 1}$

so $\frac{\partial \langle n \rangle}{\partial \beta} = \frac{-\hbar \omega e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2}$

so $\langle n^2 \rangle - \langle n \rangle^2 = \frac{-1}{\hbar \omega} \frac{\partial \langle n \rangle}{\partial \beta} = \frac{e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2}$

and

$$\frac{\langle n^2 \rangle - \langle n \rangle^2}{\langle n \rangle^2} = \frac{e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2} \cdot (e^{\beta \hbar \omega} - 1)^2$$

$$= e^{\beta \hbar \omega}$$

Relative fluctuation is

$$\frac{\sqrt{\langle n^2 \rangle - \langle n \rangle^2}}{\langle n \rangle} = e^{\beta \hbar \omega / 2} \gg 1$$

so fluctuations are always large

The bigger is $\beta \hbar \omega$, i.e. the smaller is $\frac{k_B T}{\hbar \omega}$,
the larger the fluctuations are.

2) $H = \frac{p^2}{2m} + \frac{1}{2} m \omega_0^2 x^2$ x is displacement from equilibrium position

a) Classically, use the equipartition theorem

$\langle \frac{1}{2} m \omega_0^2 x^2 \rangle = \frac{1}{2} k_B T$ as x is a quadratic degree of freedom, it contributes $\frac{1}{2} k_B T$ to the average energy

$$\Rightarrow \langle x^2 \rangle = \frac{k_B T}{m \omega_0^2}$$

by symmetry $\langle x \rangle = 0$
 so $\langle (x)^2 \rangle = \langle x^2 \rangle$

b) Quantum mechanically: use the quantum virial theorem

$$\langle \frac{p^2}{2m} \rangle = \langle \frac{1}{2} m \omega_0^2 x^2 \rangle = \frac{1}{2} \langle H \rangle$$

$$\langle x^2 \rangle = \frac{\langle H \rangle}{m \omega_0^2} \quad \text{again } \langle x \rangle = 0$$

To compute $\langle H \rangle$, we know that $\langle H \rangle = \hbar \omega_0 \left[\langle n \rangle + \frac{1}{2} \right]$

where

$$\langle n \rangle = \frac{1}{e^{\beta \hbar \omega_0} - 1}$$

is the average excitation number of the oscillator

$$\text{so } \langle x^2 \rangle = \frac{\hbar \omega_0}{m \omega_0^2} \left[\frac{1}{e^{\beta \hbar \omega_0} - 1} + \frac{1}{2} \right]$$

$$= \frac{\hbar}{2m \omega_0} \left[\frac{e^{\beta \hbar \omega_0} + 1}{e^{\beta \hbar \omega_0} - 1} \right]$$

c) We expect the quantum result to reduce to the classical result when $k_B T \gg \hbar \omega_0$ i.e. when thermal energy is much greater than the quantum energy level spacing

So when $k_B T \gg \hbar \omega_0 \Rightarrow \beta \hbar \omega_0 \ll 1$

$$\langle x^2 \rangle \approx \frac{\hbar}{m \omega_0} \left[\frac{1}{1 + \beta \hbar \omega_0} + \frac{1}{2} \right] \quad \begin{array}{l} \text{as } e^x \sim 1 + x + \dots \\ \text{as } x \rightarrow 0 \end{array}$$

$$\approx \frac{\hbar}{m \omega_0} \left[\frac{k_B T}{\hbar \omega_0} + \frac{1}{2} \right]$$

↑ ignore since $\frac{1}{\beta \hbar \omega_0} \gg 1$

$$\approx \frac{k_B T}{m \omega_0^2} \quad \text{same as classical result}$$

We could also have done the quantum calculation a more straight forward way. The position operator for the quantum oscillator is

$$\hat{x} = \sqrt{\frac{\hbar}{2m\omega_0}} (a^\dagger + a)$$

So

$$\hat{x}^2 = \frac{\hbar}{2m\omega_0} (a^\dagger + a)(a^\dagger + a)$$

$$= \frac{\hbar}{2m\omega_0} (a^\dagger a^\dagger + a a^\dagger + a^\dagger a + a a)$$

The thermal average is then

$$\langle \hat{x}^2 \rangle = \text{Tr} [\hat{\rho} \hat{x}^2]$$

which is most easily evaluated in the energy eigenstate basis

$$\langle \hat{x}^2 \rangle = \sum_n p_n \langle n | \hat{x}^2 | n \rangle$$

where

$$p_n = \frac{e^{-\beta E_n}}{\sum_n e^{-\beta E_n}}$$

and $E_n = \hbar\omega (n + 1/2)$ are the energy eigenvalues

We have

$$\langle n | \hat{X}^2 | n \rangle = \frac{\hbar}{2m\omega_0} \langle n | a^\dagger a^\dagger + a a^\dagger + a^\dagger a + a a | n \rangle$$

$$\text{now } \langle n | a^\dagger a^\dagger | n \rangle = \langle n | a a | n \rangle = 0$$

$$\text{and } [a, a^\dagger] = 1 \Rightarrow a a^\dagger = a^\dagger a + 1$$

$$\text{So } \langle n | \hat{X}^2 | n \rangle = \frac{\hbar}{2m\omega_0} \langle n | 2a^\dagger a + 1 | n \rangle$$

Now $a^\dagger a = \hat{n}$ is just the number operator, so

$$\langle n | \hat{X}^2 | n \rangle = \frac{\hbar}{2m\omega_0} (2n + 1)$$

$$= \frac{\hbar}{m\omega_0} \left(n + \frac{1}{2} \right) = \frac{E_n}{m\omega_0^2}$$

$$\text{So } \langle \hat{X}^2 \rangle = \sum_n \rho_n \frac{E_n}{m\omega_0^2} = \frac{1}{m\omega_0^2} \langle E \rangle$$

$$\text{where } \langle E \rangle = \hbar\omega_0 \left(\langle n \rangle + \frac{1}{2} \right)$$

$$\text{using } \langle n \rangle = \frac{1}{e^{\beta\hbar\omega_0} - 1}$$

we recover our previous result.