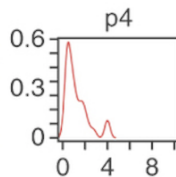
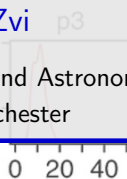
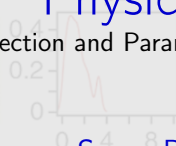


# Physics 403

Model Selection and Parameter Estimation, Part I

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## 1 Review of Last Class

## 2 Parameter Estimation: Bayesian Framework

- Reminder of the Basics
- Effect of the Prior
- Marginalization
- Comparing Models (Odds Ratio)
- Statistical Trials
- The Occam Factor

## 3 Systematic Uncertainties

# This Week in Bad Plots



# This Week in Bad Plots

Notice Anything Wrong?



# Last Time

## Generation of pseudo-random numbers for simulation

- ▶ Simulation, data challenges, parameter estimation
- ▶ Linear Congruential Generators
- ▶ Mersenne Twister and Xorshift Generators
- ▶ Word of caution about **seeding** your RNG: system clock, /dev/random, etc.

## Generating random numbers from arbitrary PDFs

- ▶ Inversion method, if PDF integrable and CDF invertible
- ▶ Acceptance/rejection method, works for most cases
- ▶ Gaussian and Poisson random numbers

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## 3 Systematic Uncertainties

## Reminder of the Basics

Recall the basic rules of probability introduced at the start of the course:

► **Sum Rule:**

$$p(H|I) + p(\bar{H}|I) = 1, \quad \sum_i p(H_i|I) = 1$$

► **Product Rule:**

$$p(H_i, D|I) = p(D|H_i, I)p(H_i|I) = p(H_i|D, I)p(D|I)$$

► **Bayes' Theorem:**

$$p(H_i|D, I) = \frac{p(D|H_i, I)p(H_i|I)}{p(D|I)}$$

► **Law of total probability:**

$$\sum_i p(H_i|D, I) = \frac{\sum_i p(D|H_i, I)p(H_i|I)}{p(D|I)} = 1$$
$$\therefore p(D|I) = \sum_i p(D|H_i, I)p(H_i|I)$$

## Reminder of the Basics

The law of total probability has a continuous counterpart. For example, given a model  $M$  with parameters  $\theta$ ,

$$p(D|M) = \int_V d\theta \, p(D|\theta, M) \, p(\theta|M)$$

Interpretation: the likelihood of model  $M$  is the weighted average likelihood for its parameters  $\theta$ .

**Parameter Estimation:** the determination of the values of model parameters  $\theta$  using data.

- ▶ Bayesian: evaluate the full posterior PDF  $p(\theta|D, M)$  or “best fit” summary values like the mean or mode. Uses **prior**  $p(\theta|M)$
- ▶ Frequentist: evaluate the best fit values from the likelihood alone
- ▶ Both approaches: give some allowed range of parameter with some probability measure (confidence interval, or credible range)



# Effect of the Prior

- ▶ The presence of a prior tends to make many people upset, because you can get different answers depending on the prior you choose.
- ▶ Bayesian answer: that's exactly right, but **so what?**
- ▶ The prior is how we **incorporate external information** about the quantities being tested
- ▶ If the posterior PDF is dominated by the prior, that just means the data are not constraining our model parameters
- ▶ **Note:** frequentists don't use priors, which in practice means that **assumptions are hidden**
- ▶ Best practice: **report likelihoods and priors separately**, and show the effect of different priors on the posterior

# Coin Flipping

## Example

From Sivia, Ch. 2 [1]: we walk into a casino and start betting on the outcomes of flipping a coin. (It's not a very impressive casino.)

- ▶ We don't know the probability  $h$  of getting heads, so we have to choose some  $p(h|I)$ .
- ▶ We do know that given  $h$ , the probability of observing heads  $r$  times in  $N$  coin flips is given by the binomial PDF

$$p(r|N, h, I) \propto h^r (1 - h)^{N-r}$$

What is the effect of  $p(h|I)$  on the posterior probability  $p(h|N, r, I)$ , the distribution of  $h$  given  $r$  heads in  $N$  tosses? From Bayes' Theorem,

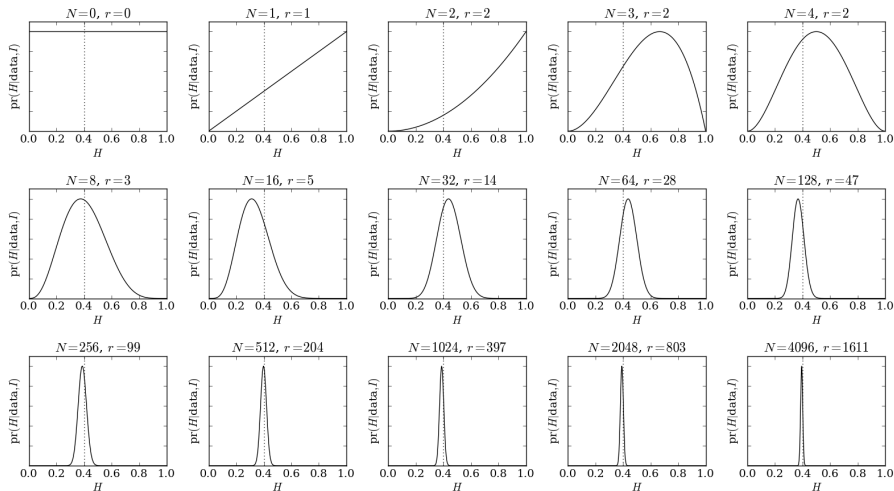
$$p(h|N, r, I) \propto p(r|h, N, I) p(h|I),$$

so let's try out different priors and see what happens.

# Coin Flipping

## Uniform “Ignorance” Prior

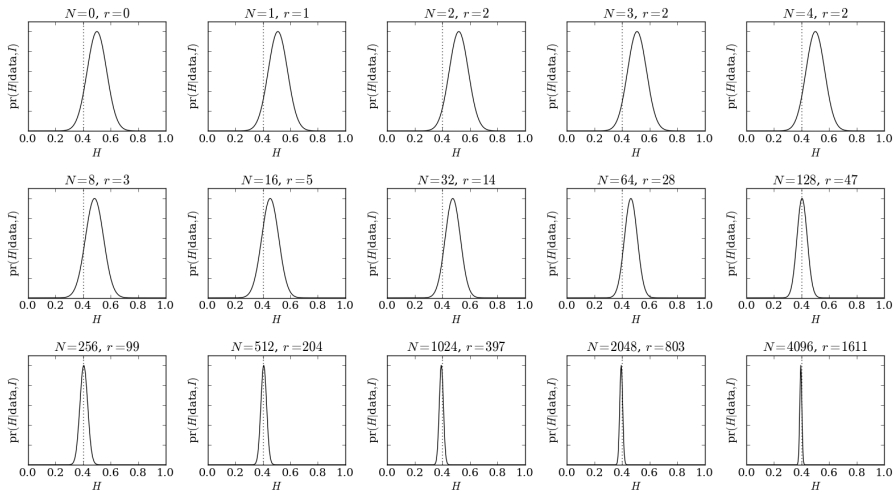
We start with **no preferred value for  $h$** :



# Coin Flipping

## Fair Coin Prior

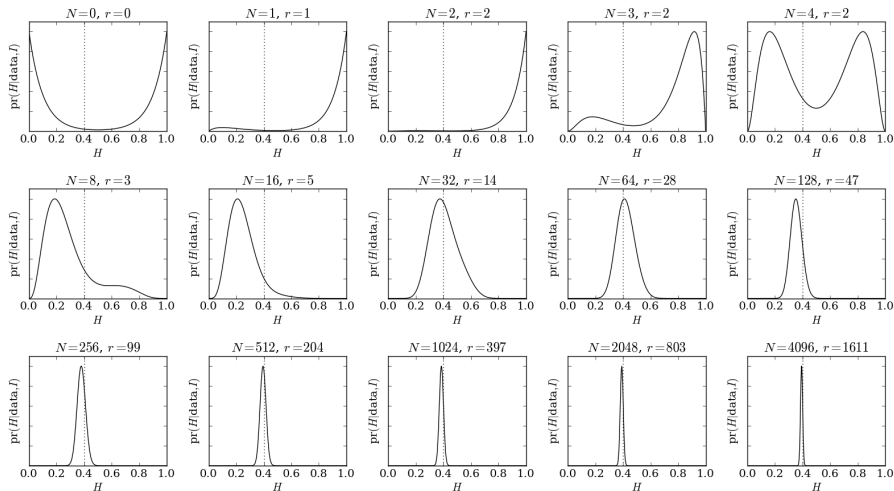
We assume the **coin is fair** ( $h = 0.5$ ) with some uncertainty:



# Coin Flipping

## Unfair Coin Prior

We assume the coin is **very unfair**, but don't know the bias.



# Marginalization

- ▶ Recall the definition of marginalization and marginal distributions: if we don't care about the effect of some parameter on a probability, we can **integrate it out**
- ▶ Example: for model  $M$  with parameters  $\theta, \varphi$ , if we are only interested in  $\theta$  then we can calculate the marginal PDF

$$p(\theta|D, M) = \int d\varphi \, p(\theta, \varphi|D, M)$$

- ▶ Marginalization is a general technique in Bayesian analysis that doesn't have an analog in frequentist statistics
- ▶ Terms that we don't care about are called **nuisance parameters** in frequentist statistics. There is no general procedure for handling them

# Model Comparison

- ▶ One topic we haven't discussed yet is **model comparison**
- ▶ The idea: compare two competing models by calculating the probability of each model given the data  $D$
- ▶ If we want to compare two or more alternative models  $M_i$ , then use Bayes' Theorem to calculate the posterior probability of each model:

$$p(M_i|D, I) = \frac{p(D|M_i, I)p(M_i|I)}{p(D|I)}$$

- ▶ This is analogous to parameter estimation, except instead of estimating  $p(\theta|D, I)$  for a parameter, we estimate  $p(M_i|D, I)$  for a model
- ▶ The math is the same, but the interpretation differs

# The Odds Ratio

To select between two models, it is useful to calculate the ratio of the posterior probabilities of the models. This is called the **odds ratio**:

$$\begin{aligned} O_{ij} &= \frac{p(D|M_i, I)}{p(D|M_j, I)} \frac{p(M_i|I)}{p(M_j|I)} \\ &= B_{ij} \frac{p(M_i|I)}{p(M_j|I)} \end{aligned}$$

The first term is called the **Bayes Factor** [2, 3] and the second is called the **prior odds ratio**. Interpretation:

- ▶ **Prior odds**: the amount by which you favor  $M_i$  over  $M_j$  *before taking data*. There is no analog in frequentist statistics.
- ▶ **Bayes Factor**: the amount that the data  $D$  causes you favor  $M_i$  over  $M_j$ . Frequentist analog: *likelihood ratio* (but frequentists can't marginalize nuisance parameters)



# The Odds Ratio

Interpreting the Bayes Factor, according to Jeffreys [2]:

$B_{ij}$	Strength of Evidence
$< 1 : 1$	negative (supports $M_j$ )
$1 : 1$ to $3 : 1$	barely worth mentioning
$3 : 1$ to $10 : 1$	substantial support for $M_i$
$10 : 1$ to $30 : 1$	strong support for $M_i$
$30 : 1$ to $100 : 1$	very strong support for $M_i$
$> 100 : 1$	decisive evidence for $M_i$

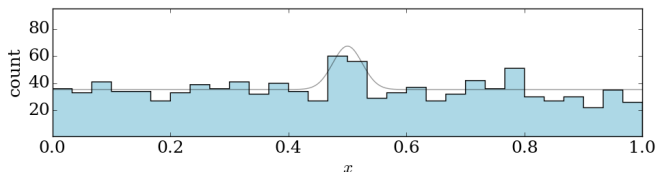
But wait, remember the “**5 $\sigma$  rule?**” That corresponds to a Gaussian *tail probability* (or **p-value**) of  $6 \times 10^{-7}$ . Isn't that MUCH stronger evidence than  $100 : 1$  odds. What's going on?

Partial answer: odds ratios and  $p$ -values are not the same thing. Not to mention the “look elsewhere effect” and other sources of **statistical trials**

# Aside: Statistical Trials

## The Look Elsewhere Effect

- ▶ Suppose you are looking for a spike in some data with background, e.g., a mass resonance or a spectral line, but you don't know the location of the feature, just a range of interest
- ▶ You **scan over the data** and find a spike which is  $> 3\sigma$  above the background ( $p$ -value:  $\sim 0.1\%$ ). Is this significant?



- ▶ Hang on: because location was a free parameter, you need to account for the fact that **any one of the bins you looked at could have been an upward fluctuation of the background**
- ▶ Conservatively,  $p \rightarrow N_{\text{bins}} \times p \approx 2\%$ , or  $\sim 2\sigma$

# Statistical Trials in the Bayesian Framework

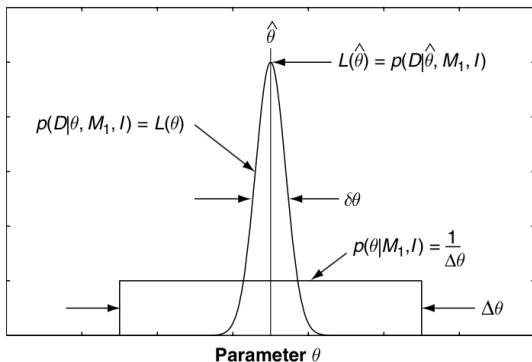
## Occam's Razor

- ▶ **Occam's Razor**: when selecting from among competing models, generally prefer the simpler model
- ▶ For model comparison, the Bayes Factor and odds ratio have a built-in Occam's razor
- ▶ Searching for a spike: in the Bayesian framework, we would treat the location of the spike as a **nuisance parameter** and marginalize it (model  $M_1$ )
- ▶ Compare this to a model with no spike ( $M_0$ )
- ▶ If we did everything correctly,  $p(D|M_1)$  should have extra terms compared to  $p(D|M_0)$  which “penalize” it for our ignorance of the location of the spike
- ▶ So **a piece of the odds ratio** should account for statistical trials and favor the simpler model!

# Occam's Razor

## The Bayesian Framework

Let's be more explicit. Imagine  $M_1$  has a single parameter  $\theta$  (e.g., the location of a spike) which is unknown.  $M_0$  has  $\theta$  fixed at  $\theta_0$ .



Suppose our prior on  $\theta$  is **uniform** in model  $M_1$ . I.e., we don't know what it is, just that it lies in some range  $\Delta\theta$ . And suppose the data tell us a lot about  $\theta$ , so  **$p(D|\theta, M_1, I)$  is very peaked** about  $\hat{\theta}$  with width  $\delta\theta$ .

# Occam's Razor

## The Bayesian Framework

The “**global likelihood**” of the data given  $M_1$  (independent of  $\theta$ ) is

$$\begin{aligned} p(D|M_1, I) &= \int d\theta \, p(D|\theta, M_1, I) \, p(\theta|M_1, I) \\ &= \int d\theta \, p(D|\theta, M_1, I) \, \frac{1}{\Delta\theta} \\ &\approx p(D|\hat{\theta}, M_1, I) \, \frac{\delta\theta}{\Delta\theta} \end{aligned}$$

Since  $M_0$  has no free parameters, its **global likelihood** is

$$\begin{aligned} p(D|M_0, I) &= \int d\theta \, p(D|\theta, M_1, I) \, \delta(\theta - \theta_0) \\ &= p(D|\theta_0, M_1, I) \end{aligned}$$

I.e., it's just the likelihood of model  $M_1$  with  $\theta$  **fixed**.

# Occam's Razor

## The Bayesian Framework

Putting it all together, the Bayes factor in favor of the more complex model  $M_1$  is

$$\begin{aligned} B_{10} &\approx \frac{p(D|\hat{\theta}, M_1, I)}{p(D|\theta_0, M_1, I)} \frac{\delta\theta}{\Delta\theta} \\ &= \frac{\mathcal{L}(\hat{\theta})}{\mathcal{L}(\theta_0)} \frac{\delta\theta}{\Delta\theta} \end{aligned}$$

The first term is a **likelihood ratio**, which **favors  $M_1$**  because of the strong peak at  $\hat{\theta}$ .

But the second term **penalizes  $M_1$**  since  $\delta\theta < \Delta\theta$ . In other words,  $M_1$  is penalized because of the wasted parameter space that gets ruled out by the data.

# The Occam Factor

- ▶ Generalizing from this specific problem, we can express any likelihood of data  $D$  given a model  $M$  as the maximum value of its likelihood times an **Occam factor**:

$$p(D|M, I) = \mathcal{L}_{\max} \Omega_{\theta}$$

- ▶ The Occam factor corrects the likelihood for the **statistical trials** incurred by scanning the parameter space for  $\hat{\theta}$ .
- ▶ The odds ratio automatically accounts for these factors. It is in this way that the Bayesian framework prevents overfitting of data with arbitrarily complicated models.
- ▶ **Note:** in frequentist statistics, statistical penalties are more of a kluge. There are many ways to calculate them (e.g., the  $N_{\text{bins}}$  factor used earlier) but no simple framework.

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## 3 Systematic Uncertainties



# Systematic Uncertainties

Recall that there are two types of experimental uncertainties:

1. **Random:** uncertainties which can be reduced by acquiring and averaging more data (details on this next class)
2. **Systematic:** uncertainties which are fixed and tend to affect all measurements equally

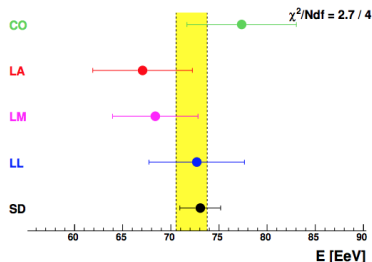
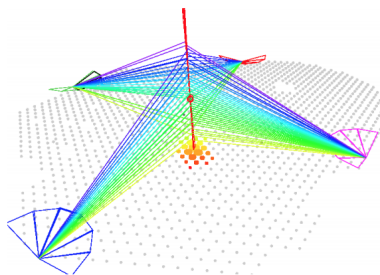
## Example

Calibrations of meters and rulers are a classic example of systematic uncertainties.

- ▶ Wooden meter sticks may shrink by several mm over time
- ▶ Energy scales in detectors may be uncertain due to other experimental or theoretical uncertainties
- ▶ Astronomical “rulers” have lots of systematic uncertainties, e.g., Hubble’s constant  $H_0$

# Systematic Uncertainties

We try to tabulate systematic uncertainties in an “error budget”:



Systematic uncertainties in the scale of cosmic-ray energy measurements at the Pierre Auger Observatory [4]:

Source	Uncertainty
Fluorescence Yield $Y$	14%
$\rho$ , $T$ , $e$ Effects on $Y$	7%
Calibration	9.5%
Atmosphere	4%
Reconstruction	10%
Invisible Energy	4%
Total	22%

# Reducing Systematic Uncertainties

- ▶ We can try to reduce systematic uncertainties by **changing our experimental procedure**
- ▶ Or, we work on **secondary measurements** to better evaluate physical quantities that affect our primary calculations
- ▶ In the case of Auger, the collaboration put a lot of effort into reducing systematic uncertainties related to the production of fluorescence light by  $N_2$ :
  1. Measurement of the absolute level of fluorescence production in the lab: FLASH (SLAC) [5] and AIRFLY (ANL) [6, 7]
  2. Characterization of the “quenching” of fluorescence by molecular collisions [8] and careful measurements of atmospheric conditions [9]
- ▶ Result: one of the largest sources of systematic uncertainties in the energy scale reduced to the few percent level

# Marginalization of Uncertainties

- ▶ What happens if you can't reduce systematic uncertainties to a negligible level?
- ▶ Bayesian approach: we almost always have some **prior information** about the accuracy of our “ruler.”
- ▶ Incorporate this prior information by **parameterizing the systematic uncertainty** and then **marginalizing the scale**

## Example

You want to compare the cosmic ray flux measured by several different experiments, but the experiments used different measurement techniques and have different systematic uncertainties. As a result, the spectra are offset from each other. What do we do?

Parameterize the systematics as Gaussians of known mean and width, and marginalize the absolute energy scale using these PDFs [10]

# Summary

- ▶ The formalism for parameter estimation and model selection in Bayesian statistics is mathematically the same
- ▶ We estimate parameters by looking at the **PDF** and its **maximum likelihood** (same as frequentist approach)
- ▶ We perform model selection by computing an **odds ratio** and making a decision about the odds. In frequentist approach: a likelihood ratio test, or Neyman-Pearson test
- ▶ The odds ratio has a built-in **Occam factor** that accounts for “scanning” for the best value in a parameter space
- ▶ Marginalization gives us a uniform way of handling unknown *nuisance parameters*, **including systematic uncertainties**

# References I

- [1] D.S. Sivia and John Skilling. *Data Analysis: A Bayesian Tutorial*. New York: Oxford University Press, 1998.
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- [3] Robert E. Kass and Adrian E. Raftery. “Bayes Factors”. In: *J. Am. Stat. Assoc.* 90.430 (1995), pp. 773–795. URL: <http://amstat.tandfonline.com/doi/abs/10.1080/01621459.1995.10476572>.
- [4] B.R. Dawson et al. “Hybrid Performance of the Pierre Auger Observatory”. In: *Proc. 30th ICRC*. Merida, Mexico, 2007. arXiv: 0706.1105 [astro-ph.HE].
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- [7] M. Ave et al. “Precise measurement of the absolute fluorescence yield of the 337 nm band in atmospheric gases”. In: *Astropart.Phys.* 42 (2013), pp. 90–102. arXiv: 1210.6734 [astro-ph.IM].
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- [10] Brian M. Connolly et al. “Comparison of the Ultrahigh Energy Cosmic Ray Flux Observed by AGASA, HiRes and Auger”. In: *Phys.Rev.* D74 (2006), p. 043001. arXiv: astro-ph/0606343 [astro-ph].